

$$I^G(J^{PC}) = 0^-(1^{--})$$

### $\phi(1020)$ MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1019.461 ± 0.016 OUR AVERAGE</b>				
1019.463 ± 0.061	2.3M	<sup>1</sup> KOZYREV 18	CMD3	$e^+e^- \rightarrow K^+K^-, K_S^0 K_L^0$
1019.462 ± 0.042 ± 0.056	28k	<sup>2</sup> LEES 14H	BABR	$e^+e^- \rightarrow K_S^0 K_L^0 \gamma$
1019.51 ± 0.02 ± 0.05		<sup>3</sup> LEES 13Q	BABR	$e^+e^- \rightarrow K^+K^-\gamma$
1019.30 ± 0.02 ± 0.10	105k	AKHMETSHIN 06	CMD2	$0.98-1.06 e^+e^- \rightarrow \pi^+\pi^-\pi^0$
1019.52 ± 0.05 ± 0.05	17.4k	AKHMETSHIN 05	CMD2	$0.60-1.38 e^+e^- \rightarrow \eta\gamma$
1019.483 ± 0.011 ± 0.025	272k	<sup>4</sup> AKHMETSHIN 04	CMD2	$e^+e^- \rightarrow K_L^0 K_S^0$
1019.42 ± 0.05	1900k	<sup>5</sup> ACHASOV 01E	SND	$e^+e^- \rightarrow K^+K^-, K_S K_L, \pi^+\pi^-\pi^0$
1019.40 ± 0.04 ± 0.05	23k	AKHMETSHIN 01B	CMD2	$e^+e^- \rightarrow \eta\gamma$
1019.36 ± 0.12		<sup>6</sup> ACHASOV 00B	SND	$e^+e^- \rightarrow \eta\gamma$
1019.38 ± 0.07 ± 0.08	2200	<sup>7</sup> AKHMETSHIN 99F	CMD2	$e^+e^- \rightarrow \pi^+\pi^-\geq 2\gamma$
1019.51 ± 0.07 ± 0.10	11169	AKHMETSHIN 98	CMD2	$e^+e^- \rightarrow \pi^+\pi^-\pi^0$
1019.5 ± 0.4		BARBERIS 98	OMEG	$450 pp \rightarrow pp2K^+2K^-$
1019.42 ± 0.06	55600	AKHMETSHIN 95	CMD2	$e^+e^- \rightarrow$ hadrons
1019.7 ± 0.3	2012	DAVENPORT 86	MPSF	$400 pA \rightarrow 4KX$
1019.7 ± 0.1 ± 0.1	5079	ALBRECHT 85D	ARG	$10 e^+e^- \rightarrow K^+K^-X$
1019.3 ± 0.1	1500	ARENTON 82	AEMS	$11.8$ polar. $pp \rightarrow KK$
1019.67 ± 0.17	25080	<sup>8</sup> PELLINEN 82	RVUE	
1019.52 ± 0.13	3681	BUKIN 78C	OLYA	$e^+e^- \rightarrow$ hadrons
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
1018.4 ± 0.5 ± 0.1		<sup>9</sup> ALBRECHT 20	CBAR	$0.9 \bar{p}p \rightarrow K^+K^-\pi^0$
1019.21 ± 0.04 ± 0.03		<sup>10</sup> HOID 20	RVUE	$e^+e^- \rightarrow \pi^0\gamma$
1019.54 ± 0.10 ± 0.51		<sup>11</sup> AAIJ 19H	LHCB	$pp \rightarrow D^\pm X$
1019.20 ± 0.02 ± 0.01		<sup>12</sup> HOFERICH... 19	RVUE	$e^+e^- \rightarrow \pi^+\pi^-\pi^0$
1019.469 ± 0.061	1.7M	KOZYREV 18	CMD3	$e^+e^- \rightarrow K^+K^-$
1019.457 ± 0.061	610k	KOZYREV 16	CMD3	$e^+e^- \rightarrow K_S^0 K_L^0$
1019.48 ± 0.01		LEES 13F	BABR	$D^+ \rightarrow K^+K^-\pi^+$
1019.441 ± 0.008 ± 0.080	542k	<sup>13</sup> AKHMETSHIN 08	CMD2	$1.02 e^+e^- \rightarrow K^+K^-$
1019.63 ± 0.07	12540	<sup>14</sup> AUBERT,B 05J	BABR	$D^0 \rightarrow \bar{K}^0 K^+K^-$
1019.8 ± 0.7		ARMSTRONG 86	OMEG	$85 \pi^+ / pp \rightarrow \pi^+ / p4Kp$
1020.1 ± 0.11	5526	<sup>14</sup> ATKINSON 86	OMEG	$20-70 \gamma p$
1019.7 ± 1.0		BEBEK 86	CLEO	$e^+e^- \rightarrow \Upsilon(4S)$
1019.411 ± 0.008	642k	<sup>15</sup> DIJKSTRA 86	SPEC	$100-200 \pi^\pm, \bar{p}, p, K^\pm, \text{on Be}$

1020.9	$\pm 0.2$		<sup>14</sup>	FRAME	86	OMEG	13	$K^+ p \rightarrow \phi K^+ p$
1021.0	$\pm 0.2$		<sup>14</sup>	ARMSTRONG	83B	OMEG	18.5	$K^- p \rightarrow K^- K^+ \Lambda$
1020.0	$\pm 0.5$		<sup>14</sup>	ARMSTRONG	83B	OMEG	18.5	$K^- p \rightarrow K^- K^+ \Lambda$
1019.7	$\pm 0.3$		<sup>14</sup>	BARATE	83	GOLI	190	$\pi^- \text{Be} \rightarrow 2\mu X$
1019.8	$\pm 0.2$	$\pm 0.5$	766	IVANOV	81	OLYA	1-1.4	$e^+ e^- \rightarrow K^+ K^-$
1019.4	$\pm 0.5$	337		COOPER	78B	HBC	0.7-0.8	$\bar{p} p \rightarrow K_S^0 K_L^0 \pi^+ \pi^-$
1020	$\pm 1$	383	<sup>14</sup>	BALDI	77	CNTR	10	$\pi^- p \rightarrow \pi^- \phi p$
1018.9	$\pm 0.6$	800		COHEN	77	ASPK	6	$\pi^\pm N \rightarrow K^+ K^- N$
1019.7	$\pm 0.5$	454		KALBFLEISCH	76	HBC	2.18	$K^- p \rightarrow \Lambda K \bar{K}$
1019.4	$\pm 0.8$	984		BESCH	74	CNTR	2	$\gamma p \rightarrow p K^+ K^-$
1020.3	$\pm 0.4$	100		BALLAM	73	HBC	2.8-9.3	$\gamma p$
1019.4	$\pm 0.7$			BINNIE	73B	CNTR		$\pi^- p \rightarrow \phi n$
1019.6	$\pm 0.5$	120	<sup>16</sup>	AGUILAR-...	72B	HBC	3.9,4.6	$K^- p \rightarrow \Lambda K^+ K^-$
1019.9	$\pm 0.5$	100	<sup>16</sup>	AGUILAR-...	72B	HBC	3.9,4.6	$K^- p \rightarrow K^- p K^+ K^-$
1020.4	$\pm 0.5$	131		COLLEY	72	HBC	10	$K^+ p \rightarrow K^+ p \phi$
1019.9	$\pm 0.3$	410		STOTTLE...	71	HBC	2.9	$K^- p \rightarrow \Sigma / \Lambda K \bar{K}$

<sup>1</sup> Average of KOZYREV 16 and KOZYREV 18 values taking into account the correlated uncertainties. Supersedes individual KOZYREV 16 and KOZYREV 18 results.

<sup>2</sup> Using a vector meson dominance model with contribution from  $\phi(1020)$  and higher mass excitations of  $\rho(770)$ ,  $\omega(782)$ , and  $\phi(1020)$ .

<sup>3</sup> Using a phenomenological model based on KUHN 90 with a sum of Breit-Wigner resonances for  $\rho(770)$ ,  $\omega(782)$ ,  $\phi(1020)$  and their higher mass excitations.

<sup>4</sup> Update of AKHMETSHIN 99D

<sup>5</sup> From the combined fit assuming that the total  $\phi(1020)$  production cross section is saturated by those of  $K^+ K^-$ ,  $K_S^0 K_L^0$ ,  $\pi^+ \pi^- \pi^0$ , and  $\eta \gamma$  decays modes and using ACHASOV 00B for the  $\eta \gamma$  decay mode.

<sup>6</sup> Using a total width of  $4.43 \pm 0.05$  MeV. Systematic uncertainty included.

<sup>7</sup> Using a total width of  $4.43 \pm 0.05$  MeV.

<sup>8</sup> PELLINEN 82 review includes AKERLOF 77, DAUM 81, BALDI 77, AYRES 74, DE-GROOT 74.

<sup>9</sup> Width fixed at 4.2 MeV.

<sup>10</sup> The values were extracted from a dispersively improved Breit-Wigner parameterization and do not include vacuum polarization. Inclusion of vacuum polarization gives  $1019.457 \pm 0.020$  MeV.

<sup>11</sup> From the  $D^\pm \rightarrow K^\pm K^+ K^-$  Dalitz plot fit with the Triple-M amplitude in the multi-meson model of AOUDE 18.

<sup>12</sup> The values were extracted from a dispersively improved Breit-Wigner parameterization and do not include vacuum polarization.

<sup>13</sup> Strongly correlated with AKHMETSHIN 04.

<sup>14</sup> Systematic errors not evaluated.

<sup>15</sup> Weighted and scaled average of 12 measurements of DIJKSTRA 86.

<sup>16</sup> Mass errors enlarged by us to  $\Gamma/\sqrt{N}$ ; see the note with the  $K^*(892)$  mass.

### $\phi(1020)$ WIDTH

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>4.249±0.013 OUR AVERAGE</b>		Error includes scale factor of 1.1.		
4.245±0.013	2.3M	<sup>1</sup> KOZYREV	18	CMD3 $e^+e^- \rightarrow K^+K^-$ , $K_S^0 K_L^0$
4.205±0.103±0.067	28k	<sup>2</sup> LEES	14H	BABR $e^+e^- \rightarrow K_S^0 K_L^0 \gamma$
4.29 ±0.04 ±0.07		<sup>3</sup> LEES	13Q	BABR $e^+e^- \rightarrow K^+K^- \gamma$
4.30 ±0.06 ±0.17	105k	AKHMETSHIN	06	CMD2 0.98–1.06 $e^+e^- \rightarrow \pi^+\pi^-\pi^0$
4.280±0.033±0.025	272k	<sup>4</sup> AKHMETSHIN	04	CMD2 $e^+e^- \rightarrow K_L^0 K_S^0$
4.21 ±0.04	1900k	<sup>5</sup> ACHASOV	01E	SND $e^+e^- \rightarrow K^+K^-$ , $K_S K_L, \pi^+\pi^-\pi^0$
4.44 ±0.09	55600	AKHMETSHIN	95	CMD2 $e^+e^- \rightarrow$ hadrons
4.5 ±0.7	1500	ARENTON	82	AEMS 11.8 polar. $pp \rightarrow KK$
4.2 ±0.6	766	<sup>6</sup> IVANOV	81	OLYA 1–1.4 $e^+e^- \rightarrow K^+K^-$
4.3 ±0.6		<sup>6</sup> CORDIER	80	DM1 $e^+e^- \rightarrow \pi^+\pi^-\pi^0$
4.36 ±0.29	3681	<sup>6</sup> BUKIN	78c	OLYA $e^+e^- \rightarrow$ hadrons
4.4 ±0.6	984	<sup>6</sup> BESCH	74	CNTR $2\gamma p \rightarrow pK^+K^-$
4.67 ±0.72	681	<sup>6</sup> BALAKIN	71	OSPK $e^+e^- \rightarrow$ hadrons
4.09 ±0.29		BIZOT	70	OSPK $e^+e^- \rightarrow$ hadrons
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
4.07 ±0.13 ±0.01		<sup>7</sup> HOID	20	RVUE $e^+e^- \rightarrow \pi^0 \gamma$
4.23 ±0.04 ±0.02		<sup>8</sup> HOFERICHT...	19	RVUE $e^+e^- \rightarrow \pi^+\pi^-\pi^0$
4.249±0.015	1.7M	KOZYREV	18	CMD3 $e^+e^- \rightarrow K^+K^-$
4.240±0.017	610k	KOZYREV	16	CMD3 $e^+e^- \rightarrow K_S^0 K_L^0$
4.37 ±0.02		LEES	13F	BABR $D^+ \rightarrow K^+K^-\pi^+$
4.24 ±0.02 ±0.03	542k	<sup>9</sup> AKHMETSHIN	08	CMD2 1.02 $e^+e^- \rightarrow K^+K^-$
4.28 ±0.13	12540	<sup>10</sup> AUBERT,B	05J	BABR $D^0 \rightarrow \bar{K}^0 K^+K^-$
4.45 ±0.06	271k	DIJKSTRA	86	SPEC 100 $\pi^- \text{Be}$
3.6 ±0.8	337	<sup>6</sup> COOPER	78B	HBC 0.7–0.8 $\bar{p}p \rightarrow K_S^0 K_L^0 \pi^+\pi^-$
4.5 ±0.50	1300	<sup>6,10</sup> AKERLOF	77	SPEC 400 $pA \rightarrow K^+K^-X$
4.5 ±0.8	500	<sup>6,10</sup> AYRES	74	ASPK 3–6 $\pi^- p \rightarrow K^+K^-n, K^-p \rightarrow K^+K^-\Lambda/\Sigma^0$
3.81 ±0.37		COSME	74B	OSPK $e^+e^- \rightarrow K_L^0 K_S^0$
3.8 ±0.7	454	<sup>6</sup> BORENSTEIN	72	HBC 2.18 $K^- p \rightarrow K\bar{K}n$

<sup>1</sup> Average of KOZYREV 16 and KOZYREV 18 values taking into account the correlated uncertainties. Supersedes individual KOZYREV 16 and KOZYREV 18 results.

<sup>2</sup> Using a vector meson dominance model with contribution from  $\phi(1020)$  and higher mass excitations of  $\rho(770)$ ,  $\omega(782)$ , and  $\phi(1020)$ .

<sup>3</sup> Using a phenomenological model based on KUHN 90 with a sum of Breit-Wigner resonances for  $\rho(770)$ ,  $\omega(782)$ ,  $\phi(1020)$  and their higher mass excitations.

<sup>4</sup> Update of AKHMETSHIN 99D

<sup>5</sup> From the combined fit assuming that the total  $\phi(1020)$  production cross section is saturated by those of  $K^+K^-$ ,  $K_S K_L$ ,  $\pi^+\pi^-\pi^0$ , and  $\eta\gamma$  decays modes and using ACHASOV 00B for the  $\eta\gamma$  decay mode.

<sup>6</sup> Width errors enlarged by us to  $4\Gamma/\sqrt{N}$ ; see the note with the  $K^*(892)$  mass.

<sup>7</sup> The values were extracted from a dispersively improved Breit-Wigner parameterization.

<sup>8</sup> The values were extracted from a dispersively improved Breit-Wigner parameterization and do not include vacuum polarization.

<sup>9</sup> Strongly correlated with AKHMETSHIN 04.

<sup>10</sup> Systematic errors not evaluated.

### $\phi(1020)$ DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor/ Confidence level
$\Gamma_1$ $K^+ K^-$	(49.1 $\pm$ 0.5 ) %	S=1.3
$\Gamma_2$ $K_L^0 K_S^0$	(33.9 $\pm$ 0.4 ) %	S=1.2
$\Gamma_3$ $\rho\pi + \pi^+\pi^-\pi^0$	(15.4 $\pm$ 0.4 ) %	S=1.2
$\Gamma_4$ $\rho\pi$		
$\Gamma_5$ $\pi^+\pi^-\pi^0$		
$\Gamma_6$ $\eta\gamma$	( 1.301 $\pm$ 0.025) %	S=1.2
$\Gamma_7$ $\pi^0\gamma$	( 1.32 $\pm$ 0.05 ) $\times 10^{-3}$	
$\Gamma_8$ $\ell^+\ell^-$	—	
$\Gamma_9$ $e^+e^-$	( 2.979 $\pm$ 0.033) $\times 10^{-4}$	S=1.3
$\Gamma_{10}$ $\mu^+\mu^-$	( 2.85 $\pm$ 0.19 ) $\times 10^{-4}$	
$\Gamma_{11}$ $\eta e^+e^-$	( 1.08 $\pm$ 0.04 ) $\times 10^{-4}$	
$\Gamma_{12}$ $\pi^+\pi^-$	( 7.3 $\pm$ 1.3 ) $\times 10^{-5}$	
$\Gamma_{13}$ $\omega\pi^0$	( 4.7 $\pm$ 0.5 ) $\times 10^{-5}$	
$\Gamma_{14}$ $\omega\gamma$	< 5 %	CL=84%
$\Gamma_{15}$ $\rho\gamma$	< 1.2 $\times 10^{-5}$	CL=90%
$\Gamma_{16}$ $\pi^+\pi^-\gamma$	( 4.1 $\pm$ 1.3 ) $\times 10^{-5}$	
$\Gamma_{17}$ $f_0(980)\gamma$	( 3.22 $\pm$ 0.19 ) $\times 10^{-4}$	S=1.1
$\Gamma_{18}$ $\pi^0\pi^0\gamma$	( 1.12 $\pm$ 0.06 ) $\times 10^{-4}$	
$\Gamma_{19}$ $\pi^+\pi^-\pi^+\pi^-$	( 3.9 $^{+2.8}_{-2.2}$ ) $\times 10^{-6}$	
$\Gamma_{20}$ $\pi^+\pi^+\pi^-\pi^-\pi^0$	< 4.6 $\times 10^{-6}$	CL=90%
$\Gamma_{21}$ $\pi^0 e^+ e^-$	( 1.33 $^{+0.07}_{-0.10}$ ) $\times 10^{-5}$	
$\Gamma_{22}$ $\pi^0\eta\gamma$	( 7.27 $\pm$ 0.30 ) $\times 10^{-5}$	S=1.5
$\Gamma_{23}$ $a_0(980)\gamma$	( 7.6 $\pm$ 0.6 ) $\times 10^{-5}$	
$\Gamma_{24}$ $K^0\bar{K}^0\gamma$	< 1.9 $\times 10^{-8}$	CL=90%
$\Gamma_{25}$ $\eta'(958)\gamma$	( 6.21 $\pm$ 0.21 ) $\times 10^{-5}$	
$\Gamma_{26}$ $\eta\pi^0\pi^0\gamma$	< 2 $\times 10^{-5}$	CL=90%
$\Gamma_{27}$ $\mu^+\mu^-\gamma$	( 1.4 $\pm$ 0.5 ) $\times 10^{-5}$	
$\Gamma_{28}$ $\rho\gamma\gamma$	< 1.2 $\times 10^{-4}$	CL=90%
$\Gamma_{29}$ $\eta\pi^+\pi^-$	< 1.8 $\times 10^{-5}$	CL=90%
$\Gamma_{30}$ $\eta\mu^+\mu^-$	< 9.4 $\times 10^{-6}$	CL=90%
$\Gamma_{31}$ $\eta U \rightarrow \eta e^+ e^-$	< 1 $\times 10^{-6}$	CL=90%
$\Gamma_{32}$ invisible	< 1.7 $\times 10^{-4}$	CL=90%

### Lepton Family number (*LF*) violating modes

$$\Gamma_{33} \quad e^\pm \mu^\mp \quad LF < 2 \quad \times 10^{-6} \quad CL=90\%$$

#### CONSTRAINED FIT INFORMATION

An overall fit to 30 branching ratios uses 80 measurements and one constraint to determine 14 parameters. The overall fit has a  $\chi^2 = 61.8$  for 67 degrees of freedom.

The following *off-diagonal* array elements are the correlation coefficients  $\langle \delta x_i \delta x_j \rangle / (\delta x_i \cdot \delta x_j)$ , in percent, from the fit to the branching fractions,  $x_i \equiv \Gamma_i / \Gamma_{\text{total}}$ . The fit constrains the  $x_i$  whose labels appear in this array to sum to one.

$x_2$	-73									
$x_3$	-60	-10								
$x_6$	-21	18	2							
$x_7$	-11	11	2	8						
$x_9$	48	-51	-8	-37	-22					
$x_{10}$	-6	6	1	4	3	-12				
$x_{12}$	-3	3	0	2	1	-6	1			
$x_{13}$	-4	4	1	3	2	-8	1	0		
$x_{17}$	0	0	0	0	0	0	0	0	0	
$x_{18}$	-10	10	1	18	4	-19	2	1	2	0
$x_{19}$	-1	1	0	1	0	-2	0	0	0	0
$x_{23}$	0	0	0	0	0	0	0	0	0	0
$x_{25}$	-7	6	1	33	3	-12	1	1	1	0
	$x_1$	$x_2$	$x_3$	$x_6$	$x_7$	$x_9$	$x_{10}$	$x_{12}$	$x_{13}$	$x_{17}$
$x_{19}$	0									
$x_{23}$	0	0								
$x_{25}$	6	0	0							
	$x_{18}$	$x_{19}$	$x_{23}$							

#### $\phi(1020)$ PARTIAL WIDTHS

$$\Gamma(\eta\gamma) \quad \Gamma_6$$

<i>VALUE</i> (keV)	<i>DOCUMENT ID</i>	<i>TECN</i>	<i>COMMENT</i>
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• • • We do not use the following data for averages, fits, limits, etc. • • •

58.9±0.5±2.4	ACHASOV	00	SND $e^+ e^- \rightarrow \eta\gamma$
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$\Gamma(\pi^0\gamma)$   $\Gamma_7$

VALUE (keV)	DOCUMENT ID	TECN	COMMENT
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• • • We do not use the following data for averages, fits, limits, etc. • • •

$5.40 \pm 0.16^{+0.43}_{-0.40}$	ACHASOV	00	SND $e^+e^- \rightarrow \pi^0\gamma$
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$\Gamma(\ell^+\ell^-)$   $\Gamma_8$

VALUE (keV)	DOCUMENT ID	TECN	COMMENT
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• • • We do not use the following data for averages, fits, limits, etc. • • •

$1.320 \pm 0.017 \pm 0.015$	<sup>1</sup> AMBROSINO	05	KLOE $1.02 e^+e^- \rightarrow \mu^+\mu^-$
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<sup>1</sup>Weighted average of  $\Gamma_{ee}$  and  $\sqrt{\Gamma_{ee}\Gamma_{\mu\mu}}$  from AMBROSINO 05 assuming lepton universality.

$\Gamma(e^+e^-)$   $\Gamma_9$

VALUE (keV)	DOCUMENT ID	TECN	COMMENT
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**1.27 ± 0.04 OUR EVALUATION**

**1.251 ± 0.021 OUR AVERAGE** Error includes scale factor of 1.1.

$1.235 \pm 0.006 \pm 0.022$	<sup>1</sup> AKHMETSHIN	11	CMD2 $1.02 e^+e^- \rightarrow \phi$
$1.32 \pm 0.05 \pm 0.03$	<sup>2</sup> AMBROSINO	05	KLOE $1.02 e^+e^- \rightarrow e^+e^-$
$1.28 \pm 0.05$	AKHMETSHIN	95	CMD2 $1.02 e^+e^- \rightarrow \phi$

<sup>1</sup>Combined analysis of the CMD-2 data on  $\phi \rightarrow K^+K^-, K_S^0K_L^0, \pi^+\pi^-\pi^0, \eta\gamma$  assuming that the sum of their branching fractions is  $0.99741 \pm 0.00007$ .

<sup>2</sup>From forward-backward asymmetry and using  $\Gamma_{\text{total}} = 4.26 \pm 0.05$  MeV from the 2004 edition of this Review.

$(\Gamma(e^+e^-) \times \Gamma(\mu^+\mu^-))^{1/2}$   $(\Gamma_9\Gamma_{10})^{1/2}$

VALUE (keV)	DOCUMENT ID	TECN	COMMENT
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<b>1.320 ± 0.018 ± 0.017</b>	AMBROSINO	05	KLOE $1.02 e^+e^- \rightarrow \mu^+\mu^-$
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$\phi(1020) \Gamma(i)\Gamma(e^+e^-)/\Gamma(\text{total})$

$\Gamma(K^+K^-) \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$   $\Gamma_1\Gamma_9/\Gamma$

VALUE (keV)	EVTS	DOCUMENT ID	TECN	COMMENT
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<b>0.6340 ± 0.0070 ± 0.0039</b>		<sup>1</sup> LEES	13Q BABR	$e^+e^- \rightarrow K^+K^-\gamma$
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• • • We do not use the following data for averages, fits, limits, etc. • • •

$0.669 \pm 0.001 \pm 0.023$	1.7M	KOZYREV	18	CMD3 $e^+e^- \rightarrow K^+K^-$
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<sup>1</sup>Using a phenomenological model based on KUHN 90 with a sum of Breit-Wigner resonances for  $\rho(770)$ ,  $\omega(782)$ ,  $\phi(1020)$  and their higher mass excitations. The first error combines statistical and systematic uncertainties. The second one is due to the parametrization of the charged kaon form factor and mass calibration.

$\Gamma(K_S^0K_S^0) \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$   $\Gamma_2\Gamma_9/\Gamma$

VALUE (keV)	EVTS	DOCUMENT ID	TECN	COMMENT
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<b>0.4200 ± 0.0033 ± 0.0123</b>	28k	<sup>1</sup> LEES	14H BABR	$e^+e^- \rightarrow K_S^0K_L^0\gamma$
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<sup>1</sup>Using a vector meson dominance model with contribution from  $\phi(1020)$  and higher mass excitations of  $\rho(770)$ ,  $\omega(782)$ , and  $\phi(1020)$ .

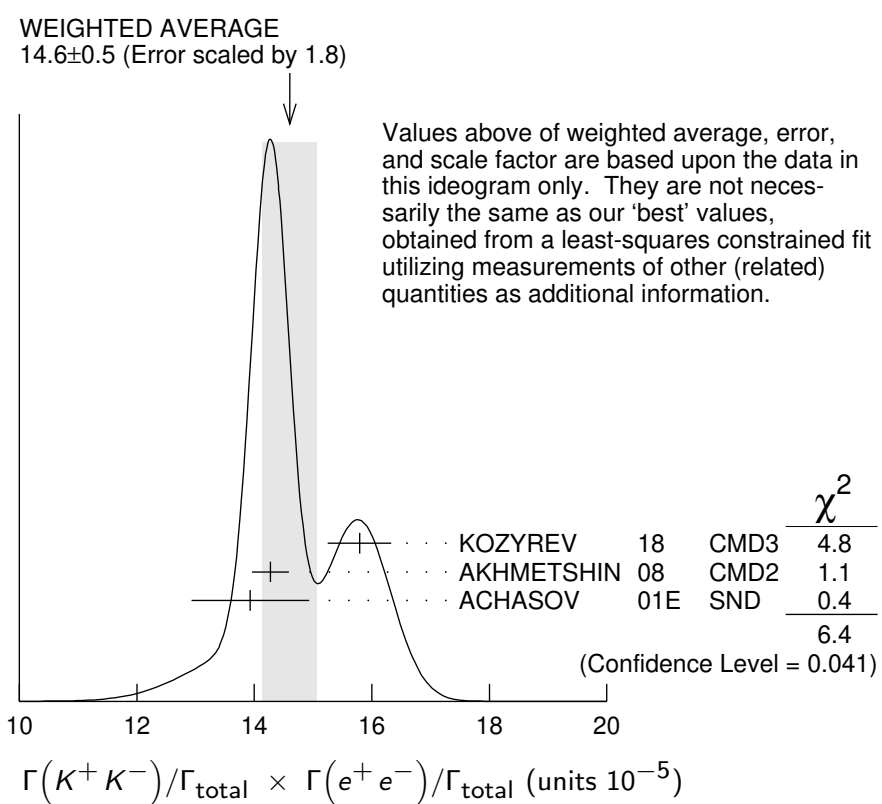
$[\Gamma(\rho\pi) + \Gamma(\pi^+\pi^-\pi^0)] \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$				$\Gamma_3\Gamma_9/\Gamma$
VALUE (eV)	DOCUMENT ID	TECN	COMMENT	
<b>184.1±2.1±8.0</b>	<sup>1</sup> LEES	21B BABR	10.5 $e^+e^- \rightarrow \pi^+\pi^-\pi^0\gamma$	

<sup>1</sup> From the cross section for  $e^+e^- \rightarrow \pi^+\pi^-\pi^0$  with contributions from  $\rho(770)$ ,  $\omega(782)$ ,  $\phi(1020)$ ,  $\omega(1420)$ , and  $\omega(1650)$ .

$\phi(1020) \Gamma(i)\Gamma(e^+e^-)/\Gamma^2(\text{total})$				
$\Gamma(K^+K^-)/\Gamma_{\text{total}} \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$				$\Gamma_1/\Gamma \times \Gamma_9/\Gamma$
VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	TECN	COMMENT

<b>14.64 ±0.28 OUR FIT</b>	Error includes scale factor of 1.4.			
<b>14.6 ±0.5 OUR AVERAGE</b>	Error includes scale factor of 1.8. See the ideogram below.			
15.789±0.541	1.7M	KOZYREV	18	CMD3 $e^+e^- \rightarrow K^+K^-$
14.27 ±0.05 ±0.31	542k	AKHMETSHIN	08	CMD2 $e^+e^- \rightarrow K^+K^-$
13.93 ±0.14 ±0.99	1000k	<sup>1</sup> ACHASOV	01E	SND $e^+e^- \rightarrow K^+K^-$ , $K_S K_L, \pi^+\pi^-\pi^0$

<sup>1</sup> From the combined fit assuming that the total  $\phi(1020)$  production cross section is saturated by those of  $K^+K^-$ ,  $K_S K_L$ ,  $\pi^+\pi^-\pi^0$ , and  $\eta\gamma$  decays modes and using ACHASOV 00B for the  $\eta\gamma$  decay mode.



$\Gamma(K_L^0 K_S^0)/\Gamma_{\text{total}} \times \Gamma(e^+ e^-)/\Gamma_{\text{total}}$   $\Gamma_2/\Gamma \times \Gamma_9/\Gamma$

VALUE (units  $10^{-5}$ )    EVTS    DOCUMENT ID    TECN    COMMENT

**10.11 ±0.12 OUR FIT**

**10.07 ±0.13 OUR AVERAGE**

10.078 ± 0.223	610k	<sup>1</sup> KOZYREV	16	CMD3	$e^+ e^- \rightarrow K_S^0 K_L^0$
10.01 ± 0.04 ± 0.17	272k	<sup>2</sup> AKHMETSHIN	04	CMD2	$e^+ e^- \rightarrow K_L^0 K_S^0$
10.27 ± 0.07 ± 0.34	500k	<sup>3</sup> ACHASOV	01E	SND	$e^+ e^- \rightarrow K^+ K^-, K_S K_L, \pi^+ \pi^- \pi^0$

<sup>1</sup> KOZYREV 16 also reports  $\Gamma(e^+ e^-) B(\phi \rightarrow K_S^0 K_L^0) = (0.428 \pm 0.001 \pm 0.009)$  keV.

<sup>2</sup> Update of AKHMETSHIN 99D

<sup>3</sup> From the combined fit assuming that the total  $\phi(1020)$  production cross section is saturated by those of  $K^+ K^-$ ,  $K_S K_L$ ,  $\pi^+ \pi^- \pi^0$ , and  $\eta\gamma$  decays modes and using ACHASOV 00B for the  $\eta\gamma$  decay mode.

$[\Gamma(\rho\pi) + \Gamma(\pi^+ \pi^- \pi^0)]/\Gamma_{\text{total}} \times \Gamma(e^+ e^-)/\Gamma_{\text{total}}$   $\Gamma_3/\Gamma \times \Gamma_9/\Gamma$

VALUE (units  $10^{-5}$ )    EVTS    DOCUMENT ID    TECN    COMMENT

**4.58 ±0.11 OUR FIT** Error includes scale factor of 1.1.

**4.51 ±0.14 OUR AVERAGE**

4.51 ± 0.16 ± 0.11	105k	AKHMETSHIN	06	CMD2	0.98–1.06 $e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$
4.665 ± 0.042 ± 0.261	400k	<sup>1</sup> ACHASOV	01E	SND	$e^+ e^- \rightarrow K^+ K^-, K_S K_L, \pi^+ \pi^- \pi^0$
4.35 ± 0.27 ± 0.08	11169	<sup>2</sup> AKHMETSHIN	98	CMD2	$e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$
• • •		We do not use the following data for averages, fits, limits, etc. • • •			
4.38 ± 0.12		BENAYOUN	10	RVUE	0.4–1.05 $e^+ e^-$
4.30 ± 0.08 ± 0.21		<sup>3</sup> AUBERT,B	04N	BABR	10.6 $e^+ e^- \rightarrow \pi^+ \pi^- \pi^0 \gamma$

<sup>1</sup> From the combined fit assuming that the total  $\phi(1020)$  production cross section is saturated by those of  $K^+ K^-$ ,  $K_S K_L$ ,  $\pi^+ \pi^- \pi^0$ , and  $\eta\gamma$  decays modes and using ACHASOV 00B for the  $\eta\gamma$  decay mode.

<sup>2</sup> Recalculated by us from the cross section in the peak.

<sup>3</sup> Superseded by LEES 21B.

$\Gamma(\eta\gamma)/\Gamma_{\text{total}} \times \Gamma(e^+ e^-)/\Gamma_{\text{total}}$   $\Gamma_6/\Gamma \times \Gamma_9/\Gamma$

VALUE (units  $10^{-6}$ )    EVTS    DOCUMENT ID    TECN    COMMENT

**3.88 ±0.07 OUR FIT** Error includes scale factor of 1.2.

**3.93 ±0.09 OUR AVERAGE** Error includes scale factor of 1.3. See the ideogram below.

4.050 ± 0.067 ± 0.118	33k	<sup>1</sup> ACHASOV	07B	SND	0.6–1.38 $e^+ e^- \rightarrow \eta\gamma$
4.093 <sup>+0.040</sup> <sub>-0.043</sub> ± 0.247	17.4k	<sup>2</sup> AKHMETSHIN	05	CMD2	0.60–1.38 $e^+ e^- \rightarrow \eta\gamma$
3.850 ± 0.041 ± 0.159	23k	<sup>3,4</sup> AKHMETSHIN	01B	CMD2	$e^+ e^- \rightarrow \eta\gamma$
4.00 ± 0.04 ± 0.11		<sup>5</sup> ACHASOV	00	SND	$e^+ e^- \rightarrow \eta\gamma$
3.53 ± 0.08 ± 0.17	2200	<sup>6,7</sup> AKHMETSHIN	99F	CMD2	$e^+ e^- \rightarrow \eta\gamma$

• • • We do not use the following data for averages, fits, limits, etc. • • •

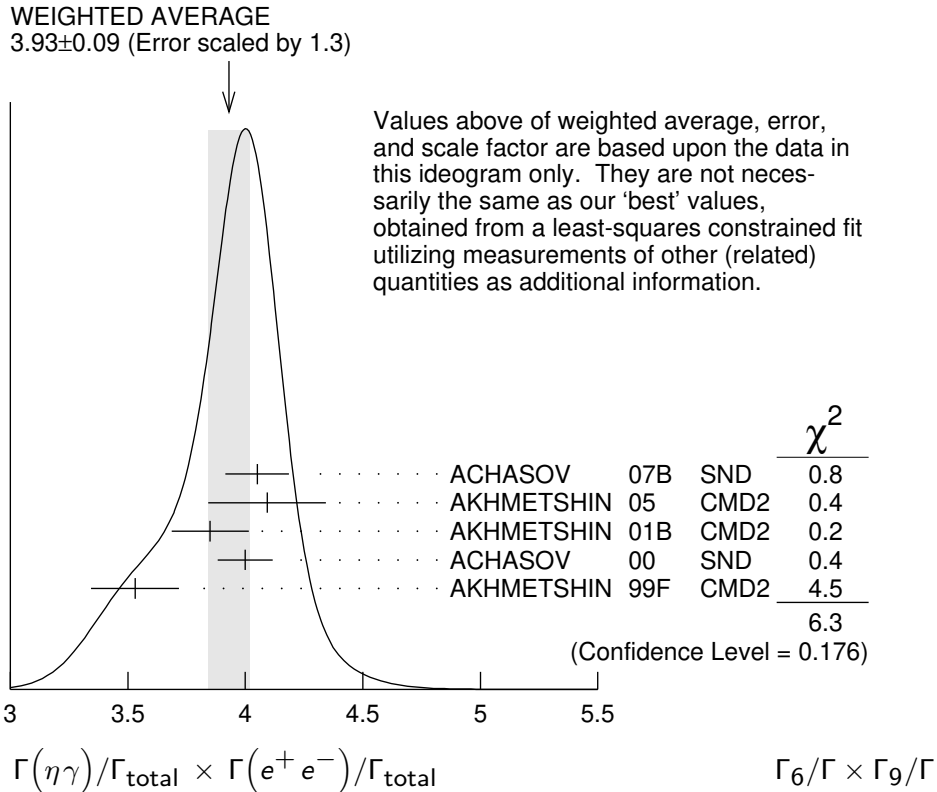
4.19 ± 0.06 <sup>8</sup> BENAYOUN 10 RVUE 0.4–1.05  $e^+ e^-$

<sup>1</sup> From a combined fit of  $\sigma(e^+ e^- \rightarrow \eta\gamma)$  with  $\eta \rightarrow 3\pi^0$  and  $\eta \rightarrow \pi^+ \pi^- \pi^0$ , and fixing  $B(\eta \rightarrow 3\pi^0) / B(\eta \rightarrow \pi^+ \pi^- \pi^0) = 1.44 \pm 0.04$ . Recalculated by us from the cross section at the peak. Supersedes ACHASOV 00D and ACHASOV 06A.

<sup>2</sup> From the  $\eta \rightarrow 2\gamma$  decay and using  $B(\eta \rightarrow \gamma\gamma) = 39.43 \pm 0.26\%$ .



- <sup>3</sup> From the  $\eta \rightarrow 3\pi^0$  decay and using  $B(\eta \rightarrow 3\pi^0) = (32.24 \pm 0.29) \times 10^{-2}$ .
- <sup>4</sup> The combined fit from 600 to 1380 MeV taking into account  $\rho(770)$ ,  $\omega(782)$ ,  $\phi(1020)$ , and  $\rho(1450)$  (mass and width fixed at 1450 MeV and 310 MeV respectively).
- <sup>5</sup> From the  $\eta \rightarrow 2\gamma$  decay and using  $B(\eta \rightarrow 2\gamma) = (39.21 \pm 0.34) \times 10^{-2}$ .
- <sup>6</sup> Recalculated by the authors from the cross section in the peak.
- <sup>7</sup> From the  $\eta \rightarrow \pi^+\pi^-\pi^0$  decay and using  $B(\eta \rightarrow \pi^+\pi^-\pi^0) = (23.1 \pm 0.5) \times 10^{-2}$ .
- <sup>8</sup> A simultaneous fit of  $e^+e^- \rightarrow \pi^+\pi^-, \pi^+\pi^-\pi^0, \pi^0\gamma, \eta\gamma$  data.



$\Gamma(\pi^0\gamma)/\Gamma_{\text{total}} \times \Gamma(e^+e^-)/\Gamma_{\text{total}} \qquad \Gamma_7/\Gamma \times \Gamma_9/\Gamma$

VALUE (units $10^{-7}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>3.94±0.16 OUR FIT</b>				
<b>3.95±0.17 OUR AVERAGE</b>				
4.04±0.09±0.19		<sup>1</sup> ACHASOV 16A	SND	0.60–1.38 $e^+e^- \rightarrow \pi^0\gamma$
3.75±0.11±0.29	18k	AKHMETSHIN 05	CMD2	0.60–1.38 $e^+e^- \rightarrow \pi^0\gamma$
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
4.29±0.11		<sup>2</sup> BENAYOUN 10	RVUE	0.4–1.05 $e^+e^-$
3.67±0.10 <sup>+0.27</sup> <sub>-0.25</sub>		<sup>3</sup> ACHASOV 00	SND	$e^+e^- \rightarrow \pi^0\gamma$

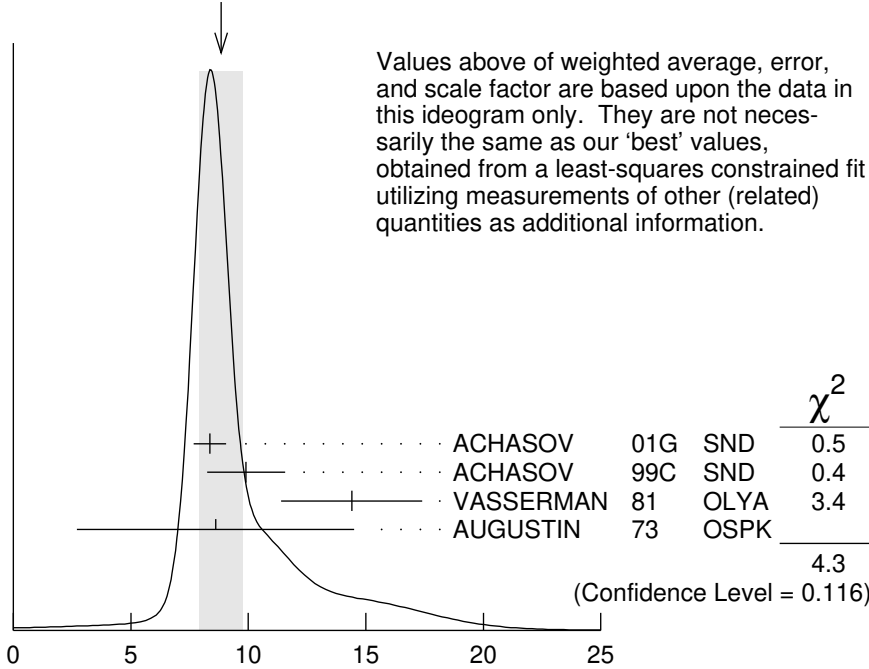
- <sup>1</sup> From the VMD model with the interfering  $\rho(770)$ ,  $\omega(782)$ ,  $\phi(1020)$  resonances, and an additional resonance describing the total contribution of the  $\rho(1450)$  and  $\omega(1420)$  states. Supersedes ACHASOV 00.
- <sup>2</sup> A simultaneous fit of  $e^+e^- \rightarrow \pi^+\pi^-, \pi^+\pi^-\pi^0, \pi^0\gamma, \eta\gamma$  data.
- <sup>3</sup> From the  $\pi^0 \rightarrow 2\gamma$  decay and using  $B(\pi^0 \rightarrow 2\gamma) = (98.798 \pm 0.032) \times 10^{-2}$ .

$\Gamma(\mu^+ \mu^-)/\Gamma_{\text{total}} \times \Gamma(e^+ e^-)/\Gamma_{\text{total}}$   $\Gamma_{10}/\Gamma \times \Gamma_9/\Gamma$

VALUE (units $10^{-8}$ )	DOCUMENT ID	TECN	COMMENT
<b>8.5 ± 0.6 OUR FIT</b>			
<b>8.8 ± 0.9 OUR AVERAGE</b>	Error includes scale factor of 1.5. See the ideogram below.		
8.36 ± 0.59 ± 0.37	ACHASOV	01G SND	$e^+ e^- \rightarrow \mu^+ \mu^-$
9.9 ± 1.4 ± 0.9	<sup>1</sup> ACHASOV	99C SND	$e^+ e^- \rightarrow \mu^+ \mu^-$
14.4 ± 3.0	<sup>2</sup> VASSERMAN	81 OLYA	$e^+ e^- \rightarrow \mu^+ \mu^-$
8.6 ± 5.9	<sup>2</sup> AUGUSTIN	73 OSPK	$e^+ e^- \rightarrow \mu^+ \mu^-$

<sup>1</sup> Recalculated by the authors from the cross section in the peak.  
<sup>2</sup> Recalculated by us from the cross section in the peak.

WEIGHTED AVERAGE  
 8.8 ± 0.9 (Error scaled by 1.5)



$\Gamma(\pi^+ \pi^-)/\Gamma_{\text{total}} \times \Gamma(e^+ e^-)/\Gamma_{\text{total}}$   $\Gamma_{12}/\Gamma \times \Gamma_9/\Gamma$

VALUE (units $10^{-8}$ )	DOCUMENT ID	TECN	COMMENT
<b>2.2 ± 0.4 OUR FIT</b>			
<b>2.2 ± 0.4 OUR AVERAGE</b>			
2.1 ± 0.3 ± 0.3	<sup>1</sup> ACHASOV	00C SND	$e^+ e^- \rightarrow \pi^+ \pi^-$
1.95 <sup>+1.15</sup> <sub>-0.87</sub>	<sup>2</sup> GOLUBEV	86 ND	$e^+ e^- \rightarrow \pi^+ \pi^-$
6.01 <sup>+3.19</sup> <sub>-2.51</sub>	<sup>2</sup> VASSERMAN	81 OLYA	$e^+ e^- \rightarrow \pi^+ \pi^-$

• • • We do not use the following data for averages, fits, limits, etc. • • •

3.31 ± 0.99	<sup>3</sup> BENAYOUN	13 RVUE	0.4–1.05 $e^+ e^-$
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<sup>1</sup> Recalculated by the authors from the cross section in the peak.  
<sup>2</sup> Recalculated by us from the cross section in the peak.  
<sup>3</sup> A simultaneous fit to  $e^+ e^- \rightarrow \pi^+ \pi^-, \pi^+ \pi^- \pi^0, \pi^0 \gamma, \eta \gamma, K \bar{K}$ , and  $\tau^- \rightarrow \pi^- \pi^0 \nu_\tau$  data.

$\Gamma(\omega\pi^0)/\Gamma_{\text{total}} \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$				$\Gamma_{13}/\Gamma \times \Gamma_9/\Gamma$
<u>VALUE (units <math>10^{-8}</math>)</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	
<b>1.40±0.15 OUR FIT</b>				
<b>1.37±0.17±0.01</b>	1,2	AMBROSINO 08G	KLOE	$e^+e^- \rightarrow \pi^+\pi^-2\pi^0, 2\pi^0\gamma$

<sup>1</sup> Recalculated by the authors from the cross section at the peak.

<sup>2</sup> AMBROSINO 08G reports  $[\Gamma(\phi(1020) \rightarrow \omega\pi^0)/\Gamma_{\text{total}} \times \Gamma(\phi(1020) \rightarrow e^+e^-)/\Gamma_{\text{total}}] \times [B(\omega(782) \rightarrow \pi^+\pi^-\pi^0)] = (1.22 \pm 0.13 \pm 0.08) \times 10^{-8}$  which we divide by our best value  $B(\omega(782) \rightarrow \pi^+\pi^-\pi^0) = (89.2 \pm 0.7) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(\pi^0\pi^0\gamma)/\Gamma_{\text{total}} \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$				$\Gamma_{18}/\Gamma \times \Gamma_9/\Gamma$
<u>VALUE (units <math>10^{-8}</math>)</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	
<b>3.34±0.17 OUR FIT</b>				
<b>3.33<sup>+0.04+0.19</sup><sub>-0.09-0.20</sub></b>	1	AMBROSINO 07	KLOE	$e^+e^- \rightarrow \pi^0\pi^0\gamma$

<sup>1</sup> Calculated by the authors from the cross section at the peak.

$\Gamma(\pi^+\pi^-\pi^+\pi^-)/\Gamma_{\text{total}} \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$				$\Gamma_{19}/\Gamma \times \Gamma_9/\Gamma$
<u>VALUE (units <math>10^{-9}</math>)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>1.2<sup>+0.8</sup><sub>-0.7</sub> OUR FIT</b>				
<b>1.17±0.52±0.64</b>	3285	1	AKHMETSHIN 00E	CMD2 $e^+e^- \rightarrow \pi^+\pi^-\pi^+\pi^-$

<sup>1</sup> Recalculated by the authors from the cross section in the peak.

### $\phi(1020)$ BRANCHING RATIOS

$\Gamma(K^+K^-)/\Gamma_{\text{total}}$				$\Gamma_1/\Gamma$
<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.491±0.005 OUR FIT</b>				Error includes scale factor of 1.3.
<b>0.493±0.010 OUR AVERAGE</b>				
0.492±0.012	2913	AKHMETSHIN 95	CMD2	$e^+e^- \rightarrow K^+K^-$
0.44 ±0.05	321	KALBFLEISCH 76	HBC	2.18 $K^-p \rightarrow \Lambda K^+K^-$
0.49 ±0.06	270	DEGROOT 74	HBC	4.2 $K^-p \rightarrow \Lambda\phi$
0.540±0.034	565	BALAKIN 71	OSPK	$e^+e^- \rightarrow K^+K^-$
0.48 ±0.04	252	LINDSEY 66	HBC	2.1-2.7 $K^-p \rightarrow \Lambda K^+K^-$
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
0.493±0.003±0.007		1	AKHMETSHIN 11	CMD2 $1.02 e^+e^- \rightarrow K^+K^-$
0.476±0.017	1000k	2	ACHASOV 01E	SND $e^+e^- \rightarrow K^+K^-, K_S K_L, \pi^+\pi^-\pi^0$

<sup>1</sup> Combined analysis of the CMD-2 data on  $\phi \rightarrow K^+K^-, K_S^0 K_L^0, \pi^+\pi^-\pi^0, \eta\gamma$  assuming that the sum of their branching fractions is  $0.99741 \pm 0.00007$ .

<sup>2</sup> Using  $B(\phi \rightarrow e^+e^-) = (2.93 \pm 0.14) \times 10^{-4}$ .

$\Gamma(K_L^0 K_S^0)/\Gamma_{\text{total}}$   $\Gamma_2/\Gamma$

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.339±0.004 OUR FIT</b>				Error includes scale factor of 1.2.
<b>0.331±0.009 OUR AVERAGE</b>				
0.335±0.010	40644	AKHMETSHIN 95	CMD2	$e^+e^- \rightarrow K_L^0 K_S^0$
0.326±0.035		DOLINSKY 91	ND	$e^+e^- \rightarrow K_L^0 K_S^0$
0.310±0.024		DRUZHININ 84	ND	$e^+e^- \rightarrow K_L^0 K_S^0$
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
0.336±0.002±0.006		<sup>1</sup> AKHMETSHIN 11	CMD2	1.02 $e^+e^- \rightarrow K_S^0 K_L^0$
0.351±0.013	500k	<sup>2</sup> ACHASOV 01E	SND	$e^+e^- \rightarrow K^+K^-, K_S^0 K_L^0, \pi^+\pi^-\pi^0$
0.27 ±0.03	133	KALBFLEISCH 76	HBC	2.18 $K^-p \rightarrow \Lambda K_L^0 K_S^0$
0.257±0.030	95	<sup>3</sup> BALAKIN 71	OSPK	$e^+e^- \rightarrow K_L^0 K_S^0$
0.40 ±0.04	167	LINDSEY 66	HBC	2.1–2.7 $K^-p \rightarrow \Lambda K_L^0 K_S^0$

<sup>1</sup> Combined analysis of the CMD-2 data on  $\phi \rightarrow K^+K^-, K_S^0 K_L^0, \pi^+\pi^-\pi^0, \eta\gamma$  assuming that the sum of their branching fractions is  $0.99741 \pm 0.00007$ .

<sup>2</sup> Using  $B(\phi \rightarrow e^+e^-) = (2.93 \pm 0.14) \times 10^{-4}$ .

<sup>3</sup> Balakin error increased by Paul.

$\Gamma(K_L^0 K_S^0)/\Gamma(K^+K^-)$   $\Gamma_2/\Gamma_1$

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.690±0.015 OUR FIT</b>				Error includes scale factor of 1.3.
<b>0.740±0.031 OUR AVERAGE</b>				
0.70 ±0.06	2732	BUKIN 78c	OLYA	$e^+e^- \rightarrow K_L^0 K_S^0$
0.82 ±0.08		LOSTY 78	HBC	4.2 $K^-p \rightarrow \phi$ hyperon
0.71 ±0.05		LAVEN 77	HBC	10 $K^-p \rightarrow K^+K^-\Lambda$
0.71 ±0.08		LYONS 77	HBC	3–4 $K^-p \rightarrow \Lambda\phi$
0.89 ±0.10	144	AGUILAR-... 72B	HBC	3.9,4.6 $K^-p$
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
0.638±0.022	2.3M	<sup>1</sup> KOZYREV 18	CMD3	$e^+e^- \rightarrow K_L^0 K_S^0, K^+K^-$
0.68 ±0.03		<sup>2</sup> AKHMETSHIN 95	CMD2	$e^+e^- \rightarrow K_L^0 K_S^0, K^+K^-$

<sup>1</sup> The prediction taking into account phase-space difference, radiative corrections, isospin breaking, and the Sommerfeld-Gamow-Sakharov factor gives 0.630.

<sup>2</sup> Theoretical analysis of BRAMON 00 taking into account phase-space difference, electromagnetic radiative corrections, as well as isospin breaking, predicts 0.62. FLOREZ-BAEZ 08 predicts 0.63 considering also structure-dependent radiative corrections. FISCHBACH 02 calculates additional corrections caused by the close threshold and predicts 0.68. See also BENAYOUN 01 and DUBYNSKIY 07. BENAYOUN 12 obtains  $0.71 \pm 0.01$  in the HLS model.

$\Gamma(K_L^0 K_S^0)/\Gamma(K\bar{K})$   $\Gamma_2/(\Gamma_1+\Gamma_2)$

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.408±0.005 OUR FIT</b>				Error includes scale factor of 1.3.
<b>0.45 ±0.04 OUR AVERAGE</b>				
0.44 ±0.07		<sup>1</sup> LONDON 66	HBC	2.24 $K^-p \rightarrow \Lambda K\bar{K}$
0.48 ±0.07	52	BADIER 65B	HBC	3 $K^-p$
0.40 ±0.10	34	SCHLEIN 63	HBC	1.95 $K^-p \rightarrow \Lambda K\bar{K}$

<sup>1</sup> This is probably not affected by their controversial background subtraction; the value is from their numbers of  $K_1 K_2$  vs  $K^+K^-$  events.

$[\Gamma(\rho\pi) + \Gamma(\pi^+\pi^-\pi^0)]/\Gamma_{\text{total}}$   $\Gamma_3/\Gamma$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
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**0.154±0.004 OUR FIT** Error includes scale factor of 1.2.

**0.151±0.009 OUR AVERAGE** Error includes scale factor of 1.7.

0.161±0.008	11761	AKHMETSHIN 95	CMD2	$e^+e^- \rightarrow \pi^+\pi^-\pi^0$
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0.143±0.007		DOLINSKY 91	ND	$e^+e^- \rightarrow \pi^+\pi^-\pi^0$
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• • • We do not use the following data for averages, fits, limits, etc. • • •

0.155±0.002±0.005		<sup>1</sup> AKHMETSHIN 11	CMD2	1.02 $e^+e^- \rightarrow \pi^+\pi^-\pi^0$
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0.159±0.008	400k	<sup>2</sup> ACHASOV 01E	SND	$e^+e^- \rightarrow K^+K^-, K_S^0K_L^0, \pi^+\pi^-\pi^0$
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0.145±0.009±0.003	11169	<sup>3</sup> AKHMETSHIN 98	CMD2	$e^+e^- \rightarrow \pi^+\pi^-\pi^0$
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0.139±0.007		<sup>4</sup> PARROUR 76B	OSPK	$e^+e^-$
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<sup>1</sup> Combined analysis of the CMD-2 data on  $\phi \rightarrow K^+K^-, K_S^0K_L^0, \pi^+\pi^-\pi^0, \eta\gamma$  assuming that the sum of their branching fractions is  $0.99741 \pm 0.00007$ .

<sup>2</sup> Using  $B(\phi \rightarrow e^+e^-) = (2.93 \pm 0.14) \times 10^{-4}$ .

<sup>3</sup> Using  $B(\phi \rightarrow e^+e^-) = (2.99 \pm 0.08) \times 10^{-4}$ .

<sup>4</sup> Using  $\Gamma(\phi) = 4.1$  MeV. If interference between the  $\rho\pi$  and  $3\pi$  modes is neglected, the fraction of the  $\rho\pi$  is more than 80% at the 90% confidence level.

$[\Gamma(\rho\pi) + \Gamma(\pi^+\pi^-\pi^0)]/\Gamma(K^+K^-)$   $\Gamma_3/\Gamma_1$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
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**0.313±0.010 OUR FIT** Error includes scale factor of 1.2.

<b>0.28 ±0.09</b>	34	AGUILAR-...	72B HBC	3.9,4.6 $K^-p$
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$[\Gamma(\rho\pi) + \Gamma(\pi^+\pi^-\pi^0)]/\Gamma(K\bar{K})$   $\Gamma_3/(\Gamma_1+\Gamma_2)$

VALUE	DOCUMENT ID	TECN	COMMENT
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**0.185±0.005 OUR FIT** Error includes scale factor of 1.2.

**0.24 ±0.04 OUR AVERAGE**

0.237±0.039	CERRADA 77B	HBC	4.2 $K^-p \rightarrow \Lambda 3\pi$
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0.30 ±0.15	LONDON 66	HBC	2.24 $K^-p \rightarrow \Lambda \pi^+\pi^-\pi^0$
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$[\Gamma(\rho\pi) + \Gamma(\pi^+\pi^-\pi^0)]/\Gamma(K_L^0K_S^0)$   $\Gamma_3/\Gamma_2$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
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**0.453±0.012 OUR FIT** Error includes scale factor of 1.1.

**0.51 ±0.05 OUR AVERAGE**

0.56 ±0.07	3681	BUKIN 78c	OLYA	$e^+e^- \rightarrow K_L^0K_S^0, \pi^+\pi^-\pi^0$
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0.47 ±0.06	516	COSME 74	OSPK	$e^+e^- \rightarrow \pi^+\pi^-\pi^0$
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$\Gamma(\pi^+\pi^-\pi^0)/\Gamma_{\text{total}}$   $\Gamma_5/\Gamma$

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
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• • • We do not use the following data for averages, fits, limits, etc. • • •

$\simeq 0.0087$	1.98M	<sup>1,2</sup> ALOISIO 03	KLOE	1.02 $e^+e^- \rightarrow \pi^+\pi^-\pi^0$
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$<0.0006$	90	<sup>3</sup> ACHASOV 02	SND	1.02 $e^+e^- \rightarrow \pi^+\pi^-\pi^0$
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$<0.23$	90	<sup>3</sup> CORDIER 80	DM1	$e^+e^- \rightarrow \pi^+\pi^-\pi^0$
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$<0.20$	90	<sup>3</sup> PARROUR 76B	OSPK	$e^+e^- \rightarrow \pi^+\pi^-\pi^0$
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<sup>1</sup> From a fit without limitations on charged and neutral  $\rho$  masses and widths.

<sup>2</sup> Adding the direct and  $\omega\pi$  contributions and considering the interference between the  $\rho\pi$  and  $\pi^+\pi^-\pi^0$ .

<sup>3</sup> Neglecting the interference between the  $\rho\pi$  and  $\pi^+\pi^-\pi^0$ .

$\Gamma(\eta\gamma)/\Gamma_{\text{total}}$   $\Gamma_6/\Gamma$

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
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**1.301 ± 0.025 OUR FIT** Error includes scale factor of 1.2.

**1.26 ± 0.04 OUR AVERAGE**

1.246 ± 0.025 ± 0.057	10k	<sup>1</sup> ACHASOV	98F SND	$e^+e^- \rightarrow 7\gamma$
1.18 ± 0.11	279	<sup>2</sup> AKHMETSHIN	95 CMD2	$e^+e^- \rightarrow \pi^+\pi^-3\gamma$
1.30 ± 0.06		<sup>3</sup> DRUZHININ	84 ND	$e^+e^- \rightarrow 3\gamma$
1.4 ± 0.2		<sup>4</sup> DRUZHININ	84 ND	$e^+e^- \rightarrow 6\gamma$
0.88 ± 0.20	290	KURDADZE	83C OLYA	$e^+e^- \rightarrow 3\gamma$
1.35 ± 0.29		ANDREWS	77 CNTR	6.7–10 $\gamma$ Cu
1.5 ± 0.4	54	<sup>3</sup> COSME	76 OSPK	$e^+e^-$

• • • We do not use the following data for averages, fits, limits, etc. • • •

1.38 ± 0.02 ± 0.02		<sup>5</sup> AKHMETSHIN	11 CMD2	$1.02 e^+e^- \rightarrow \eta\gamma$
1.36 ± 0.05 ± 0.02	33k	<sup>6</sup> ACHASOV	07B SND	$0.6\text{--}1.38 e^+e^- \rightarrow \eta\gamma$
1.373 ± 0.014 ± 0.085	17.4k	<sup>7,8</sup> AKHMETSHIN	05 CMD2	$0.60\text{--}1.38 e^+e^- \rightarrow \eta\gamma$
1.287 ± 0.013 ± 0.063		<sup>9,10</sup> AKHMETSHIN	01B CMD2	$e^+e^- \rightarrow \eta\gamma$
1.338 ± 0.012 ± 0.052		<sup>11</sup> ACHASOV	00 SND	$e^+e^- \rightarrow \eta\gamma$
1.18 ± 0.03 ± 0.06	2200	<sup>12</sup> AKHMETSHIN	99F CMD2	$e^+e^- \rightarrow \eta\gamma$
1.21 ± 0.07		<sup>13</sup> BENAYOUN	96 RVUE	$0.54\text{--}1.04 e^+e^- \rightarrow \eta\gamma$

<sup>1</sup> Using  $B(\phi \rightarrow e^+e^-) = (2.99 \pm 0.08) \times 10^{-4}$  and  $B(\eta \rightarrow 3\pi^0) = (32.2 \pm 0.4) \times 10^{-2}$ .

<sup>2</sup> From  $\pi^+\pi^-\pi^0$  decay mode of  $\eta$ .

<sup>3</sup> From  $2\gamma$  decay mode of  $\eta$ .

<sup>4</sup> From  $3\pi^0$  decay mode of  $\eta$ .

<sup>5</sup> Combined analysis of the CMD-2 data on  $\phi \rightarrow K^+K^-, K_S^0 K_L^0, \pi^+\pi^-\pi^0, \eta\gamma$  assuming that the sum of their branching fractions is  $0.99741 \pm 0.00007$ .

<sup>6</sup> ACHASOV 07B reports  $[\Gamma(\phi(1020) \rightarrow \eta\gamma)/\Gamma_{\text{total}}] \times [B(\phi(1020) \rightarrow e^+e^-)] = (4.050 \pm 0.067 \pm 0.118) \times 10^{-6}$  which we divide by our best value  $B(\phi(1020) \rightarrow e^+e^-) = (2.979 \pm 0.033) \times 10^{-4}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value. Supersedes ACHASOV 00D and ACHASOV 06A.

<sup>7</sup> Using  $B(\phi \rightarrow e^+e^-) = (2.98 \pm 0.04) \times 10^{-4}$  and  $B(\eta \rightarrow \gamma\gamma) = 39.43 \pm 0.26\%$ .

<sup>8</sup> Not independent of the corresponding  $\Gamma(e^+e^-) \times \Gamma(\eta\gamma)/\Gamma_{\text{total}}^2$ .

<sup>9</sup> Using  $B(\phi \rightarrow e^+e^-) = (2.99 \pm 0.08) \times 10^{-4}$  and  $B(\eta \rightarrow 3\pi^0) = (32.24 \pm 0.29) \times 10^{-2}$ .

<sup>10</sup> The combined fit from 600 to 1380 MeV taking into account  $\rho(770), \omega(782), \phi(1020)$ , and  $\rho(1450)$  (mass and width fixed at 1450 MeV and 310 MeV respectively).

<sup>11</sup> From the  $\eta \rightarrow 2\gamma$  decay and using  $B(\phi \rightarrow e^+e^-) = (2.99 \pm 0.08) \times 10^{-4}$ .

<sup>12</sup> From  $\pi^+\pi^-\pi^0$  decay mode of  $\eta$  and using  $B(\phi \rightarrow e^+e^-) = (2.99 \pm 0.08) \times 10^{-4}$ .

<sup>13</sup> Reanalysis of DRUZHININ 84, DOLINSKY 89, and DOLINSKY 91 taking into account a triangle anomaly contribution.

$\Gamma(\pi^0\gamma)/\Gamma_{\text{total}}$   $\Gamma_7/\Gamma$

VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
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**1.32 ± 0.05 OUR FIT**

**1.31 ± 0.13 OUR AVERAGE**

1.30 ± 0.13		DRUZHININ	84 ND	$e^+e^- \rightarrow 3\gamma$
1.4 ± 0.5	32	COSME	76 OSPK	$e^+e^-$

• • • We do not use the following data for averages, fits, limits, etc. • • •

1.367 ± 0.072		<sup>1</sup> ACHASOV	16A SND	$0.60\text{--}1.38 e^+e^- \rightarrow \pi^0\gamma$
1.258 ± 0.037 ± 0.077	18k	<sup>2,3</sup> AKHMETSHIN	05 CMD2	$0.60\text{--}1.38 e^+e^- \rightarrow \pi^0\gamma$

$1.226 \pm 0.036^{+0.096}_{-0.089}$	<sup>4</sup> ACHASOV	00	SND	$e^+e^- \rightarrow \pi^0\gamma$
$1.26 \pm 0.17$	<sup>5</sup> BENAYOUN	96	RVUE	$0.54-1.04 e^+e^- \rightarrow \pi^0\gamma$

<sup>1</sup> Using  $B(\phi \rightarrow e^+e^-)$  from PDG 15. Supersedes ACHASOV 00.

<sup>2</sup> Using  $B(\phi \rightarrow e^+e^-) = (2.98 \pm 0.04) \times 10^{-4}$ .

<sup>3</sup> Not independent of the corresponding  $\Gamma(e^+e^-) \times \Gamma(\pi^0\gamma)/\Gamma_{\text{total}}^2$ .

<sup>4</sup> From the  $\pi^0 \rightarrow 2\gamma$  decay and using  $B(\phi \rightarrow e^+e^-) = (2.99 \pm 0.08) \times 10^{-4}$ .

<sup>5</sup> Reanalysis of DRUZHININ 84, DOLINSKY 89, and DOLINSKY 91 taking into account a triangle anomaly contribution.

### $\Gamma(\eta\gamma)/\Gamma(\pi^0\gamma)$ $\Gamma_6/\Gamma_7$

<u>VALUE</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
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• • • We do not use the following data for averages, fits, limits, etc. • • •

$10.9 \pm 0.3^{+0.7}_{-0.8}$	ACHASOV	00	SND	$e^+e^- \rightarrow \eta\gamma, \pi^0\gamma$
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### $\Gamma(e^+e^-)/\Gamma_{\text{total}}$ $\Gamma_9/\Gamma$

<u>VALUE (units <math>10^{-4}</math>)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
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**2.979 ± 0.033 OUR FIT** Error includes scale factor of 1.3.

**2.98 ± 0.07 OUR AVERAGE** Error includes scale factor of 1.1.

$2.93 \pm 0.14$	1900k	<sup>1</sup> ACHASOV	01E	SND	$e^+e^- \rightarrow K^+K^-, K_S^0K_L^0, \pi^+\pi^-\pi^0$
$2.88 \pm 0.09$	55600	AKHMETSHIN	95	CMD2	$e^+e^- \rightarrow \text{hadrons}$
$3.00 \pm 0.21$	3681	BUKIN	78C	OLYA	$e^+e^- \rightarrow \text{hadrons}$
$3.10 \pm 0.14$		<sup>2</sup> PARROUR	76	OSPK	$e^+e^-$
$3.3 \pm 0.3$		COSME	74	OSPK	$e^+e^- \rightarrow \text{hadrons}$
$2.81 \pm 0.25$	681	BALAKIN	71	OSPK	$e^+e^- \rightarrow \text{hadrons}$
$3.50 \pm 0.27$		CHATELUS	71	OSPK	$e^+e^-$

<sup>1</sup> From the combined fit assuming that the total  $\phi(1020)$  production cross section is saturated by those of  $K^+K^-$ ,  $K_S^0K_L^0$ ,  $\pi^+\pi^-\pi^0$ , and  $\eta\gamma$  decays modes and using ACHASOV 00B for the  $\eta\gamma$  decay mode.

<sup>2</sup> Using total width 4.2 MeV. They detect  $3\pi$  mode and observe significant interference with  $\omega$  tail. This is accounted for in the result quoted above.

### $\Gamma(\mu^+\mu^-)/\Gamma_{\text{total}}$ $\Gamma_{10}/\Gamma$

<u>VALUE (units <math>10^{-4}</math>)</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
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**2.85 ± 0.19 OUR FIT**

**2.5 ± 0.4 OUR AVERAGE**

$2.69 \pm 0.46$	<sup>1</sup> HAYES	71	CNTR	$8.3, 9.8 \gamma C \rightarrow \mu^+\mu^- X$
$2.17 \pm 0.60$	<sup>1</sup> EARLES	70	CNTR	$6.0 \gamma C \rightarrow \mu^+\mu^- X$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$2.87 \pm 0.20 \pm 0.14$	<sup>2</sup> ACHASOV	01G	SND	$e^+e^- \rightarrow \mu^+\mu^-$
$3.30 \pm 0.45 \pm 0.32$	<sup>3</sup> ACHASOV	99C	SND	$e^+e^- \rightarrow \mu^+\mu^-$
$4.83 \pm 1.02$	<sup>4</sup> VASSERMAN	81	OLYA	$e^+e^- \rightarrow \mu^+\mu^-$
$2.87 \pm 1.98$	<sup>4</sup> AUGUSTIN	73	OSPK	$e^+e^- \rightarrow \mu^+\mu^-$

<sup>1</sup> Neglecting interference between resonance and continuum.

<sup>2</sup> Using  $B(\phi \rightarrow e^+e^-) = (2.91 \pm 0.07) \times 10^{-4}$ .

<sup>3</sup> Using  $B(\phi \rightarrow e^+e^-) = (2.99 \pm 0.08) \times 10^{-4}$ .

<sup>4</sup> Recalculated by us using  $B(\phi \rightarrow e^+e^-) = (2.99 \pm 0.08) \times 10^{-4}$ .

$\Gamma(\eta e^+ e^-)/\Gamma_{\text{total}}$   $\Gamma_{11}/\Gamma$

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
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**1.08 ± 0.04 OUR AVERAGE**

1.075 ± 0.007 ± 0.038	30k	<sup>1</sup> BABUSCI 15	KLOE	1.02 $e^+ e^- \rightarrow \eta e^+ e^-$
1.19 ± 0.19 ± 0.12	213	<sup>2</sup> ACHASOV 01B	SND	$e^+ e^- \rightarrow \eta e^+ e^-$
1.14 ± 0.10 ± 0.06	355	<sup>3</sup> AKHMETSHIN 01	CMD2	$e^+ e^- \rightarrow \eta e^+ e^-$

• • • We do not use the following data for averages, fits, limits, etc. • • •

1.13 ± 0.14 ± 0.07	183	<sup>4</sup> AKHMETSHIN 01	CMD2	$e^+ e^- \rightarrow \eta e^+ e^-$
1.21 ± 0.14 ± 0.09	130	<sup>5</sup> AKHMETSHIN 01	CMD2	$e^+ e^- \rightarrow \eta e^+ e^-$
1.04 ± 0.20 ± 0.08	42	<sup>6</sup> AKHMETSHIN 01	CMD2	$e^+ e^- \rightarrow \eta e^+ e^-$
1.3 <sup>+0.8</sup> <sub>-0.6</sub>	7	GOLUBEV 85	ND	$e^+ e^- \rightarrow \eta e^+ e^-$

<sup>1</sup> Using  $B(\eta \rightarrow 3\pi^0) = (32.57 \pm 0.23)\%$  from PDG 12.

<sup>2</sup> Using  $B(\eta \rightarrow \gamma\gamma) = (39.25 \pm 0.32)\%$ ,  $B(\phi \rightarrow \eta\gamma) = (1.26 \pm 0.06)\%$ , and  $B(\phi \rightarrow e^+ e^-) = (3.00 \pm 0.06) \times 10^{-4}$ .

<sup>3</sup> The average of the branching ratios separately obtained from the  $\eta \rightarrow \gamma\gamma$ ,  $3\pi^0$ ,  $\pi^+ \pi^- \pi^0$  decays.

<sup>4</sup> From  $\eta \rightarrow \gamma\gamma$  decays and using  $B(\eta \rightarrow \gamma\gamma) = (39.33 \pm 0.25) \times 10^{-2}$ ,  $B(\eta \rightarrow \pi^+ \pi^- \gamma) = (4.75 \pm 11) \times 10^{-2}$ , and  $B(\phi \rightarrow \eta\gamma) = (1.297 \pm 0.033) \times 10^{-2}$ .

<sup>5</sup> From  $\eta \rightarrow 3\pi^0$  decays and using  $B(\pi^0 \rightarrow \gamma\gamma) = (98.798 \pm 0.033) \times 10^{-2}$ ,  $B(\eta \rightarrow 3\pi^0) = (32.24 \pm 0.29) \times 10^{-2}$ ,  $B(\eta \rightarrow \pi^+ \pi^- \gamma) = (4.75 \pm 0.11) \times 10^{-2}$ , and  $B(\phi \rightarrow \eta\gamma) = (1.297 \pm 0.033) \times 10^{-2}$ .

<sup>6</sup> From  $\eta \rightarrow \pi^+ \pi^- \pi^0$  decays and using  $B(\pi^0 \rightarrow \gamma\gamma) = (98.798 \pm 0.033) \times 10^{-2}$ ,  $B(\pi^0 \rightarrow e^+ e^- \gamma) = (1.198 \pm 0.032) \times 10^{-2}$ ,  $B(\eta \rightarrow \pi^+ \pi^- \pi^0) = (23.0 \pm 0.4) \times 10^{-2}$ ,  $B(\phi \rightarrow \pi^+ \pi^- \pi^0) = (15.5 \pm 0.6) \times 10^{-2}$ , and  $B(\phi \rightarrow \eta\gamma) = (1.297 \pm 0.033) \times 10^{-2}$ .

$\Gamma(\pi^+ \pi^-)/\Gamma_{\text{total}}$   $\Gamma_{12}/\Gamma$

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
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• • • We do not use the following data for averages, fits, limits, etc. • • •

0.71 ± 0.11 ± 0.09		<sup>1</sup> ACHASOV 00C	SND	$e^+ e^- \rightarrow \pi^+ \pi^-$
0.65 <sup>+0.38</sup> <sub>-0.29</sub>		<sup>1</sup> GOLUBEV 86	ND	$e^+ e^- \rightarrow \pi^+ \pi^-$
2.01 <sup>+1.07</sup> <sub>-0.84</sub>		<sup>1</sup> VASSERMAN 81	OLYA	$e^+ e^- \rightarrow \pi^+ \pi^-$
<6.6	95	BUKIN 78B	OLYA	$e^+ e^- \rightarrow \pi^+ \pi^-$
<2.7	95	ALVENSLEB... 72	CNTR	$6.7 \gamma C \rightarrow C \pi^+ \pi^-$

<sup>1</sup> Using  $B(\phi \rightarrow e^+ e^-) = (2.99 \pm 0.08) \times 10^{-4}$ .

$\Gamma(\omega\pi^0)/\Gamma_{\text{total}}$   $\Gamma_{13}/\Gamma$

VALUE (units $10^{-5}$ )	DOCUMENT ID	TECN	COMMENT
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**4.7 ± 0.5 OUR FIT**

5.2 <sup>+1.3</sup> <sub>-1.1</sub>	<sup>1,2</sup> AULCHENKO 00A	SND	$e^+ e^- \rightarrow \pi^+ \pi^- \pi^0 \pi^0$
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• • • We do not use the following data for averages, fits, limits, etc. • • •

4.4 ± 0.6	<sup>3</sup> AMBROSINO 08G	KLOE	$e^+ e^- \rightarrow \pi^+ \pi^- 2\pi^0, 2\pi^0 \gamma$
~ 5.4	<sup>4</sup> ACHASOV 00E	SND	$e^+ e^- \rightarrow \pi^0 \pi^0 \gamma$
5.5 <sup>+1.6</sup> <sub>-1.4</sub> ± 0.3	<sup>2,5</sup> AULCHENKO 00A	SND	$e^+ e^- \rightarrow \pi^+ \pi^- \pi^0 \pi^0$



$$4.8^{+1.9}_{-1.7} \pm 0.8 \quad {}^4 \text{ACHASOV} \quad 99 \quad \text{SND} \quad e^+ e^- \rightarrow \pi^+ \pi^- \pi^0 \pi^0$$

<sup>1</sup> Using the 1996 and 1998 data.

<sup>2</sup> (2.3 ± 0.3)% correction for other decay modes of the  $\omega(782)$  applied.

<sup>3</sup> Not independent of the corresponding  $\Gamma(\omega\pi^0) \times \Gamma(e^+e^-) / \Gamma^2(\text{total})$ .

<sup>4</sup> Using the 1996 data.

<sup>5</sup> Using the 1998 data.

$\Gamma(\omega\gamma)/\Gamma_{\text{total}}$						$\Gamma_{14}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT		
<0.05	84	LINDSEY	66	HBC	2.1–2.7 $K^- p \rightarrow \Lambda \pi^+ \pi^-$	neutrals

$\Gamma(\rho\gamma)/\Gamma_{\text{total}}$						$\Gamma_{15}/\Gamma$
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT		

< 0.12	90	<sup>1</sup> AKHMETSHIN 99B	CMD2	$e^+ e^- \rightarrow \pi^+ \pi^- \gamma$		
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●						
< 7	90	AKHMETSHIN 97C	CMD2	$e^+ e^- \rightarrow \pi^+ \pi^- \gamma$		
<200	84	LINDSEY	66	HBC	2.1–2.7 $K^- p \rightarrow \Lambda \pi^+ \pi^-$	neutrals

<sup>1</sup> Supersedes AKHMETSHIN 97C.

$\Gamma(\pi^+ \pi^- \gamma)/\Gamma_{\text{total}}$						$\Gamma_{16}/\Gamma$
VALUE (units $10^{-4}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT	

<b>0.41 ± 0.12 ± 0.04</b>		30175	<sup>1</sup> AKHMETSHIN 99B	CMD2	$e^+ e^- \rightarrow \pi^+ \pi^- \gamma$	
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●						
< 0.3	90		<sup>2</sup> AKHMETSHIN 97C	CMD2	$e^+ e^- \rightarrow \pi^+ \pi^- \gamma$	
<600	90		KALBFLEISCH 75	HBC	2.18 $K^- p \rightarrow \Lambda \pi^+ \pi^- \gamma$	
< 70	90		COSME	74	OSPK	$e^+ e^- \rightarrow \pi^+ \pi^- \gamma$
<400	90		LINDSEY	65	HBC	2.1–2.7 $K^- p \rightarrow \Lambda \pi^+ \pi^-$ neutrals

<sup>1</sup> For  $E_\gamma > 20$  MeV and assuming that  $B(\phi(1020) \rightarrow f_0(980)\gamma)$  is negligible. Supersedes AKHMETSHIN 97C.

<sup>2</sup> For  $E_\gamma > 20$  MeV and assuming that  $B(\phi(1020) \rightarrow f_0(980)\gamma)$  is negligible.

$\Gamma(f_0(980)\gamma)/\Gamma_{\text{total}}$						$\Gamma_{17}/\Gamma$
VALUE (units $10^{-4}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT	

<b>3.22 ± 0.19 OUR FIT</b>	Error includes scale factor of 1.1.					
<b>3.21 ± 0.19 OUR AVERAGE</b>						
$3.21^{+0.03}_{-0.09} \pm 0.18$			<sup>1</sup> AMBROSINO 07	KLOE	$e^+ e^- \rightarrow \pi^0 \pi^0 \gamma$	
$2.90 \pm 0.21 \pm 1.54$			<sup>2</sup> AKHMETSHIN 99C	CMD2	$e^+ e^- \rightarrow \pi^+ \pi^- \gamma, \pi^0 \pi^0 \gamma$	
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●						
$4.47 \pm 0.21$	2438		<sup>3</sup> ALOISIO 02D	KLOE	$e^+ e^- \rightarrow \pi^0 \pi^0 \gamma$	
$3.5 \pm 0.3 \pm 1.3_{-0.5}$	419		<sup>4,5</sup> ACHASOV 00H	SND	$e^+ e^- \rightarrow \pi^0 \pi^0 \gamma$	
$1.93 \pm 0.46 \pm 0.50$	27188		<sup>6</sup> AKHMETSHIN 99B	CMD2	$e^+ e^- \rightarrow \pi^+ \pi^- \gamma$	
$3.05 \pm 0.25 \pm 0.72$	268		<sup>7</sup> AKHMETSHIN 99C	CMD2	$e^+ e^- \rightarrow \pi^0 \pi^0 \gamma$	

1.5 ± 0.5	268	8	AKHMETSHIN 99C	CMD2	$e^+e^- \rightarrow \pi^0\pi^0\gamma$
3.42 ± 0.30 ± 0.36	164	4	ACHASOV 98I	SND	$e^+e^- \rightarrow 5\gamma$
< 1	90	9	AKHMETSHIN 97C	CMD2	$e^+e^- \rightarrow \pi^+\pi^-\gamma$
< 7	90	10	AKHMETSHIN 97C	CMD2	$e^+e^- \rightarrow \pi^+\pi^-\gamma$
< 20	90		DRUZHININ 87	ND	$e^+e^- \rightarrow \pi^0\pi^0\gamma$

<sup>1</sup> Obtained by the authors taking into account the  $\pi^+\pi^-$  decay mode. Includes a component due to  $\pi\pi$  production via the  $f_0(500)$  meson. Supersedes ALOISIO 02D.

<sup>2</sup> From the combined fit of the photon spectra in the reactions  $e^+e^- \rightarrow \pi^+\pi^-\gamma$ ,  $\pi^0\pi^0\gamma$ .

<sup>3</sup> From the negative interference with the  $f_0(500)$  meson of AITALA 01B using the ACHASOV 89 parameterization for the  $f_0(980)$ , a Breit-Wigner for the  $f_0(500)$ , and ACHASOV 01F for the  $\rho\pi$  contribution. Superseded by AMBROSINO 07.

<sup>4</sup> Assuming that the  $\pi^0\pi^0\gamma$  final state is completely determined by the  $f_0\gamma$  mechanism, neglecting the decay  $B(\phi \rightarrow K\bar{K}\gamma)$  and using  $B(f_0 \rightarrow \pi^+\pi^-) = 2B(f_0 \rightarrow \pi^0\pi^0)$ .

<sup>5</sup> Using the value  $B(\phi \rightarrow \eta\gamma) = (1.338 \pm 0.053) \times 10^{-2}$ .

<sup>6</sup> For  $E_\gamma > 20$  MeV. Supersedes AKHMETSHIN 97C.

<sup>7</sup> Neglecting other intermediate mechanisms ( $\rho\pi$ ,  $\sigma\gamma$ ).

<sup>8</sup> A narrow pole fit taking into account  $f_0(980)$  and  $f_0(1200)$  intermediate mechanisms.

<sup>9</sup> For destructive interference with the Bremsstrahlung process

<sup>10</sup> For constructive interference with the Bremsstrahlung process

### $\Gamma(f_0(980)\gamma)/\Gamma(\eta\gamma)$ $\Gamma_{17}/\Gamma_6$

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2.48 ± 0.15 OUR FIT</b>	Error includes scale factor of 1.1.			

<b>2.6 ± 0.2 <sup>+0.8</sup>/<sub>-0.3</sub></b>	419	<sup>1</sup> ACHASOV 00H	SND	$e^+e^- \rightarrow \pi^0\pi^0\gamma$
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<sup>1</sup> Assuming that the  $\pi^0\pi^0\gamma$  final state is completely determined by the  $f_0\gamma$  mechanism, neglecting the decay  $B(\phi \rightarrow K\bar{K}\gamma)$  and using  $B(f_0 \rightarrow \pi^+\pi^-) = 2B(f_0 \rightarrow \pi^0\pi^0)$ .

### $\Gamma(\pi^0\pi^0\gamma)/\Gamma_{total}$ $\Gamma_{18}/\Gamma$

VALUE (units $10^{-4}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.07 ± 0.06 OUR AVERAGE</b>					

1.07 <sup>+0.01</sup> / <sub>-0.03</sub> <sup>+0.06</sup> / <sub>-0.06</sub>		<sup>1</sup> AMBROSINO 07	KLOE	$e^+e^- \rightarrow \pi^0\pi^0\gamma$
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1.08 ± 0.17 ± 0.09	268	AKHMETSHIN 99C	CMD2	$e^+e^- \rightarrow \pi^0\pi^0\gamma$
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• • • We do not use the following data for averages, fits, limits, etc. • • •

1.09 ± 0.03 ± 0.05	2438	ALOISIO 02D	KLOE	$e^+e^- \rightarrow \pi^0\pi^0\gamma$
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1.158 ± 0.093 ± 0.052	419	<sup>2,3</sup> ACHASOV 00H	SND	$e^+e^- \rightarrow \pi^0\pi^0\gamma$
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< 10	90	DRUZHININ 87	ND	$e^+e^- \rightarrow 5\gamma$
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<sup>1</sup> Supersedes ALOISIO 02D.

<sup>2</sup> Using the value  $B(\phi \rightarrow \eta\gamma) = (1.338 \pm 0.053) \times 10^{-2}$ .

<sup>3</sup> Supersedes ACHASOV 98I. Excluding  $\omega\pi^0$ .

### $\Gamma(\pi^0\pi^0\gamma)/\Gamma(\eta\gamma)$ $\Gamma_{18}/\Gamma_6$

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.86 ± 0.04 OUR FIT</b>				

<b>0.865 ± 0.070 ± 0.017</b>	419	<sup>1</sup> ACHASOV 00H	SND	$e^+e^- \rightarrow \pi^0\pi^0\gamma$
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• • • We do not use the following data for averages, fits, limits, etc. • • •

0.90 ± 0.08 ± 0.07	164	ACHASOV 98I	SND	$e^+e^- \rightarrow 5\gamma$
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<sup>1</sup> Supersedes ACHASOV 98I. Excluding  $\omega\pi^0$ .

$\Gamma(\pi^+\pi^-\pi^+\pi^-)/\Gamma_{\text{total}}$   $\Gamma_{19}/\Gamma$

VALUE (units $10^{-6}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>6.5 ± 2.7 ± 1.6</b>		6.8k	<sup>1</sup> AKHMETSHIN 17	CMD3	$e^+e^- \rightarrow \pi^+\pi^-\pi^+\pi^-$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
3.93 ± 1.74 ± 2.14		3.3k	AKHMETSHIN 00E	CMD2	$e^+e^- \rightarrow \pi^+\pi^-\pi^+\pi^-$
< 870	90		CORDIER 79	WIRE	$e^+e^- \rightarrow \pi^+\pi^-\pi^+\pi^-$

<sup>1</sup> Using the cross section at the  $\phi$  meson peak  $\sigma(\phi) = 4172 \pm 42$  nb, the nonresonant cross section  $\sigma(0) = 1.263 \pm 0.027$  nb and  $\text{Re}(Z) = 0.146 \pm 0.030$ ,  $\text{Im}(Z) = -0.002 \pm 0.024$  for the complex amplitude of the  $\phi \rightarrow \pi^+\pi^-\pi^+\pi^-$  transition.

$\Gamma(\pi^+\pi^+\pi^-\pi^-\pi^0)/\Gamma_{\text{total}}$   $\Gamma_{20}/\Gamma$

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt; 4.6</b>	90	AKHMETSHIN 00E	CMD2	$e^+e^- \rightarrow \pi^+\pi^-\pi^+\pi^-\pi^0$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
< 150	95	BARKOV 88	CMD	$e^+e^- \rightarrow \pi^+\pi^-\pi^+\pi^-\pi^0$

$\Gamma(\pi^0 e^+ e^-)/\Gamma_{\text{total}}$   $\Gamma_{21}/\Gamma$

VALUE (units $10^{-5}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.33<sup>+0.07</sup><sub>-0.10</sub> OUR AVERAGE</b>					

1.35 ± 0.05<sup>+0.05</sup><sub>-0.10</sub> 9.5k <sup>1</sup> ANASTASI 16B KLOE  $e^+e^- \rightarrow \pi^0 e^+ e^-$

1.01 ± 0.28 ± 0.29 52 <sup>2</sup> ACHASOV 02D SND  $e^+e^- \rightarrow \pi^0 e^+ e^-$

1.22 ± 0.34 ± 0.21 46 <sup>3</sup> AKHMETSHIN 01C CMD2  $e^+e^- \rightarrow \pi^0 e^+ e^-$

• • • We do not use the following data for averages, fits, limits, etc. • • •

< 12 90 DOLINSKY 88 ND  $e^+e^- \rightarrow \pi^0 e^+ e^-$

<sup>1</sup> Using  $B(\pi^0 \rightarrow \gamma\gamma)$  from the 2014 Edition of this Review (PDG 14).

<sup>2</sup> Using various branching ratios from the 2000 Edition of this Review (PDG 00).

<sup>3</sup> Using  $B(\pi^0 \rightarrow \gamma\gamma) = 0.98798 \pm 0.00032$ ,  $B(\phi \rightarrow \eta\gamma) = (1.297 \pm 0.033) \times 10^{-2}$ , and  $B(\eta \rightarrow \pi^+\pi^-\gamma) = (4.75 \pm 0.11) \times 10^{-2}$ .

$\Gamma(\pi^0 \eta \gamma)/\Gamma_{\text{total}}$   $\Gamma_{22}/\Gamma$

VALUE (units $10^{-5}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>7.27 ± 0.30 OUR AVERAGE</b>					Error includes scale factor of 1.5. See the ideogram below.

7.06 ± 0.22 16.9k <sup>1</sup> AMBROSINO 09F KLOE  $1.02 e^+e^- \rightarrow \eta\pi^0\gamma$

8.51 ± 0.51 ± 0.57 607 <sup>2</sup> ALOISIO 02C KLOE  $e^+e^- \rightarrow \eta\pi^0\gamma$

7.96 ± 0.60 ± 0.40 197 <sup>3</sup> ALOISIO 02C KLOE  $e^+e^- \rightarrow \eta\pi^0\gamma$

8.8 ± 1.4 ± 0.9 36 <sup>4</sup> ACHASOV 00F SND  $e^+e^- \rightarrow \eta\pi^0\gamma$

9.0 ± 2.4 ± 1.0 80 AKHMETSHIN 99C CMD2  $e^+e^- \rightarrow \eta\pi^0\gamma$

• • • We do not use the following data for averages, fits, limits, etc. • • •

7.01 ± 0.10 ± 0.20 13.3k <sup>2,5</sup> AMBROSINO 09F KLOE  $1.02 e^+e^- \rightarrow \eta\pi^0\gamma$

7.12 ± 0.13 ± 0.22 3.6k <sup>3,6</sup> AMBROSINO 09F KLOE  $1.02 e^+e^- \rightarrow \eta\pi^0\gamma$

8.3 ± 2.3 ± 1.2 20 ACHASOV 98B SND  $e^+e^- \rightarrow 5\gamma$

< 250 90 DOLINSKY 91 ND  $e^+e^- \rightarrow \pi^0\eta\gamma$

<sup>1</sup> Combined results of  $\eta \rightarrow \gamma\gamma$  and  $\eta \rightarrow \pi^+\pi^-\pi^0$  decay modes measurements.

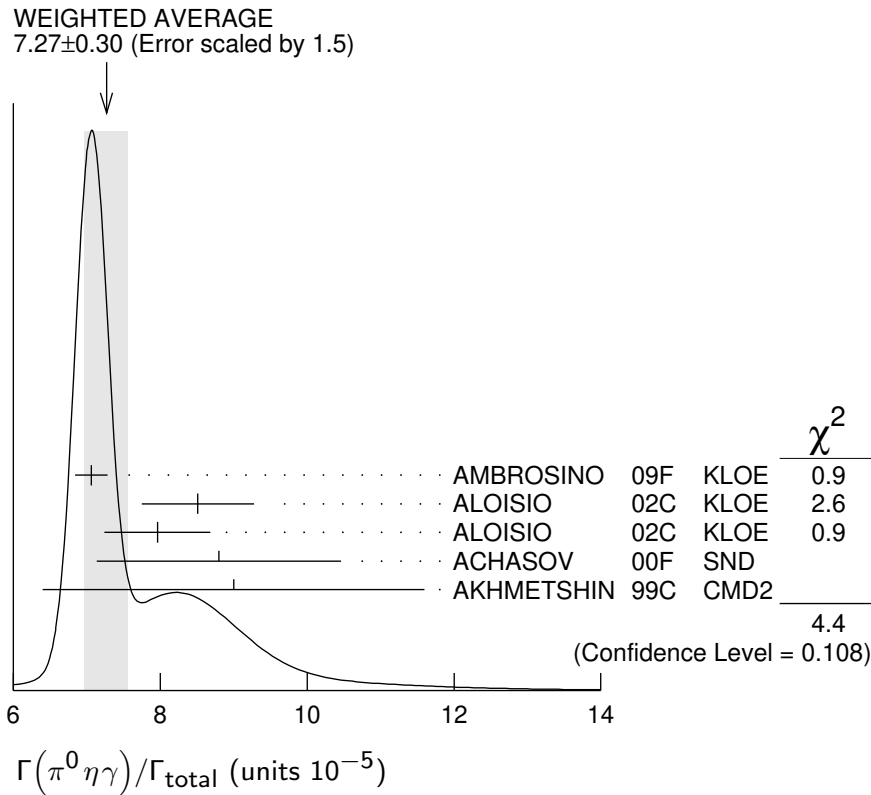
<sup>2</sup> From the decay mode  $\eta \rightarrow \gamma\gamma$ .

<sup>3</sup> From the decay mode  $\eta \rightarrow \pi^+\pi^-\pi^0$ .

<sup>4</sup> Supersedes ACHASOV 98B.

<sup>5</sup> Using  $B(\phi \rightarrow \eta\gamma) = (1.304 \pm 0.025)\%$ ,  $B(\eta \rightarrow 3\pi^0) = (32.56 \pm 0.23)\%$ , and  $B(\eta \rightarrow \gamma\gamma) = (39.31 \pm 0.20)\%$ .

<sup>6</sup> Using  $B(\phi \rightarrow \eta\gamma) = (1.304 \pm 0.025)\%$ ,  $B(\eta \rightarrow 3\pi^0) = (32.56 \pm 0.23)\%$ , and  $B(\eta \rightarrow \pi^+\pi^-\pi^0) = (22.73 \pm 0.28)\%$ .



### $\Gamma(a_0(980)\gamma)/\Gamma_{\text{total}}$

$\Gamma_{23}/\Gamma$

VALUE (units $10^{-5}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>7.6±0.6 OUR FIT</b>					
<b>7.6±0.6 OUR AVERAGE</b>					
7.4±0.7			1 ALOISIO	02C	KLOE $e^+e^- \rightarrow \eta\pi^0\gamma$
8.8±1.7		36	2 ACHASOV	00F	SND $e^+e^- \rightarrow \eta\pi^0\gamma$
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●					
11 ±2			3 GOKALP	02	RVUE $e^+e^- \rightarrow \eta\pi^0\gamma$
<500	90		DOLINSKY	91	ND $e^+e^- \rightarrow \pi^0\eta\gamma$

<sup>1</sup> Using  $M_{a_0(980)}=984.8$  MeV and assuming  $a_0(980)\gamma$  dominance.

<sup>2</sup> Assuming  $a_0(980)\gamma$  dominance in the  $\eta\pi^0\gamma$  final state.

<sup>3</sup> Using data of ACHASOV 00F.

### $\Gamma(f_0(980)\gamma)/\Gamma(a_0(980)\gamma)$

$\Gamma_{17}/\Gamma_{23}$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>6.1±0.6</b>	1 ALOISIO	02C	KLOE $e^+e^- \rightarrow \eta\pi^0\gamma$

<sup>1</sup> Using results of ALOISIO 02D and assuming that  $f_0(980)$  decays into  $\pi\pi$  only and  $a_0(980)$  into  $\eta\pi$  only.

$\Gamma(K^0 \bar{K}^0 \gamma) / \Gamma_{\text{total}}$			$\Gamma_{24} / \Gamma$		
<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	
$<1.9 \times 10^{-8}$	90	AMBROSINO 09C	KLOE	$e^+ e^- \rightarrow K_S^0 K_S^0 \gamma$	

$\Gamma(\eta'(958)\gamma) / \Gamma_{\text{total}}$			$\Gamma_{25} / \Gamma$		
<u>VALUE (units <math>10^{-5}</math>)</u>	<u>CL%</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>

**6.21 ± 0.21 OUR FIT**  
**6.21 ± 0.30 OUR AVERAGE**

6.21 ± 0.27 ± 0.12	3407	<sup>1</sup> AMBROSINO 07A	KLOE	1.02 $e^+ e^- \rightarrow \pi^+ \pi^- 7\gamma$	
6.7 $^{+2.8}_{-2.4} \pm 0.8$	12	<sup>2</sup> AULCHENKO 03B	SND	$e^+ e^- \rightarrow \eta' \gamma$	
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●					
6.7 $^{+5.0}_{-4.2} \pm 1.5$	7	AULCHENKO 03B	SND	$e^+ e^- \rightarrow 7\gamma$	
6.10 ± 0.61 ± 0.43	120	<sup>3</sup> ALOISIO 02E	KLOE	1.02 $e^+ e^- \rightarrow \pi^+ \pi^- 3\gamma$	
8.2 $^{+2.1}_{-1.9} \pm 1.1$	21	<sup>4</sup> AKHMETSHIN 00B	CMD2	$e^+ e^- \rightarrow \pi^+ \pi^- 3\gamma$	
4.9 $^{+2.2}_{-1.8} \pm 0.6$	9	<sup>5</sup> AKHMETSHIN 00F	CMD2	$e^+ e^- \rightarrow \pi^+ \pi^- \pi^+ \pi^- \geq 2\gamma$	
6.4 ± 1.6	30	<sup>6</sup> AKHMETSHIN 00F	CMD2	$e^+ e^- \rightarrow \eta'(958)\gamma$	
6.7 $^{+3.4}_{-2.9} \pm 1.0$	5	<sup>7</sup> AULCHENKO 99	SND	$e^+ e^- \rightarrow \pi^+ \pi^- 3\gamma$	
<11	90	AULCHENKO 98	SND	$e^+ e^- \rightarrow 7\gamma$	
12 $^{+7}_{-5} \pm 2$	6	<sup>4</sup> AKHMETSHIN 97B	CMD2	$e^+ e^- \rightarrow \pi^+ \pi^- 3\gamma$	
<41	90	DRUZHININ 87	ND	$e^+ e^- \rightarrow \gamma \eta \pi^+ \pi^-$	

<sup>1</sup> AMBROSINO 07A reports  $[\Gamma(\phi(1020) \rightarrow \eta'(958)\gamma) / \Gamma_{\text{total}}] / [B(\phi(1020) \rightarrow \eta\gamma)] = (4.77 \pm 0.09 \pm 0.19) \times 10^{-3}$  which we multiply by our best value  $B(\phi(1020) \rightarrow \eta\gamma) = (1.301 \pm 0.025) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>2</sup> Averaging AULCHENKO 03B with AULCHENKO 99.

<sup>3</sup> Using  $B(\phi \rightarrow \eta\gamma) = (1.297 \pm 0.033)\%$ .

<sup>4</sup> Using the value  $B(\phi \rightarrow \eta\gamma) = (1.26 \pm 0.06) \times 10^{-2}$ .

<sup>5</sup> Using  $B(\phi \rightarrow K_L^0 K_S^0) = (33.8 \pm 0.6)\%$ .

<sup>6</sup> Averaging AKHMETSHIN 00B with AKHMETSHIN 00F.

<sup>7</sup> Using the value  $B(\eta' \rightarrow \eta \pi^+ \pi^-) = (43.7 \pm 1.5) \times 10^{-2}$  and  $B(\eta \rightarrow \gamma\gamma) = (39.25 \pm 0.31) \times 10^{-2}$ .

$\Gamma(\eta'(958)\gamma) / \Gamma(K_L^0 K_S^0)$			$\Gamma_{25} / \Gamma_2$		
<u>VALUE (units <math>10^{-4}</math>)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	

**1.83 ± 0.06 OUR FIT**

1.46 $^{+0.64}_{-0.54} \pm 0.18$	9	<sup>1</sup> AKHMETSHIN 00F	CMD2	$e^+ e^- \rightarrow \pi^+ \pi^- \pi^+ \pi^- \geq 2\gamma$	
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<sup>1</sup> Using various branching ratios of  $K_S^0$ ,  $K_L^0$ ,  $\eta$ ,  $\eta'$  from the 2000 edition (The European Physical Journal **C15** 1 (2000)) of this Review.

### $\Gamma(\eta'(958)\gamma)/\Gamma(\eta\gamma)$

$\Gamma_{25}/\Gamma_6$

<u>VALUE (units <math>10^{-3}</math>)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>4.77±0.15 OUR FIT</b>				
<b>4.78±0.20 OUR AVERAGE</b>				
4.77±0.09±0.19	3407	AMBROSINO 07A	KLOE	1.02 $e^+e^- \rightarrow \pi^+\pi^-7\gamma$
4.70±0.47±0.31	120	<sup>1</sup> ALOISIO 02E	KLOE	1.02 $e^+e^- \rightarrow \pi^+\pi^-3\gamma$
6.5 $\begin{smallmatrix} +1.7 \\ -1.5 \end{smallmatrix} \pm 0.8$	21	AKHMETSHIN 00B	CMD2	$e^+e^- \rightarrow \pi^+\pi^-3\gamma$
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
9.5 $\begin{smallmatrix} +5.2 \\ -4.0 \end{smallmatrix} \pm 1.4$	6	<sup>2</sup> AKHMETSHIN 97B	CMD2	$e^+e^- \rightarrow \pi^+\pi^-3\gamma$

<sup>1</sup>From the decay mode  $\eta' \rightarrow \eta\pi^+\pi^-$ ,  $\eta \rightarrow \gamma\gamma$ .

<sup>2</sup>Superseded by AKHMETSHIN 00B.

### $\Gamma(\eta\pi^0\pi^0\gamma)/\Gamma_{\text{total}}$

$\Gamma_{26}/\Gamma$

<u>VALUE (units <math>10^{-5}</math>)</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>&lt;2</b>	90	AULCHENKO 98	SND	$e^+e^- \rightarrow 7\gamma$

### $\Gamma(\mu^+\mu^-\gamma)/\Gamma_{\text{total}}$

$\Gamma_{27}/\Gamma$

<u>VALUE (units <math>10^{-5}</math>)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>1.43±0.45±0.14</b>	27188	<sup>1</sup> AKHMETSHIN 99B	CMD2	$e^+e^- \rightarrow \mu^+\mu^-\gamma$
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
2.3 ±1.0	824 ± 33	<sup>2</sup> AKHMETSHIN 97C	CMD2	$e^+e^- \rightarrow \mu^+\mu^-\gamma$
<sup>1</sup> For $E_\gamma > 20$ MeV. Supersedes AKHMETSHIN 97C.				
<sup>2</sup> For $E_\gamma > 20$ MeV.				

### $\Gamma(\rho\gamma\gamma)/\Gamma_{\text{total}}$

$\Gamma_{28}/\Gamma$

<u>VALUE (units <math>10^{-4}</math>)</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>&lt;1.2</b>	90	AULCHENKO 08	CMD2	$\phi \rightarrow \pi^+\pi^-\gamma\gamma$
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
<5	90	AKHMETSHIN 98	CMD2	$e^+e^- \rightarrow \pi^+\pi^-\gamma\gamma$

### $\Gamma(\eta\pi^+\pi^-)/\Gamma_{\text{total}}$

$\Gamma_{29}/\Gamma$

<u>VALUE (units <math>10^{-5}</math>)</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>&lt; 1.8</b>	90	AKHMETSHIN 00E	CMD2	$e^+e^- \rightarrow \pi^+\pi^-\pi^+\pi^-\pi^0$
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
< 6.1	90	AULCHENKO 08	CMD2	$\phi \rightarrow \eta\pi^+\pi^-$
<30	90	AKHMETSHIN 98	CMD2	$e^+e^- \rightarrow \pi^+\pi^-\gamma\gamma$

### $\Gamma(\eta\mu^+\mu^-)/\Gamma_{\text{total}}$

$\Gamma_{30}/\Gamma$

<u>VALUE (units <math>10^{-6}</math>)</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>&lt;9.4</b>	90	AKHMETSHIN 01	CMD2	$e^+e^- \rightarrow \eta e^+e^-$

### $\Gamma(\eta U \rightarrow \eta e^+e^-)/\Gamma_{\text{total}}$

$\Gamma_{31}/\Gamma$

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>&lt;1 × 10<sup>-6</sup></b>	90	<sup>1</sup> BABUSCI 13B	KLOE	1.02 $e^+e^- \rightarrow \eta e^+e^-$

<sup>1</sup>For a narrow vector  $U$  with mass between 5 and 470 MeV, from the combined analysis of  $\eta \rightarrow \pi^+\pi^-\pi^0$  and  $\eta \rightarrow \pi^0\pi^0\pi^0$  from ARCHILLI 12. Measured 90% CL limits as a function of  $m_U$  range from  $2.2 \times 10^{-8}$  to  $10^{-6}$ .

$\Gamma(\text{invisible})/\Gamma(K^+K^-)$   $\Gamma_{32}/\Gamma_1$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<3.4 \times 10^{-4}$	90	ABLIKIM	18S BES3	$J/\psi \rightarrow \phi \eta \rightarrow \phi \pi^+ \pi^- \pi^0$

———— Lepton Family number (LF) violating modes ————

$\Gamma(e^\pm \mu^\mp)/\Gamma_{\text{total}}$   $\Gamma_{33}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<2 \times 10^{-6}$	90	ACHASOV	10A SND	$e^+ e^- \rightarrow e^\pm \mu^\mp$

$\pi^+ \pi^- \pi^0 / \rho\pi$  AMPLITUDE RATIO  $a_1$  IN DECAY OF  $\phi \rightarrow \pi^+ \pi^- \pi^0$

NIECKNIG 12 describes final-state interactions between the three pions in a dispersive framework using data on the  $\pi\pi$   $P$ -wave scattering phase shift.

VALUE (units $10^{-2}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>9.1±1.2 OUR AVERAGE</b>					
$10.1 \pm 4.4 \pm 1.7$		80k	<sup>1</sup> AKHMETSHIN 06	CMD2	$1.017-1.021 e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$
$9.0 \pm 1.1 \pm 0.6$		1.98M	<sup>2,3</sup> ALOISIO	03 KLOE	$1.02 e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$-6 < a_1 < 6$		500k	<sup>3</sup> ACHASOV	02 SND	$e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$
$-16 < a_1 < 11$	90	9.8k	<sup>1,4</sup> AKHMETSHIN 98	CMD2	$e^+ e^- \rightarrow \pi^+ \pi^- \gamma \gamma$

<sup>1</sup> Dalitz plot analysis taking into account interference between the contact and  $\rho\pi$  amplitudes.

<sup>2</sup> From a fit without limitations on charged and neutral  $\rho$  masses and widths.

<sup>3</sup> Recalculated by us to match the notations of AKHMETSHIN 98.

<sup>4</sup> Assuming zero phase for the contact term.

**PARAMETER  $\beta$  IN  $\phi \rightarrow P e^+ e^-$  DECAYS**

In the one-pole approximation the electromagnetic transition form factor for  $\phi \rightarrow P e^+ e^-$  ( $P = \pi, \eta$ ) is given as a function of the  $e^+ e^-$  invariant mass squared,  $q^2$ , by the expression:

$$|F(q^2)|^2 = (1 - q^2/\Lambda^2)^{-2},$$

where vector meson dominance predicts parameter  $\Lambda \approx 0.770$  GeV ( $\Lambda^{-2} \approx 1.687$  GeV<sup>-2</sup>). The slope of this form factor,  $\beta = dF/dq^2(q^2=0)$ , equals  $\Lambda^{-2}$  in this approximation.

The measurements below obtain  $\beta$  in the one-pole approximation.

**PARAMETER  $\beta$  IN  $\phi \rightarrow \pi^0 e^+ e^-$  DECAY**

VALUE (GeV <sup>-2</sup> )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2.02±0.11</b>	9.5k	<sup>1</sup> ANASTASI	16B KLOE	$1.02 e^+ e^- \rightarrow \pi^0 e^+ e^-$

<sup>1</sup> The error combines statistical and systematic uncertainties.

## PARAMETER $\beta$ IN $\phi \rightarrow \eta e^+ e^-$ DECAY

<u>VALUE (GeV<sup>-2</sup>)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>1.29 ± 0.13 OUR AVERAGE</b>				
1.28 ± 0.10 <sup>+0.09</sup> <sub>-0.08</sub>	30k	BABUSCI	15 KLOE	1.02 e <sup>+</sup> e <sup>-</sup> → ηe <sup>+</sup> e <sup>-</sup>
3.8 ± 1.8	213	<sup>1</sup> ACHASOV	01B SND	1.02 e <sup>+</sup> e <sup>-</sup> → ηe <sup>+</sup> e <sup>-</sup>

<sup>1</sup>The uncertainty is statistical only. The systematic one is negligible, in comparison.

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PDG	12	PR D86 010001	J. Beringer <i>et al.</i>	(PDG Collab.)
AKHMETSHIN	11	PL B695 412	R.R. Akhmetshin <i>et al.</i>	(CMD-2 Collab.)
ACHASOV	10A	PR D81 057102	M.N. Achasov <i>et al.</i>	(Novosibirsk SND Collab.)
BENAYOUN	10	EPJ C65 211	M. Benayoun <i>et al.</i>	
AMBROSINO	09C	PL B679 10	F. Ambrosino <i>et al.</i>	(KLOE Collab.)
AMBROSINO	09F	PL B681 5	F. Ambrosino <i>et al.</i>	(KLOE Collab.)
AKHMETSHIN	08	PL B669 217	R.R. Akhmetshin <i>et al.</i>	(CMD-2 Collab.)
AMBROSINO	08G	PL B669 223	F. Ambrosino <i>et al.</i>	(KLOE Collab.)
AULCHENKO	08	JETPL 88 85	V. Aulchenko <i>et al.</i>	(CMD-2 Collab.)
		Translated from ZETFP 88 93.		
FLOREZ-BAEZ	08	PR D78 077301	F.V. Florez-Baez, G. Lopez Castro	
ACHASOV	07B	PR D76 077101	M.N. Achasov <i>et al.</i>	(SND Collab.)
AMBROSINO	07	EPJ C49 473	F. Ambrosino <i>et al.</i>	(KLOE Collab.)
AMBROSINO	07A	PL B648 267	F. Ambrosino <i>et al.</i>	(KLOE Collab.)
DUBYNSKIY	07	PR D75 113001	S. Dubynskiy <i>et al.</i>	
ACHASOV	06A	PR D74 014016	M.N. Achasov <i>et al.</i>	(SND Collab.)
AKHMETSHIN	06	PL B642 203	R.R. Akhmetshin <i>et al.</i>	(CMD-2 Collab.)
AKHMETSHIN	05	PL B605 26	R.R. Akhmetshin <i>et al.</i>	(Novosibirsk CMD-2 Collab.)
AMBROSINO	05	PL B608 199	F. Ambrosino <i>et al.</i>	(KLOE Collab.)
AUBERT,B	05J	PR D72 052008	B. Aubert <i>et al.</i>	(BABAR Collab.)
AKHMETSHIN	04	PL B578 285	R.R. Akhmetshin <i>et al.</i>	(Novosibirsk CMD-2 Collab.)
AUBERT,B	04N	PR D70 072004	B. Aubert <i>et al.</i>	(BABAR Collab.)
ALOISIO	03	PL B561 55	A. Aloisio <i>et al.</i>	(KLOE Collab.)
AULCHENKO	03B	JETP 97 24	V.M. Aulchenko <i>et al.</i>	(Novosibirsk SND Collab.)
		Translated from ZETF 124 28.		
ACHASOV	02	PR D65 032002	M.N. Achasov <i>et al.</i>	(Novosibirsk SND Collab.)
ACHASOV	02D	JETPL 75 449	M.N. Achasov <i>et al.</i>	(Novosibirsk SND Collab.)
		Translated from ZETFP 75 539.		
ALOISIO	02C	PL B536 209	A. Aloisio <i>et al.</i>	(KLOE Collab.)
ALOISIO	02D	PL B537 21	A. Aloisio <i>et al.</i>	(KLOE Collab.)
ALOISIO	02E	PL B541 45	A. Aloisio <i>et al.</i>	(KLOE Collab.)
FISCHBACH	02	PL B526 355	E. Fischbach, A.W. Overhauser, B. Woodahl	
GOKALP	02	JP G28 2783	A. Gokalp <i>et al.</i>	
ACHASOV	01B	PL B504 275	M.N. Achasov <i>et al.</i>	(Novosibirsk SND Collab.)



ACHASOV	01E	PR D63 072002	M.N. Achasov <i>et al.</i>	(Novosibirsk SND Collab.)
ACHASOV	01F	PR D63 094007	N.N. Achasov, V.V. Gubin	(Novosibirsk SND Collab.)
ACHASOV	01G	PRL 86 1698	M.N. Achasov <i>et al.</i>	(Novosibirsk SND Collab.)
AITALA	01B	PRL 86 770	E.M. Aitala <i>et al.</i>	(FNAL E791 Collab.)
AKHMETSHIN	01	PL B501 191	R.R. Akhmetshin <i>et al.</i>	(Novosibirsk CMD-2 Collab.)
AKHMETSHIN	01B	PL B509 217	R.R. Akhmetshin <i>et al.</i>	(Novosibirsk CMD-2 Collab.)
AKHMETSHIN	01C	PL B503 237	R.R. Akhmetshin <i>et al.</i>	(Novosibirsk CMD-2 Collab.)
BENAYOUN	01	EPJ C22 503	M. Benayoun, H.B. O'Connell	
ACHASOV	00	EPJ C12 25	M.N. Achasov <i>et al.</i>	(Novosibirsk SND Collab.)
ACHASOV	00B	JETP 90 17	M.N. Achasov <i>et al.</i>	(Novosibirsk SND Collab.)
		Translated from ZETF 117 22.		
ACHASOV	00C	PL B474 188	M.N. Achasov <i>et al.</i>	(Novosibirsk SND Collab.)
ACHASOV	00D	JETPL 72 282	M.N. Achasov <i>et al.</i>	(Novosibirsk SND Collab.)
		Translated from ZETFP 72 411.		
ACHASOV	00E	NP B569 158	M.N. Achasov <i>et al.</i>	(Novosibirsk SND Collab.)
ACHASOV	00F	PL B479 53	M.N. Achasov <i>et al.</i>	(Novosibirsk SND Collab.)
ACHASOV	00H	PL B485 349	M.N. Achasov <i>et al.</i>	(Novosibirsk SND Collab.)
AKHMETSHIN	00B	PL B473 337	R.R. Akhmetshin <i>et al.</i>	(Novosibirsk CMD-2 Collab.)
AKHMETSHIN	00E	PL B491 81	R.R. Akhmetshin <i>et al.</i>	(Novosibirsk CMD-2 Collab.)
AKHMETSHIN	00F	PL B494 26	R.R. Akhmetshin <i>et al.</i>	(Novosibirsk CMD-2 Collab.)
AULCHENKO	00A	JETP 90 927	V.M. Aulchenko <i>et al.</i>	(Novosibirsk SND Collab.)
		Translated from ZETF 117 1067.		
BRAMON	00	PL B486 406	A. Bramon <i>et al.</i>	
PDG	00	EPJ C15 1	D.E. Groom <i>et al.</i>	(PDG Collab.)
ACHASOV	99	PL B449 122	M.N. Achasov <i>et al.</i>	
ACHASOV	99C	PL B456 304	M.N. Achasov <i>et al.</i>	
AKHMETSHIN	99B	PL B462 371	R.R. Akhmetshin <i>et al.</i>	(Novosibirsk CMD-2 Collab.)
AKHMETSHIN	99C	PL B462 380	R.R. Akhmetshin <i>et al.</i>	(Novosibirsk CMD-2 Collab.)
AKHMETSHIN	99D	PL B466 385	R.R. Akhmetshin <i>et al.</i>	
	Also	PL B508 217 (errat.)	R.R. Akhmetshin <i>et al.</i>	(Novosibirsk CMD-2 Collab.)
AKHMETSHIN	99F	PL B460 242	R.R. Akhmetshin <i>et al.</i>	(Novosibirsk CMD-2 Collab.)
AULCHENKO	99	JETPL 69 97	V.M. Aulchenko <i>et al.</i>	
		Translated from ZETFP 69 87.		
ACHASOV	98B	PL B438 441	M.N. Achasov <i>et al.</i>	(Novosibirsk SND Collab.)
ACHASOV	98F	JETPL 68 573	M.N. Achasov <i>et al.</i>	(Novosibirsk SND Collab.)
ACHASOV	98I	PL B440 442	M.N. Achasov <i>et al.</i>	
AKHMETSHIN	98	PL B434 426	R.R. Akhmetshin <i>et al.</i>	(CMD-2 Collab.)
AULCHENKO	98	PL B436 199	V.M. Aulchenko <i>et al.</i>	(Novosibirsk SND Collab.)
BARBERIS	98	PL B432 436	D. Barberis <i>et al.</i>	(Omega Expt.)
AKHMETSHIN	97B	PL B415 445	R.R. Akhmetshin <i>et al.</i>	(NOVO, BOST, PITT+)
AKHMETSHIN	97C	PL B415 452	R.R. Akhmetshin <i>et al.</i>	(Novosibirsk CMD-2 Collab.)
BENAYOUN	96	ZPHY C72 221	M. Benayoun <i>et al.</i>	(IPNP, NOVO)
AKHMETSHIN	95	PL B364 199	R.R. Akhmetshin <i>et al.</i>	(Novosibirsk CMD-2 Collab.)
DOLINSKY	91	PRPL 202 99	S.I. Dolinsky <i>et al.</i>	(NOVO)
KUHN	90	ZPHY C48 445	J.H. Kuhn <i>et al.</i>	(MPIM)
ACHASOV	89	NP B315 465	N.N. Achasov, V.N. Ivanchenko	
DOLINSKY	89	ZPHY C42 511	S.I. Dolinsky <i>et al.</i>	(NOVO)
BARKOV	88	SJNP 47 248	L.M. Barkov <i>et al.</i>	(NOVO)
		Translated from YAF 47 393.		
DOLINSKY	88	SJNP 48 277	S.I. Dolinsky <i>et al.</i>	(NOVO)
		Translated from YAF 48 442.		
DRUZHININ	87	ZPHY C37 1	V.P. Druzhinin <i>et al.</i>	(NOVO)
ARMSTRONG	86	PL 166B 245	T.A. Armstrong <i>et al.</i>	(ATHU, BARI, BIRM+)
ATKINSON	86	ZPHY C30 521	M. Atkinson <i>et al.</i>	(BONN, CERN, GLAS+)
BEBEK	86	PRL 56 1893	C. Bebek <i>et al.</i>	(CLEO Collab.)
DAVENPORT	86	PR D33 2519	T.F. Davenport	(TUFTS, ARIZ, FNAL, FSU, NDAM+)
DIJKSTRA	86	ZPHY C31 375	H. Dijkstra <i>et al.</i>	(ANIK, BRIS, CERN+)
FRAME	86	NP B276 667	D. Frame <i>et al.</i>	(GLAS)
GOLUBEV	86	SJNP 44 409	V.B. Golubev <i>et al.</i>	(NOVO)
		Translated from YAF 44 633.		
ALBRECHT	85D	PL 153B 343	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
GOLUBEV	85	SJNP 41 756	V.B. Golubev <i>et al.</i>	(NOVO)
		Translated from YAF 41 1183.		
DRUZHININ	84	PL 144B 136	V.P. Druzhinin <i>et al.</i>	(NOVO)
ARMSTRONG	83B	NP B224 193	T.A. Armstrong <i>et al.</i>	(BARI, BIRM, CERN+)
BARATE	83	PL 121B 449	R. Barate <i>et al.</i>	(SACL, LOIC, SHMP, IND)
KURDADZE	83C	JETPL 38 366	L.M. Kurdadze <i>et al.</i>	(NOVO)
		Translated from ZETFP 38 306.		

ARENTON	82	PR D25 2241	M.W. Arenton <i>et al.</i>	(ANL, ILL)
PELLINEN	82	PS 25 599	A. Pellinen, M. Roos	(HELS)
DAUM	81	PL 100B 439	C. Daum <i>et al.</i>	(AMST, BRIS, CERN, CRAC+)
IVANOV	81	PL 107B 297	P.M. Ivanov <i>et al.</i>	(NOVO)
Also		Private Comm.	S.I. Eidelman	(NOVO)
VASSERMAN	81	PL 99B 62	I.B. Vasserman <i>et al.</i>	(NOVO)
Also		SJNP 35 240	L.M. Kurdadze <i>et al.</i>	
		Translated from YAF 35	352.	
CORDIER	80	NP B172 13	A. Cordier <i>et al.</i>	(LALO)
CORDIER	79	PL 81B 389	A. Cordier <i>et al.</i>	(LALO)
BUKIN	78B	SJNP 27 521	A.D. Bukin <i>et al.</i>	(NOVO)
		Translated from YAF 27	985.	
BUKIN	78C	SJNP 27 516	A.D. Bukin <i>et al.</i>	(NOVO)
		Translated from YAF 27	976.	
COOPER	78B	NP B146 1	A.M. Cooper <i>et al.</i>	(TATA, CERN, CDEF+)
LOSTY	78	NP B133 38	M.J. Losty <i>et al.</i>	(CERN, AMST, NIJM+)
AKERLOF	77	PRL 39 861	C.W. Akerlof <i>et al.</i>	(FNAL, MICH, PURD)
ANDREWS	77	PRL 38 198	D.E. Andrews <i>et al.</i>	(ROCH)
BALDI	77	PL 68B 381	R. Baldi <i>et al.</i>	(GEVA)
CERRADA	77B	NP B126 241	M. Cerrada <i>et al.</i>	(AMST, CERN, NIJM+)
COHEN	77	PRL 38 269	D. Cohen <i>et al.</i>	(ANL)
LAVEN	77	NP B127 43	H. Laven <i>et al.</i>	(AACH3, BERL, CERN, LOIC+)
LYONS	77	NP B125 207	L. Lyons, A.M. Cooper, A.G. Clark	(OXF)
COSME	76	PL 63B 352	G. Cosme <i>et al.</i>	(ORSAY)
KALBFLEISCH	76	PR D13 22	G.R. Kalbfleisch, R.C. Strand, J.W. Chapman	(BNL+)
PARROUR	76	PL 63B 357	G. Parrour <i>et al.</i>	(ORSAY)
PARROUR	76B	PL 63B 362	G. Parrour <i>et al.</i>	(ORSAY)
KALBFLEISCH	75	PR D11 987	G.R. Kalbfleisch, R.C. Strand, J.W. Chapman	(BNL+)
AYRES	74	PRL 32 1463	D.S. Ayres <i>et al.</i>	(ANL)
BESCH	74	NP B70 257	H.J. Besch <i>et al.</i>	(BONN)
COSME	74	PL 48B 155	G. Cosme <i>et al.</i>	(ORSAY)
COSME	74B	PL 48B 159	G. Cosme <i>et al.</i>	(ORSAY)
DEGROOT	74	NP B74 77	A.J. de Groot <i>et al.</i>	(AMST, NIJM)
AUGUSTIN	73	PRL 30 462	J.E. Augustin <i>et al.</i>	(ORSAY)
BALLAM	73	PR D7 3150	J. Ballam <i>et al.</i>	(SLAC, LBL)
BINNIE	73B	PR D8 2789	D.M. Binnie <i>et al.</i>	(LOIC, SHMP)
AGUILAR-...	72B	PR D6 29	M. Aguilar-Benitez <i>et al.</i>	(BNL)
ALVENSLEB...	72	PRL 28 66	H. Alvensleben <i>et al.</i>	(MIT, DESY)
BORENSTEIN	72	PR D5 1559	S.R. Borenstein <i>et al.</i>	(BNL, MICH)
COLLEY	72	NP B50 1	D.C. Colley <i>et al.</i>	(BIRM, GLAS)
BALAKIN	71	PL 34B 328	V.E. Balakin <i>et al.</i>	(NOVO)
CHATELUS	71	Thesis LAL 1247	Y. Chatelus	(STRB)
Also		PL 32B 416	J.C. Bizot <i>et al.</i>	(ORSAY)
HAYES	71	PR D4 899	S. Hayes <i>et al.</i>	(CORN)
STOTTLE...	71	Thesis ORO 2504 170	A.R. Stottlemyer	(UMD)
BIZOT	70	PL 32B 416	J.C. Bizot <i>et al.</i>	(ORSAY)
Also		Liverpool Sym. 69	J.P. Perez-y-Jorba	
EARLES	70	PRL 25 1312	D.R. Earles <i>et al.</i>	(NEAS)
LINDSEY	66	PR 147 913	J.S. Lindsey, G. Smith	(LRL)
LONDON	66	PR 143 1034	G.W. London <i>et al.</i>	(BNL, SYRA) IGJPC
BADIER	65B	PL 17 337	J. Badier <i>et al.</i>	(EPOL, SACL, AMST)
LINDSEY	65	PRL 15 221	J.S. Lindsey, G.A. Smith	(LRL)
LINDSEY 65 data		included in LINDSEY 66.		
SCHLEIN	63	PRL 10 368	P.E. Schlein <i>et al.</i>	(UCLA) IGJP