



$$J = \frac{1}{2}$$

$\tau$  discovery paper was PERL 75.  $e^+e^- \rightarrow \tau^+\tau^-$  cross-section threshold behavior and magnitude are consistent with pointlike spin-1/2 Dirac particle. BRANDELIK 78 ruled out pointlike spin-0 or spin-1 particle. FELDMAN 78 ruled out  $J = 3/2$ . KIRKBY 79 also ruled out  $J=\text{integer}$ ,  $J = 3/2$ .

### $\tau$ MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1776.93±0.09 OUR AVERAGE</b>				
1777.09±0.08±0.11	175M	<sup>1</sup> ADACHI	23C BELL	190 fb <sup>-1</sup> , $E_{\text{cm}}^{ee} = 10.6$ GeV
1776.69 <sup>+0.17</sup> <sub>-0.19</sub> ±0.15		<sup>2</sup> ANASHIN	23A KEDR	(6.7+ 8.5) pb <sup>-1</sup> , $E_{\text{cm}}^{ee} = 3.54\text{--}3.78$ GeV
1776.91±0.12 <sup>+0.10</sup> <sub>-0.13</sub>	1171	<sup>3</sup> ABLIKIM	14D BES3	23.3 pb <sup>-1</sup> , $E_{\text{cm}}^{ee} = 3.54\text{--}3.60$ GeV
1776.68±0.12±0.41	682k	<sup>1</sup> AUBERT	09AK BABR	423 fb <sup>-1</sup> , $E_{\text{cm}}^{ee} = 10.6$ GeV
1776.61±0.13±0.35		<sup>1</sup> BELOUS	07 BELL	414 fb <sup>-1</sup> $E_{\text{cm}}^{ee} = 10.6$ GeV
1775.1 ±1.6 ±1.0	13.3k	<sup>4</sup> ABBIENDI	00A OPAL	1990–1995 LEP runs
1778.2 ±0.8 ±1.2		ANASTASSOV	97 CLEO	$E_{\text{cm}}^{ee} = 10.6$ GeV
1776.96 <sup>+0.18</sup> <sub>-0.21</sub> <sup>+0.25</sup> <sub>-0.17</sub>	65	<sup>5</sup> BAI	96 BES	$E_{\text{cm}}^{ee} = 3.54\text{--}3.57$ GeV
1776.3 ±2.4 ±1.4	11k	<sup>6</sup> ALBRECHT	92M ARG	$E_{\text{cm}}^{ee} = 9.4\text{--}10.6$ GeV
1783 <sup>+3</sup> <sub>-4</sub>	692	<sup>7</sup> BACINO	78B DLCO	$E_{\text{cm}}^{ee} = 3.1\text{--}7.4$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1776.81 <sup>+0.25</sup> <sub>-0.23</sub> ±0.15	81	ANASHIN	07 KEDR	6.7 pb <sup>-1</sup> , $E_{\text{cm}}^{ee} = 3.54\text{--}3.78$ GeV
1777.8 ±0.7 ±1.7	35k	<sup>8</sup> BALEST	93 CLEO	Repl. by ANASTASSOV 97
1776.9 <sup>+0.4</sup> <sub>-0.5</sub> ±0.2	14	<sup>9</sup> BAI	92 BES	Repl. by BAI 96

<sup>1</sup> ADACHI 23C, AUBERT 09AK and BELOUS 07 fit  $\tau$  pseudomass spectrum in  $\tau \rightarrow \pi\pi^+\pi^-\nu_\tau$  decays. Result assumes  $m_{\nu_\tau} = 0$ .

<sup>2</sup> Previously also reported LEVICHEV 14. Superseeds ANASHIN 07.

<sup>3</sup> ABLIKIM 14D fit  $\sigma(e^+e^- \rightarrow \tau^+\tau^-)$  at different energies near threshold.

<sup>4</sup> ABBIENDI 00A fit  $\tau$  pseudomass spectrum in  $\tau \rightarrow \pi^\pm \leq 2\pi^0\nu_\tau$  and  $\tau \rightarrow \pi^\pm\pi^+\pi^- \leq 1\pi^0\nu_\tau$  decays. Result assumes  $m_{\nu_\tau} = 0$ .

<sup>5</sup> BAI 96 fit  $\sigma(e^+e^- \rightarrow \tau^+\tau^-)$  at different energies near threshold.

<sup>6</sup> ALBRECHT 92M fit  $\tau$  pseudomass spectrum in  $\tau^- \rightarrow 2\pi^-\pi^+\nu_\tau$  decays. Result assumes  $m_{\nu_\tau} = 0$ .

<sup>7</sup> BACINO 78B value comes from  $e^\pm X^\mp$  threshold. Published mass 1782 MeV increased by 1 MeV using the high precision  $\psi(2S)$  mass measurement of ZHOLENTZ 80 to eliminate the absolute SPEAR energy calibration uncertainty.

<sup>8</sup> BALEST 93 fit spectra of minimum kinematically allowed  $\tau$  mass in events of the type  $e^+e^- \rightarrow \tau^+\tau^- \rightarrow (\pi^+n\pi^0\nu_\tau)(\pi^-m\pi^0\nu_\tau)$   $n \leq 2$ ,  $m \leq 2$ ,  $1 \leq n+m \leq 3$ . If  $m_{\nu_\tau} \neq 0$ , result increases by  $(m_{\nu_\tau}^2/1100)$  MeV.

<sup>9</sup>BAI 92 fit  $\sigma(e^+e^- \rightarrow \tau^+\tau^-)$  near threshold using  $e\mu$  events.

$$(m_{\tau^+} - m_{\tau^-})/m_{\text{average}}$$

A test of *CPT* invariance.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;2.8 × 10<sup>-4</sup></b>	90	BELOUS	07 BELL	414 fb <sup>-1</sup> , E <sub>cm</sub> <sup>ee</sup> =10.6 GeV
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
<5.5 × 10 <sup>-4</sup>	90	<sup>1</sup> AUBERT	09AK BABR	423 fb <sup>-1</sup> , E <sub>cm</sub> <sup>ee</sup> =10.6 GeV
<3.0 × 10 <sup>-3</sup>	90	ABBIENDI	00A OPAL	1990–1995 LEP runs
<sup>1</sup> AUBERT 09AK quote both the listed upper limit and $(m_{\tau^+} - m_{\tau^-})/m_{\text{average}} = (-3.4 \pm 1.3 \pm 0.3) \times 10^{-4}$ .				

### $\tau$ MEAN LIFE

VALUE (10 <sup>-15</sup> s)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>290.3 ± 0.5 OUR AVERAGE</b>				
290.17 ± 0.53 ± 0.33	1.1M	BELOUS	14 BELL	711 fb <sup>-1</sup> E <sub>cm</sub> <sup>ee</sup> =10.6 GeV
290.9 ± 1.4 ± 1.0		ABDALLAH	04T DLPH	1991-1995 LEP runs
293.2 ± 2.0 ± 1.5		ACCIARRI	00B L3	1991–1995 LEP runs
290.1 ± 1.5 ± 1.1		BARATE	97R ALEP	1989–1994 LEP runs
289.2 ± 1.7 ± 1.2		ALEXANDER	96E OPAL	1990–1994 LEP runs
289.0 ± 2.8 ± 4.0	57.4k	BALEST	96 CLEO	E <sub>cm</sub> <sup>ee</sup> = 10.6 GeV
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
291.2 ± 2.0 ± 1.2		BARATE	97I ALEP	Repl. by BARATE 97R
291.4 ± 3.0		ABREU	96B DLPH	Repl. by ABDALLAH 04T
290.1 ± 4.0	34k	ACCIARRI	96K L3	Repl. by ACCIARRI 00B
297 ± 9 ± 5	1671	ABE	95Y SLD	1992–1993 SLC runs
304 ± 14 ± 7	4100	BATTLE	92 CLEO	E <sub>cm</sub> <sup>ee</sup> = 10.6 GeV
301 ± 29	3780	KLEINWORT	89 JADE	E <sub>cm</sub> <sup>ee</sup> = 35–46 GeV
288 ± 16 ± 17	807	AMIDEI	88 MRK2	E <sub>cm</sub> <sup>ee</sup> = 29 GeV
306 ± 20 ± 14	695	BRAUNSCH...	88C TASS	E <sub>cm</sub> <sup>ee</sup> = 36 GeV
299 ± 15 ± 10	1311	ABACHI	87C HRS	E <sub>cm</sub> <sup>ee</sup> = 29 GeV
295 ± 14 ± 11	5696	ALBRECHT	87P ARG	E <sub>cm</sub> <sup>ee</sup> = 9.3–10.6 GeV
309 ± 17 ± 7	3788	BAND	87B MAC	E <sub>cm</sub> <sup>ee</sup> = 29 GeV
325 ± 14 ± 18	8470	BEBEK	87C CLEO	E <sub>cm</sub> <sup>ee</sup> = 10.5 GeV
460 ± 190	102	FELDMAN	82 MRK2	E <sub>cm</sub> <sup>ee</sup> = 29 GeV

$$(\tau_{\tau^+} - \tau_{\tau^-}) / \tau_{\text{average}}$$

Test of *CPT* invariance.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;7.0 × 10<sup>-3</sup></b>	90	<sup>1</sup> BELOUS	14 BELL	711 fb <sup>-1</sup> E <sub>cm</sub> <sup>ee</sup> = 10.6 GeV
<sup>1</sup> BELOUS 14 quote limit on the absolute value of the relative lifetime difference.				

**$\tau$  MAGNETIC MOMENT ANOMALY**

The  $q^2$  dependence is expected to be small providing no thresholds are nearby.

$$\mu_\tau / (e\hbar/2m_\tau) - 1 = (g_\tau - 2)/2$$

For a theoretical calculation  $[(g_\tau - 2)/2 = 117\,721(5) \times 10^{-8}]$ , see EIDELMAN 07.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>-0.057 to 0.024</b>	95	<sup>1</sup> AAD	23BMATLS	$\gamma\gamma \rightarrow \tau^+\tau^-$ , Pb-Pb
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
-0.041 <sup>+0.012</sup> <sub>-0.009</sub>		<sup>1,2</sup> AAD	23BMATLS	$\gamma\gamma \rightarrow \tau^+\tau^-$ , Pb-Pb
0.001 <sup>+0.055</sup> <sub>-0.089</sub>		<sup>2,3</sup> TUMASYAN	23AS CMS	$\gamma\gamma \rightarrow \tau^+\tau^-$ , Pb-Pb
-0.018 ± 0.017		<sup>2,4</sup> ABDALLAH	04K DLPH	$e^+e^- \rightarrow e^+e^-\tau^+\tau^-$
< 0.107	95	<sup>5</sup> ACHARD	04G L3	$e^+e^- \rightarrow e^+e^-\tau^+\tau^-$
-0.007 to 0.005	95	<sup>6</sup> GONZALEZ-S..00	RVUE	$e^+e^- \rightarrow \tau^+\tau^-$ and $W \rightarrow \tau\nu_\tau$
-0.052 to 0.058	95	<sup>7</sup> ACCIARRI	98E L3	1991–1995 LEP runs
-0.068 to 0.065	95	<sup>8</sup> ACKERSTAFF	98N OPAL	1990–1995 LEP runs
-0.004 to 0.006	95	<sup>9</sup> ESCRIBANO	97 RVUE	$Z \rightarrow \tau^+\tau^-$ at LEP
< 0.01	95	<sup>10</sup> ESCRIBANO	93 RVUE	$Z \rightarrow \tau^+\tau^-$ at LEP
< 0.12	90	GRIFOLS	91 RVUE	$Z \rightarrow \tau\tau\gamma$ at LEP
< 0.023	95	<sup>11</sup> SILVERMAN	83 RVUE	$e^+e^- \rightarrow \tau^+\tau^-$ at PETRA

<sup>1</sup> AAD 23BM measurement is derived from  $\gamma\gamma \rightarrow \tau^+\tau^-$  total cross-section from 1.44 nb<sup>-1</sup> LHC Pb-Pb collisions at  $\sqrt{s_{NN}} = 5.02$  TeV. Authors report both the measured value and the corresponding 95% CL limit.

<sup>2</sup> Measurement ill-suited for a standard average because its likelihood appears to be remarkably non-Gaussian and asymmetric according to the model-dependent extraction procedure and the reported 95% CL limits.

<sup>3</sup> TUMASYAN 23AS measurement is derived from  $\gamma\gamma \rightarrow \tau^+\tau^-$  total cross-section from 404  $\mu\text{b}^{-1}$  LHC Pb-Pb collisions at  $\sqrt{s_{NN}} = 5.02$  TeV.

<sup>4</sup> ABDALLAH 04K measurement is derived from  $e^+e^- \rightarrow e^+e^-\tau^+\tau^-$  total cross-section measurements at  $\sqrt{s}$  between 183 and 208 GeV. In addition to the measurement, the authors also quote 95% CL limits of  $> -0.052$  and  $< 0.013$ .

<sup>5</sup> ACHARD 04G limit is derived from  $e^+e^- \rightarrow e^+e^-\tau^+\tau^-$  total cross-section measurements at  $\sqrt{s}$  between 189 and 206 GeV, and is on the absolute value of the magnetic moment anomaly.

<sup>6</sup> GONZALEZ-SPRINBERG 00 use data on tau lepton production at LEP1, SLC, and LEP2, and data from colliders and LEP2 to determine limits. Assume imaginary component is zero.

<sup>7</sup> ACCIARRI 98E use  $Z \rightarrow \tau^+\tau^-\gamma$  events. In addition to the limits, the authors also quote a value of  $0.004 \pm 0.027 \pm 0.023$ .

<sup>8</sup> ACKERSTAFF 98N use  $Z \rightarrow \tau^+\tau^-\gamma$  events. The limit applies to an average of the form factor for off-shell  $\tau$ 's having  $p^2$  ranging from  $m_\tau^2$  to  $(M_Z - m_\tau)^2$ .

<sup>9</sup> ESCRIBANO 97 use preliminary experimental results.

<sup>10</sup> ESCRIBANO 93 limit derived from  $\Gamma(Z \rightarrow \tau^+\tau^-)$ , and is on the absolute value of the magnetic moment anomaly.

<sup>11</sup> SILVERMAN 83 limit is derived from  $e^+e^- \rightarrow \tau^+\tau^-$  total cross-section measurements for  $q^2$  up to  $(37 \text{ GeV})^2$ .

**$\tau$  ELECTRIC DIPOLE MOMENT ( $d_\tau$ )**

A nonzero value is forbidden by both  $T$  invariance and  $P$  invariance.

The  $q^2$  dependence is expected to be small providing no thresholds are nearby.

**Re( $d_\tau$ )**

VALUE ( $10^{-16}$ ecm)	CL%	DOCUMENT ID	TECN	COMMENT
– <b>0.185 to 0.061</b>	95	<sup>1</sup> INAMI	22 BELL	$E_{\text{cm}}^{ee} = 10.6$ GeV
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
< 2.3	90	<sup>2</sup> GROZIN	09A RVUE	From $e$ EDM limit
< 3.7	95	<sup>3</sup> ABDALLAH	04K DLPH	$e^+e^- \rightarrow e^+e^-\tau^+\tau^-$ at LEP2
< 11.4	95	<sup>4</sup> ACHARD	04G L3	$e^+e^- \rightarrow e^+e^-\tau^+\tau^-$ at LEP2
– 0.22 to 0.45	95	<sup>5</sup> INAMI	03 BELL	$E_{\text{cm}}^{ee} = 10.6$ GeV
< 4.6	95	<sup>6</sup> ALBRECHT	00 ARG	$E_{\text{cm}}^{ee} = 10.4$ GeV
> –3.1 and < 3.1	95	ACCIARRI	98E L3	1991–1995 LEP runs
> –3.8 and < 3.6	95	<sup>7</sup> ACKERSTAFF	98N OPAL	1990–1995 LEP runs
< 0.11	95	<sup>8,9</sup> ESCRIBANO	97 RVUE	$Z \rightarrow \tau^+\tau^-$ at LEP
< 0.5	95	<sup>10</sup> ESCRIBANO	93 RVUE	$Z \rightarrow \tau^+\tau^-$ at LEP
< 7	90	GRIFOLS	91 RVUE	$Z \rightarrow \tau\tau\gamma$ at LEP
< 1.6	90	DELAGUILA	90 RVUE	$e^+e^- \rightarrow \tau^+\tau^-$ $E_{\text{cm}}^{ee} = 35$ GeV

<sup>1</sup> INAMI 22 use  $e^+e^- \rightarrow \tau^+\tau^-$  events from  $833 \text{ fb}^{-1}$  of data. Also report a measurement of  $\text{Re}(d_\tau) = (-0.62 \pm 0.63) \times 10^{-17}$  ecm.

<sup>2</sup> GROZIN 09A calculate the contribution to the electron electric dipole moment from the  $\tau$  electric dipole moment appearing in loops, which is  $\Delta d_e = 6.9 \times 10^{-12} d_\tau$ . Dividing the REGAN 02 upper limit  $|d_e| \leq 1.6 \times 10^{-27}$  e cm at CL=90% by  $6.9 \times 10^{-12}$  gives this limit.

<sup>3</sup> ABDALLAH 04K limit is derived from  $e^+e^- \rightarrow e^+e^-\tau^+\tau^-$  total cross-section measurements at  $\sqrt{s}$  between 183 and 208 GeV and is on the absolute value of  $d_\tau$ .

<sup>4</sup> ACHARD 04G limit is derived from  $e^+e^- \rightarrow e^+e^-\tau^+\tau^-$  total cross-section measurements at  $\sqrt{s}$  between 189 and 206 GeV, and is on the absolute value of  $d_\tau$ .

<sup>5</sup> INAMI 03 use  $e^+e^- \rightarrow \tau^+\tau^-$  events.

<sup>6</sup> ALBRECHT 00 use  $e^+e^- \rightarrow \tau^+\tau^-$  events. Limit is on the absolute value of  $\text{Re}(d_\tau)$ .

<sup>7</sup> ACKERSTAFF 98N use  $Z \rightarrow \tau^+\tau^-\gamma$  events. The limit applies to an average of the form factor for off-shell  $\tau$ 's having  $p^2$  ranging from  $m_\tau^2$  to  $(M_Z - m_\tau)^2$ .

<sup>8</sup> ESCRIBANO 97 derive the relationship  $|d_\tau| = \cot \theta_W |d_\tau^W|$  using effective Lagrangian methods, and use a conference result  $|d_\tau^W| < 5.8 \times 10^{-18}$  e cm at 95% CL (L. Silvestris, ICHEP96) to obtain this result.

<sup>9</sup> ESCRIBANO 97 use preliminary experimental results.

<sup>10</sup> ESCRIBANO 93 limit derived from  $\Gamma(Z \rightarrow \tau^+\tau^-)$ , and is on the absolute value of the electric dipole moment.

**Im( $d_\tau$ )**

VALUE ( $10^{-16}$ ecm)	CL%	DOCUMENT ID	TECN	COMMENT
<b>-0.103 to 0.023</b>	95	<sup>1</sup> INAMI 22	BELL	$E_{\text{cm}}^{ee} = 10.6$ GeV
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
-0.25 to 0.008	95	<sup>2</sup> INAMI 03	BELL	$E_{\text{cm}}^{ee} = 10.6$ GeV
< 1.8	95	<sup>3</sup> ALBRECHT 00	ARG	$E_{\text{cm}}^{ee} = 10.4$ GeV
<sup>1</sup> INAMI 22 use $e^+e^- \rightarrow \tau^+\tau^-$ events from $833 \text{ fb}^{-1}$ of data. Also report a measurement of $\text{Im}(d_\tau) = (-0.40 \pm 0.32) \times 10^{-17}$ ecm.				
<sup>2</sup> INAMI 03 use $e^+e^- \rightarrow \tau^+\tau^-$ events.				
<sup>3</sup> ALBRECHT 00 use $e^+e^- \rightarrow \tau^+\tau^-$ events. Limit is on the absolute value of $\text{Im}(d_\tau)$ .				

 **$\tau$  WEAK DIPOLE MOMENT ( $d_\tau^W$ )**

A nonzero value is forbidden by  $CP$  invariance.

The  $q^2$  dependence is expected to be small providing no thresholds are nearby.

**Re( $d_\tau^W$ )**

VALUE ( $10^{-17}$ ecm)	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.50</b>	95	<sup>1</sup> HEISTER 03F	ALEP	1990–1995 LEP runs
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
<3.0	90	<sup>1</sup> ACCIARRI 98C	L3	1991–1995 LEP runs
<0.56	95	ACKERSTAFF 97L	OPAL	1991–1995 LEP runs
<0.78	95	<sup>2</sup> AKERS 95F	OPAL	Repl. by ACKERSTAFF 97L
<1.5	95	<sup>2</sup> BUSKULIC 95C	ALEP	Repl. by HEISTER 03F
<7.0	95	<sup>2</sup> ACTON 92F	OPAL	$Z \rightarrow \tau^+\tau^-$ at LEP
<3.7	95	<sup>2</sup> BUSKULIC 92J	ALEP	Repl. by BUSKULIC 95C
<sup>1</sup> Limit is on the absolute value of the real part of the weak dipole moment.				
<sup>2</sup> Limit is on the absolute value of the real part of the weak dipole moment, and applies for $q^2 = m_Z^2$ .				

**Im( $d_\tau^W$ )**

VALUE ( $10^{-17}$ ecm)	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;1.1</b>	95	<sup>1</sup> HEISTER 03F	ALEP	1990–1995 LEP runs
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
<1.5	95	ACKERSTAFF 97L	OPAL	1991–1995 LEP runs
<4.5	95	<sup>2</sup> AKERS 95F	OPAL	Repl. by ACKERSTAFF 97L
<sup>1</sup> HEISTER 03F limit is on the absolute value of the imaginary part of the weak dipole moment.				
<sup>2</sup> Limit is on the absolute value of the imaginary part of the weak dipole moment, and applies for $q^2 = m_Z^2$ .				

## $\tau$ WEAK ANOMALOUS MAGNETIC DIPOLE MOMENT ( $\alpha_\tau^W$ )

Electroweak radiative corrections are expected to contribute at the  $10^{-6}$  level. See BERNABEU 95.

The  $q^2$  dependence is expected to be small providing no thresholds are nearby.

### Re( $\alpha_\tau^W$ )

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<1.1 \times 10^{-3}$	95	<sup>1</sup> HEISTER 03F	ALEP	1990–1995 LEP runs
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
$> -0.0024$ and $< 0.0025$	95	<sup>2</sup> GONZALEZ-S..00	RVUE	$e^+ e^- \rightarrow \tau^+ \tau^-$ and $W \rightarrow \tau \nu_\tau$
$<4.5 \times 10^{-3}$	90	<sup>1</sup> ACCIARRI 98C	L3	1991–1995 LEP runs

<sup>1</sup>Limit is on the absolute value of the real part of the weak anomalous magnetic dipole moment.

<sup>2</sup>GONZALEZ-SPRINBERG 00 use data on tau lepton production at LEP1, SLC, and LEP2, and data from colliders and LEP2 to determine limits. Assume imaginary component is zero.

### Im( $\alpha_\tau^W$ )

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<2.7 \times 10^{-3}$	95	<sup>1</sup> HEISTER 03F	ALEP	1990–1995 LEP runs
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
$<9.9 \times 10^{-3}$	90	<sup>1</sup> ACCIARRI 98C	L3	1991–1995 LEP runs

<sup>1</sup>Limit is on the absolute value of the imaginary part of the weak anomalous magnetic dipole moment.

## $\tau^-$ DECAY MODES

$\tau^+$  modes are charge conjugates of the modes below. “ $h^\pm$ ” stands for  $\pi^\pm$  or  $K^\pm$ . “ $\ell$ ” stands for  $e$  or  $\mu$ . “Neutrals” stands for  $\gamma$ 's and/or  $\pi^0$ 's.

Mode	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor/ Confidence level
<b>Modes with one charged particle</b>		
$\Gamma_1$ particle <sup>-</sup> $\geq 0$ neutrals $\geq 0 K^0 \nu_\tau$ (“1-prong”)	$(85.24 \pm 0.06) \%$	
$\Gamma_2$ particle <sup>-</sup> $\geq 0$ neutrals $\geq 0 K_L^0 \nu_\tau$	$(84.58 \pm 0.06) \%$	
$\Gamma_3$ $\mu^- \bar{\nu}_\mu \nu_\tau$	[a] $(17.39 \pm 0.04) \%$	
$\Gamma_4$ $\mu^- \bar{\nu}_\mu \nu_\tau \gamma$	[b] $(3.67 \pm 0.08) \times 10^{-3}$	
$\Gamma_5$ $e^- \bar{\nu}_e \nu_\tau$	[a] $(17.82 \pm 0.04) \%$	
$\Gamma_6$ $e^- \bar{\nu}_e \nu_\tau \gamma$	[b] $(1.83 \pm 0.05) \%$	
$\Gamma_7$ $h^- \geq 0 K_L^0 \nu_\tau$	$(12.03 \pm 0.05) \%$	
$\Gamma_8$ $h^- \nu_\tau$	$(11.51 \pm 0.05) \%$	
$\Gamma_9$ $\pi^- \nu_\tau$	[a] $(10.82 \pm 0.05) \%$	
$\Gamma_{10}$ $K^- \nu_\tau$	[a] $(6.96 \pm 0.10) \times 10^{-3}$	

$\Gamma_{11}$	$h^- \geq 1 \text{ neutrals } \nu_\tau$		$(37.00 \pm 0.09) \%$	
$\Gamma_{12}$	$h^- \geq 1 \pi^0 \nu_\tau (\text{ex. } K^0)$		$(36.50 \pm 0.09) \%$	
$\Gamma_{13}$	$h^- \pi^0 \nu_\tau$		$(25.93 \pm 0.09) \%$	
$\Gamma_{14}$	$\pi^- \pi^0 \nu_\tau$	[a]	$(25.49 \pm 0.09) \%$	
$\Gamma_{15}$	$\pi^- \pi^0 \text{non-}\rho(770) \nu_\tau$		$(3.0 \pm 3.2) \times 10^{-3}$	
$\Gamma_{16}$	$K^- \pi^0 \nu_\tau$	[a]	$(4.33 \pm 0.15) \times 10^{-3}$	
$\Gamma_{17}$	$h^- \geq 2 \pi^0 \nu_\tau$		$(10.81 \pm 0.09) \%$	
$\Gamma_{18}$	$h^- 2 \pi^0 \nu_\tau$		$(9.48 \pm 0.10) \%$	
$\Gamma_{19}$	$h^- 2 \pi^0 \nu_\tau (\text{ex. } K^0)$		$(9.32 \pm 0.10) \%$	
$\Gamma_{20}$	$\pi^- 2 \pi^0 \nu_\tau (\text{ex. } K^0)$	[a]	$(9.26 \pm 0.10) \%$	
$\Gamma_{21}$	$\pi^- 2 \pi^0 \nu_\tau (\text{ex. } K^0),$		$< 9 \times 10^{-3}$	CL=95%
$\Gamma_{22}$	scalar $\pi^- 2 \pi^0 \nu_\tau (\text{ex. } K^0),$		$< 7 \times 10^{-3}$	CL=95%
$\Gamma_{23}$	vector $K^- 2 \pi^0 \nu_\tau (\text{ex. } K^0)$	[a]	$(6.5 \pm 2.2) \times 10^{-4}$	
$\Gamma_{24}$	$h^- \geq 3 \pi^0 \nu_\tau$		$(1.34 \pm 0.07) \%$	
$\Gamma_{25}$	$h^- \geq 3 \pi^0 \nu_\tau (\text{ex. } K^0)$		$(1.25 \pm 0.07) \%$	
$\Gamma_{26}$	$h^- 3 \pi^0 \nu_\tau$		$(1.18 \pm 0.07) \%$	
$\Gamma_{27}$	$\pi^- 3 \pi^0 \nu_\tau (\text{ex. } K^0)$	[a]	$(1.04 \pm 0.07) \%$	
$\Gamma_{28}$	$K^- 3 \pi^0 \nu_\tau (\text{ex. } K^0, \eta)$	[a]	$(4.8 \pm 2.1) \times 10^{-4}$	
$\Gamma_{29}$	$h^- 4 \pi^0 \nu_\tau (\text{ex. } K^0)$		$(1.6 \pm 0.4) \times 10^{-3}$	
$\Gamma_{30}$	$h^- 4 \pi^0 \nu_\tau (\text{ex. } K^0, \eta)$	[a]	$(1.1 \pm 0.4) \times 10^{-3}$	
$\Gamma_{31}$	$a_1(1260) \nu_\tau \rightarrow \pi^- \gamma \nu_\tau$		$(4.0 \pm 1.5) \times 10^{-4}$	
$\Gamma_{32}$	$K^- \geq 0 \pi^0 \geq 0 K^0 \geq 0 \gamma \nu_\tau$		$(1.552 \pm 0.029) \%$	
$\Gamma_{33}$	$K^- \geq 1 (\pi^0 \text{ or } K^0 \text{ or } \gamma) \nu_\tau$		$(8.59 \pm 0.28) \times 10^{-3}$	

### Modes with $K^0$ 's

$\Gamma_{34}$	$K_S^0 (\text{particles})^- \nu_\tau$		$(9.43 \pm 0.28) \times 10^{-3}$	
$\Gamma_{35}$	$h^- \bar{K}^0 \nu_\tau$		$(9.87 \pm 0.14) \times 10^{-3}$	
$\Gamma_{36}$	$\pi^- \bar{K}^0 \nu_\tau$	[a]	$(8.38 \pm 0.14) \times 10^{-3}$	
$\Gamma_{37}$	$\pi^- \bar{K}^0 (\text{non-} K^*(892)^-) \nu_\tau$		$(5.4 \pm 2.1) \times 10^{-4}$	
$\Gamma_{38}$	$K^- K^0 \nu_\tau$	[a]	$(1.486 \pm 0.034) \times 10^{-3}$	
$\Gamma_{39}$	$K^- K^0 \geq 0 \pi^0 \nu_\tau$		$(2.99 \pm 0.07) \times 10^{-3}$	
$\Gamma_{40}$	$h^- \bar{K}^0 \pi^0 \nu_\tau$		$(5.32 \pm 0.13) \times 10^{-3}$	
$\Gamma_{41}$	$\pi^- \bar{K}^0 \pi^0 \nu_\tau$	[a]	$(3.82 \pm 0.13) \times 10^{-3}$	
$\Gamma_{42}$	$\bar{K}^0 \rho^- \nu_\tau$		$(2.2 \pm 0.5) \times 10^{-3}$	
$\Gamma_{43}$	$K^- K^0 \pi^0 \nu_\tau$	[a]	$(1.50 \pm 0.07) \times 10^{-3}$	
$\Gamma_{44}$	$\pi^- \bar{K}^0 \geq 1 \pi^0 \nu_\tau$		$(4.08 \pm 0.25) \times 10^{-3}$	
$\Gamma_{45}$	$\pi^- \bar{K}^0 \pi^0 \pi^0 \nu_\tau (\text{ex. } K^0)$	[a]	$(2.6 \pm 2.3) \times 10^{-4}$	
$\Gamma_{46}$	$K^- K^0 \pi^0 \pi^0 \nu_\tau$		$< 1.6 \times 10^{-4}$	CL=95%
$\Gamma_{47}$	$\pi^- K^0 \bar{K}^0 \nu_\tau$		$(1.55 \pm 0.24) \times 10^{-3}$	
$\Gamma_{48}$	$\pi^- K_S^0 K_S^0 \nu_\tau$	[a]	$(2.35 \pm 0.06) \times 10^{-4}$	
$\Gamma_{49}$	$\pi^- K_S^0 K_L^0 \nu_\tau$	[a]	$(1.08 \pm 0.24) \times 10^{-3}$	
$\Gamma_{50}$	$\pi^- K_L^0 K_L^0 \nu_\tau$		$(2.35 \pm 0.06) \times 10^{-4}$	

$\Gamma_{51}$	$\pi^- K^0 \bar{K}^0 \pi^0 \nu_\tau$		$( 3.6 \pm 1.2 ) \times 10^{-4}$	
$\Gamma_{52}$	$\pi^- K_S^0 K_S^0 \pi^0 \nu_\tau$	[a]	$( 1.82 \pm 0.21 ) \times 10^{-5}$	
$\Gamma_{53}$	$K^{*-} K^0 \pi^0 \nu_\tau \rightarrow$ $\pi^- K_S^0 K_S^0 \pi^0 \nu_\tau$		$( 1.08 \pm 0.21 ) \times 10^{-5}$	
$\Gamma_{54}$	$f_1(1285) \pi^- \nu_\tau \rightarrow$ $\pi^- K_S^0 K_S^0 \pi^0 \nu_\tau$		$( 6.8 \pm 1.5 ) \times 10^{-6}$	
$\Gamma_{55}$	$f_1(1420) \pi^- \nu_\tau \rightarrow$ $\pi^- K_S^0 K_S^0 \pi^0 \nu_\tau$		$( 2.4 \pm 0.8 ) \times 10^{-6}$	
$\Gamma_{56}$	$\pi^- K_S^0 K_L^0 \pi^0 \nu_\tau$	[a]	$( 3.2 \pm 1.2 ) \times 10^{-4}$	
$\Gamma_{57}$	$\pi^- K_L^0 K_L^0 \pi^0 \nu_\tau$		$( 1.82 \pm 0.21 ) \times 10^{-5}$	
$\Gamma_{58}$	$K^- K_S^0 K_S^0 \nu_\tau$		$< 6.3 \times 10^{-7}$	CL=90%
$\Gamma_{59}$	$K^- K_S^0 K_S^0 \pi^0 \nu_\tau$		$< 4.0 \times 10^{-7}$	CL=90%
$\Gamma_{60}$	$K^0 h^+ h^- h^- \geq 0$ neutrals $\nu_\tau$		$< 1.7 \times 10^{-3}$	CL=95%
$\Gamma_{61}$	$K^0 h^+ h^- h^- \nu_\tau$	[a]	$( 2.5 \pm 2.0 ) \times 10^{-4}$	

### Modes with three charged particles

$\Gamma_{62}$	$h^- h^- h^+ \geq 0$ neutrals $\geq 0 K_L^0 \nu_\tau$		$( 15.20 \pm 0.06 ) \%$	
$\Gamma_{63}$	$h^- h^- h^+ \geq 0$ neutrals $\nu_\tau$ (ex. $K_S^0 \rightarrow \pi^+ \pi^-$ ) ("3-prong")		$( 14.55 \pm 0.06 ) \%$	
$\Gamma_{64}$	$h^- h^- h^+ \nu_\tau$		$( 9.80 \pm 0.05 ) \%$	
$\Gamma_{65}$	$h^- h^- h^+ \nu_\tau$ (ex. $K^0$ )		$( 9.46 \pm 0.05 ) \%$	
$\Gamma_{66}$	$h^- h^- h^+ \nu_\tau$ (ex. $K^0, \omega$ )		$( 9.43 \pm 0.05 ) \%$	
$\Gamma_{67}$	$\pi^- \pi^+ \pi^- \nu_\tau$		$( 9.31 \pm 0.05 ) \%$	
$\Gamma_{68}$	$\pi^- \pi^+ \pi^- \nu_\tau$ (ex. $K^0$ )		$( 9.02 \pm 0.05 ) \%$	
$\Gamma_{69}$	$\pi^- \pi^+ \pi^- \nu_\tau$ (ex. $K^0$ ), non-axial vector		$< 2.4 \%$	CL=95%
$\Gamma_{70}$	$\pi^- \pi^+ \pi^- \nu_\tau$ (ex. $K^0, \omega$ )	[a]	$( 8.99 \pm 0.05 ) \%$	
$\Gamma_{71}$	$h^- h^- h^+ \geq 1$ neutrals $\nu_\tau$		$( 5.29 \pm 0.05 ) \%$	
$\Gamma_{72}$	$h^- h^- h^+ \geq 1 \pi^0 \nu_\tau$ (ex. $K^0$ )		$( 5.09 \pm 0.05 ) \%$	
$\Gamma_{73}$	$h^- h^- h^+ \pi^0 \nu_\tau$		$( 4.76 \pm 0.05 ) \%$	
$\Gamma_{74}$	$h^- h^- h^+ \pi^0 \nu_\tau$ (ex. $K^0$ )		$( 4.57 \pm 0.05 ) \%$	
$\Gamma_{75}$	$h^- h^- h^+ \pi^0 \nu_\tau$ (ex. $K^0, \omega$ )		$( 2.79 \pm 0.07 ) \%$	
$\Gamma_{76}$	$\pi^- \pi^+ \pi^- \pi^0 \nu_\tau$		$( 4.62 \pm 0.05 ) \%$	
$\Gamma_{77}$	$\pi^- \pi^+ \pi^- \pi^0 \nu_\tau$ (ex. $K^0$ )		$( 4.49 \pm 0.05 ) \%$	
$\Gamma_{78}$	$\pi^- \pi^+ \pi^- \pi^0 \nu_\tau$ (ex. $K^0, \omega$ )	[a]	$( 2.74 \pm 0.07 ) \%$	
$\Gamma_{79}$	$h^- \rho \pi^0 \nu_\tau$			
$\Gamma_{80}$	$h^- \rho^+ h^- \nu_\tau$			
$\Gamma_{81}$	$h^- \rho^- h^+ \nu_\tau$			
$\Gamma_{82}$	$h^- h^- h^+ \geq 2 \pi^0 \nu_\tau$ (ex. $K^0$ )		$( 5.17 \pm 0.31 ) \times 10^{-3}$	
$\Gamma_{83}$	$h^- h^- h^+ 2 \pi^0 \nu_\tau$		$( 5.05 \pm 0.31 ) \times 10^{-3}$	
$\Gamma_{84}$	$h^- h^- h^+ 2 \pi^0 \nu_\tau$ (ex. $K^0$ )		$( 4.95 \pm 0.31 ) \times 10^{-3}$	
$\Gamma_{85}$	$h^- h^- h^+ 2 \pi^0 \nu_\tau$ (ex. $K^0, \omega, \eta$ )	[a]	$( 10 \pm 4 ) \times 10^{-4}$	



$\Gamma_{86}$	$h^- h^- h^+ 3\pi^0 \nu_\tau$		$(2.13 \pm 0.30) \times 10^{-4}$	
$\Gamma_{87}$	$2\pi^- \pi^+ 3\pi^0 \nu_\tau$ (ex. $K^0$ )		$(1.94 \pm 0.30) \times 10^{-4}$	
$\Gamma_{88}$	$2\pi^- \pi^+ 3\pi^0 \nu_\tau$ (ex. $K^0, \eta, f_1(1285)$ )		$(1.7 \pm 0.4) \times 10^{-4}$	
$\Gamma_{89}$	$2\pi^- \pi^+ 3\pi^0 \nu_\tau$ (ex. $K^0, \eta, \omega, f_1(1285)$ )	[a]	$(1.4 \pm 2.7) \times 10^{-5}$	
$\Gamma_{90}$	$K^- h^+ h^- \geq 0$ neutrals $\nu_\tau$		$(6.29 \pm 0.14) \times 10^{-3}$	
$\Gamma_{91}$	$K^- h^+ \pi^- \nu_\tau$ (ex. $K^0$ )		$(4.37 \pm 0.07) \times 10^{-3}$	
$\Gamma_{92}$	$K^- h^+ \pi^- \pi^0 \nu_\tau$ (ex. $K^0$ )		$(8.6 \pm 1.2) \times 10^{-4}$	
$\Gamma_{93}$	$K^- \pi^+ \pi^- \geq 0$ neutrals $\nu_\tau$		$(4.77 \pm 0.14) \times 10^{-3}$	
$\Gamma_{94}$	$K^- \pi^+ \pi^- \geq 0 \pi^0 \nu_\tau$ (ex. $K^0$ )		$(3.73 \pm 0.13) \times 10^{-3}$	
$\Gamma_{95}$	$K^- \pi^+ \pi^- \nu_\tau$		$(3.45 \pm 0.07) \times 10^{-3}$	
$\Gamma_{96}$	$K^- \pi^+ \pi^- \nu_\tau$ (ex. $K^0$ )		$(2.93 \pm 0.07) \times 10^{-3}$	
$\Gamma_{97}$	$K^- \pi^+ \pi^- \nu_\tau$ (ex. $K^0, \omega$ )	[a]	$(2.93 \pm 0.07) \times 10^{-3}$	
$\Gamma_{98}$	$K^- \rho^0 \nu_\tau \rightarrow K^- \pi^+ \pi^- \nu_\tau$		$(1.4 \pm 0.5) \times 10^{-3}$	
$\Gamma_{99}$	$K^- \pi^+ \pi^- \pi^0 \nu_\tau$		$(1.31 \pm 0.12) \times 10^{-3}$	
$\Gamma_{100}$	$K^- \pi^+ \pi^- \pi^0 \nu_\tau$ (ex. $K^0$ )		$(7.9 \pm 1.2) \times 10^{-4}$	
$\Gamma_{101}$	$K^- \pi^+ \pi^- \pi^0 \nu_\tau$ (ex. $K^0, \eta$ )		$(7.6 \pm 1.2) \times 10^{-4}$	
$\Gamma_{102}$	$K^- \pi^+ \pi^- \pi^0 \nu_\tau$ (ex. $K^0, \omega$ )		$(3.7 \pm 0.9) \times 10^{-4}$	
$\Gamma_{103}$	$K^- \pi^+ \pi^- \pi^0 \nu_\tau$ (ex. $K^0, \omega, \eta$ ) [a]		$(3.9 \pm 1.4) \times 10^{-4}$	
$\Gamma_{104}$	$K^- \pi^+ K^- \geq 0$ neut. $\nu_\tau$		$< 9 \times 10^{-4}$	CL=95%
$\Gamma_{105}$	$K^- K^+ \pi^- \geq 0$ neut. $\nu_\tau$		$(1.496 \pm 0.033) \times 10^{-3}$	
$\Gamma_{106}$	$K^- K^+ \pi^- \nu_\tau$	[a]	$(1.435 \pm 0.027) \times 10^{-3}$	
$\Gamma_{107}$	$K^- K^+ \pi^- \pi^0 \nu_\tau$	[a]	$(6.1 \pm 1.8) \times 10^{-5}$	
$\Gamma_{108}$	$K^- K^+ K^- \nu_\tau$		$(2.2 \pm 0.8) \times 10^{-5}$	S=5.4
$\Gamma_{109}$	$K^- K^+ K^- \nu_\tau$ (ex. $\phi$ )		$< 2.5 \times 10^{-6}$	CL=90%
$\Gamma_{110}$	$K^- K^+ K^- \pi^0 \nu_\tau$		$< 4.8 \times 10^{-6}$	CL=90%
$\Gamma_{111}$	$\pi^- K^+ \pi^- \geq 0$ neut. $\nu_\tau$		$< 2.5 \times 10^{-3}$	CL=95%
$\Gamma_{112}$	$e^- e^- e^+ \bar{\nu}_e \nu_\tau$		$(2.8 \pm 1.5) \times 10^{-5}$	
$\Gamma_{113}$	$\mu^- e^- e^+ \bar{\nu}_\mu \nu_\tau$		$< 3.2 \times 10^{-5}$	CL=90%
$\Gamma_{114}$	$\pi^- e^- e^+ \nu_\tau$		seen	
$\Gamma_{115}$	$\pi^- \mu^- \mu^+ \nu_\tau$		$< 1.14 \times 10^{-5}$	CL=90%

### Modes with five charged particles

$\Gamma_{116}$	$3h^- 2h^+ \geq 0$ neutrals $\nu_\tau$ (ex. $K_S^0 \rightarrow \pi^- \pi^+$ ) ("5-prong")		$(9.9 \pm 0.4) \times 10^{-4}$	
$\Gamma_{117}$	$3h^- 2h^+ \nu_\tau$ (ex. $K^0$ )		$(8.29 \pm 0.31) \times 10^{-4}$	
$\Gamma_{118}$	$3\pi^- 2\pi^+ \nu_\tau$ (ex. $K^0, \omega$ )		$(8.27 \pm 0.31) \times 10^{-4}$	
$\Gamma_{119}$	$3\pi^- 2\pi^+ \nu_\tau$ (ex. $K^0, \omega, f_1(1285)$ )	[a]	$(7.75 \pm 0.30) \times 10^{-4}$	
$\Gamma_{120}$	$K^- 2\pi^- 2\pi^+ \nu_\tau$ (ex. $K^0$ )	[a]	$(6 \pm 12) \times 10^{-7}$	
$\Gamma_{121}$	$K^+ 3\pi^- \pi^+ \nu_\tau$		$< 5.0 \times 10^{-6}$	CL=90%

$\Gamma_{122}$	$K^+ K^- 2\pi^- \pi^+ \nu_\tau$	$< 4.5$	$\times 10^{-7}$	CL=90%
$\Gamma_{123}$	$3h^- 2h^+ \pi^0 \nu_\tau$ (ex. $K^0$ )	$(1.65 \pm 0.11)$	$\times 10^{-4}$	
$\Gamma_{124}$	$3\pi^- 2\pi^+ \pi^0 \nu_\tau$ (ex. $K^0$ )	$(1.64 \pm 0.11)$	$\times 10^{-4}$	
$\Gamma_{125}$	$3\pi^- 2\pi^+ \pi^0 \nu_\tau$ (ex. $K^0, \eta,$ $f_1(1285)$ )	$(1.11 \pm 0.10)$	$\times 10^{-4}$	
$\Gamma_{126}$	$3\pi^- 2\pi^+ \pi^0 \nu_\tau$ (ex. $K^0, \eta, \omega,$ [a] $f_1(1285)$ )	$(3.8 \pm 0.9)$	$\times 10^{-5}$	
$\Gamma_{127}$	$K^- 2\pi^- 2\pi^+ \pi^0 \nu_\tau$ (ex. $K^0$ ) [a]	$(1.1 \pm 0.6)$	$\times 10^{-6}$	
$\Gamma_{128}$	$K^+ 3\pi^- \pi^+ \pi^0 \nu_\tau$	$< 8$	$\times 10^{-7}$	CL=90%
$\Gamma_{129}$	$3h^- 2h^+ 2\pi^0 \nu_\tau$	$< 3.4$	$\times 10^{-6}$	CL=90%

**Miscellaneous other allowed modes**

$\Gamma_{130}$	$(5\pi)^- \nu_\tau$	$(7.8 \pm 0.5)$	$\times 10^{-3}$	
$\Gamma_{131}$	$4h^- 3h^+ \geq 0$ neutrals $\nu_\tau$ ("7-prong")	$< 3.0$	$\times 10^{-7}$	CL=90%
$\Gamma_{132}$	$4h^- 3h^+ \nu_\tau$	$< 4.3$	$\times 10^{-7}$	CL=90%
$\Gamma_{133}$	$4h^- 3h^+ \pi^0 \nu_\tau$	$< 2.5$	$\times 10^{-7}$	CL=90%
$\Gamma_{134}$	$X^-(S=-1) \nu_\tau$	$(2.92 \pm 0.04)$	%	
$\Gamma_{135}$	$K^*(892)^- \geq 0$ neutrals $\geq$ $0K_L^0 \nu_\tau$	$(1.42 \pm 0.18)$	%	S=1.4
$\Gamma_{136}$	$K^*(892)^- \nu_\tau$	$(1.20 \pm 0.07)$	%	S=1.8
$\Gamma_{137}$	$K^*(892)^- \nu_\tau \rightarrow \pi^- \bar{K}^0 \nu_\tau$	$(7.82 \pm 0.26)$	$\times 10^{-3}$	
$\Gamma_{138}$	$K^*(892)^0 K^- \geq 0$ neutrals $\nu_\tau$	$(3.2 \pm 1.4)$	$\times 10^{-3}$	
$\Gamma_{139}$	$K^*(892)^0 K^- \nu_\tau$	$(2.1 \pm 0.4)$	$\times 10^{-3}$	
$\Gamma_{140}$	$\bar{K}^*(892)^0 \pi^- \geq 0$ neutrals $\nu_\tau$	$(3.8 \pm 1.7)$	$\times 10^{-3}$	
$\Gamma_{141}$	$\bar{K}^*(892)^0 \pi^- \nu_\tau$	$(2.2 \pm 0.5)$	$\times 10^{-3}$	
$\Gamma_{142}$	$(\bar{K}^*(892)\pi)^- \nu_\tau \rightarrow$ $\pi^- \bar{K}^0 \pi^0 \nu_\tau$	$(1.0 \pm 0.4)$	$\times 10^{-3}$	
$\Gamma_{143}$	$K_1(1270)^- \nu_\tau$	$(4.7 \pm 1.1)$	$\times 10^{-3}$	
$\Gamma_{144}$	$K_1(1400)^- \nu_\tau$	$(1.7 \pm 2.6)$	$\times 10^{-3}$	S=1.7
$\Gamma_{145}$	$K^*(1410)^- \nu_\tau$	$(1.5 \pm 1.4)$ $- 1.0$	$\times 10^{-3}$	
$\Gamma_{146}$	$K_0^*(1430)^- \nu_\tau$	$< 5$	$\times 10^{-4}$	CL=95%
$\Gamma_{147}$	$K_2^*(1430)^- \nu_\tau$	$< 3$	$\times 10^{-3}$	CL=95%
$\Gamma_{148}$	$a_0(980)^- \geq 0$ neutrals $\nu_\tau$	$< 9.9$	$\times 10^{-5}$	CL=95%
$\Gamma_{149}$	$\eta \pi^- \nu_\tau$	$< 9.9$	$\times 10^{-5}$	CL=95%
$\Gamma_{150}$	$\eta \pi^- \pi^0 \nu_\tau$ [a]	$(1.39 \pm 0.07)$	$\times 10^{-3}$	
$\Gamma_{151}$	$\eta \pi^- \pi^0 \pi^0 \nu_\tau$ [a]	$(1.9 \pm 0.4)$	$\times 10^{-4}$	
$\Gamma_{152}$	$\eta K^- \nu_\tau$ [a]	$(1.55 \pm 0.08)$	$\times 10^{-4}$	
$\Gamma_{153}$	$\eta K^*(892)^- \nu_\tau$	$(1.38 \pm 0.15)$	$\times 10^{-4}$	
$\Gamma_{154}$	$\eta K^- \pi^0 \nu_\tau$ [a]	$(4.8 \pm 1.2)$	$\times 10^{-5}$	
$\Gamma_{155}$	$\eta K^- \pi^0$ (non- $K^*(892)$ ) $\nu_\tau$	$< 3.5$	$\times 10^{-5}$	CL=90%
$\Gamma_{156}$	$\eta \bar{K}^0 \pi^- \nu_\tau$ [a]	$(9.4 \pm 1.5)$	$\times 10^{-5}$	
$\Gamma_{157}$	$\eta \bar{K}^0 \pi^- \pi^0 \nu_\tau$	$< 5.0$	$\times 10^{-5}$	CL=90%
$\Gamma_{158}$	$\eta K^- K^0 \nu_\tau$	$< 9.0$	$\times 10^{-6}$	CL=90%

$\Gamma_{159}$	$\eta\pi^+\pi^-\pi^-\geq 0$ neutrals $\nu_\tau$	$< 3$	$\times 10^{-3}$	CL=90%
$\Gamma_{160}$	$\eta\pi^-\pi^+\pi^-\nu_\tau$ (ex. $K^0$ )	[a] $(2.20 \pm 0.13)$	$\times 10^{-4}$	
$\Gamma_{161}$	$\eta\pi^-\pi^+\pi^-\nu_\tau$ (ex. $K^0, f_1(1285)$ )	$(9.9 \pm 1.6)$	$\times 10^{-5}$	
$\Gamma_{162}$	$\eta a_1(1260)^-\nu_\tau \rightarrow \eta\pi^-\rho^0\nu_\tau$	$< 3.9$	$\times 10^{-4}$	CL=90%
$\Gamma_{163}$	$\eta\eta\pi^-\nu_\tau$	$< 7.4$	$\times 10^{-6}$	CL=90%
$\Gamma_{164}$	$\eta\eta\pi^-\pi^0\nu_\tau$	$< 2.0$	$\times 10^{-4}$	CL=95%
$\Gamma_{165}$	$\eta\eta K^-\nu_\tau$	$< 3.0$	$\times 10^{-6}$	CL=90%
$\Gamma_{166}$	$\eta'(958)\pi^-\nu_\tau$	$< 4.0$	$\times 10^{-6}$	CL=90%
$\Gamma_{167}$	$\eta'(958)\pi^-\pi^0\nu_\tau$	$< 1.2$	$\times 10^{-5}$	CL=90%
$\Gamma_{168}$	$\eta'(958)K^-\nu_\tau$	$< 2.4$	$\times 10^{-6}$	CL=90%
$\Gamma_{169}$	$\phi\pi^-\nu_\tau$	$(3.4 \pm 0.6)$	$\times 10^{-5}$	
$\Gamma_{170}$	$\phi K^-\nu_\tau$	[a] $(4.4 \pm 1.6)$	$\times 10^{-5}$	
$\Gamma_{171}$	$f_1(1285)\pi^-\nu_\tau$	$(3.9 \pm 0.5)$	$\times 10^{-4}$	S=1.9
$\Gamma_{172}$	$f_1(1285)\pi^-\nu_\tau \rightarrow \eta\pi^-\pi^+\pi^-\nu_\tau$	$(1.18 \pm 0.07)$	$\times 10^{-4}$	S=1.3
$\Gamma_{173}$	$f_1(1285)\pi^-\nu_\tau \rightarrow 3\pi^-2\pi^+\nu_\tau$	[a] $(5.2 \pm 0.4)$	$\times 10^{-5}$	
$\Gamma_{174}$	$\pi(1300)^-\nu_\tau \rightarrow (\rho\pi)^-\nu_\tau \rightarrow (3\pi)^-\nu_\tau$	$< 1.0$	$\times 10^{-4}$	CL=90%
$\Gamma_{175}$	$\pi(1300)^-\nu_\tau \rightarrow ((\pi\pi)_{S\text{-wave}}\pi)^-\nu_\tau \rightarrow (3\pi)^-\nu_\tau$	$< 1.9$	$\times 10^{-4}$	CL=90%
$\Gamma_{176}$	$h^-\omega \geq 0$ neutrals $\nu_\tau$	$(2.40 \pm 0.08)$	%	
$\Gamma_{177}$	$h^-\omega\nu_\tau$	$(1.99 \pm 0.06)$	%	
$\Gamma_{178}$	$\pi^-\omega\nu_\tau$	[a] $(1.95 \pm 0.06)$	%	
$\Gamma_{179}$	$K^-\omega\nu_\tau$	[a] $(4.1 \pm 0.9)$	$\times 10^{-4}$	
$\Gamma_{180}$	$h^-\omega\pi^0\nu_\tau$	[a] $(4.1 \pm 0.4)$	$\times 10^{-3}$	
$\Gamma_{181}$	$h^-\omega 2\pi^0\nu_\tau$	$(1.4 \pm 0.5)$	$\times 10^{-4}$	
$\Gamma_{182}$	$\pi^-\omega 2\pi^0\nu_\tau$	[a] $(7.2 \pm 1.6)$	$\times 10^{-5}$	
$\Gamma_{183}$	$h^-2\omega\nu_\tau$	$< 5.4$	$\times 10^{-7}$	CL=90%
$\Gamma_{184}$	$2h^-h^+\omega\nu_\tau$	$(1.20 \pm 0.22)$	$\times 10^{-4}$	
$\Gamma_{185}$	$2\pi^-\pi^+\omega\nu_\tau$ (ex. $K^0$ )	[a] $(8.4 \pm 0.6)$	$\times 10^{-5}$	

### Lepton Family number ( $LF$ ), Lepton number ( $L$ ), or Baryon number ( $B$ ) violating modes

$L$  means lepton number violation (e.g.  $\tau^- \rightarrow e^+\pi^-\pi^-$ ). Following common usage,  $LF$  means lepton family violation *and not* lepton number violation (e.g.  $\tau^- \rightarrow e^-\pi^+\pi^-$ ).  $B$  means baryon number violation.

$\Gamma_{186}$	$e^-\gamma$	$LF$	$< 3.3$	$\times 10^{-8}$	CL=90%
$\Gamma_{187}$	$e^-\gamma\gamma$	$LF$	$< 2.5$	$\times 10^{-4}$	CL=90%
$\Gamma_{188}$	$\mu^-\gamma$	$LF$	$< 4.2$	$\times 10^{-8}$	CL=90%
$\Gamma_{189}$	$\mu^-\gamma\gamma$	$LF$	$< 5.8$	$\times 10^{-4}$	CL=90%
$\Gamma_{190}$	$e^-\pi^0$	$LF$	$< 8.0$	$\times 10^{-8}$	CL=90%
$\Gamma_{191}$	$\mu^-\pi^0$	$LF$	$< 1.1$	$\times 10^{-7}$	CL=90%
$\Gamma_{192}$	$e^-K_S^0$	$LF$	$< 2.6$	$\times 10^{-8}$	CL=90%

Γ <sub>193</sub>	$\mu^- K_S^0$	LF	< 2.3	$\times 10^{-8}$	CL=90%
Γ <sub>194</sub>	$e^- \eta$	LF	< 9.2	$\times 10^{-8}$	CL=90%
Γ <sub>195</sub>	$\mu^- \eta$	LF	< 6.5	$\times 10^{-8}$	CL=90%
Γ <sub>196</sub>	$e^- \rho^0$	LF	< 2.2	$\times 10^{-8}$	CL=90%
Γ <sub>197</sub>	$\mu^- \rho^0$	LF	< 1.7	$\times 10^{-8}$	CL=90%
Γ <sub>198</sub>	$e^- \omega$	LF	< 2.4	$\times 10^{-8}$	CL=90%
Γ <sub>199</sub>	$\mu^- \omega$	LF	< 3.9	$\times 10^{-8}$	CL=90%
Γ <sub>200</sub>	$e^- K^*(892)^0$	LF	< 1.9	$\times 10^{-8}$	CL=90%
Γ <sub>201</sub>	$\mu^- K^*(892)^0$	LF	< 2.9	$\times 10^{-8}$	CL=90%
Γ <sub>202</sub>	$e^- \bar{K}^*(892)^0$	LF	< 1.7	$\times 10^{-8}$	CL=90%
Γ <sub>203</sub>	$\mu^- \bar{K}^*(892)^0$	LF	< 4.3	$\times 10^{-8}$	CL=90%
Γ <sub>204</sub>	$e^- \eta'(958)$	LF	< 1.6	$\times 10^{-7}$	CL=90%
Γ <sub>205</sub>	$\mu^- \eta'(958)$	LF	< 1.3	$\times 10^{-7}$	CL=90%
Γ <sub>206</sub>	$e^- f_0(980) \rightarrow e^- \pi^+ \pi^-$	LF	< 3.2	$\times 10^{-8}$	CL=90%
Γ <sub>207</sub>	$\mu^- f_0(980) \rightarrow \mu^- \pi^+ \pi^-$	LF	< 3.4	$\times 10^{-8}$	CL=90%
Γ <sub>208</sub>	$e^- \phi$	LF	< 2.0	$\times 10^{-8}$	CL=90%
Γ <sub>209</sub>	$\mu^- \phi$	LF	< 2.3	$\times 10^{-8}$	CL=90%
Γ <sub>210</sub>	$e^- e^+ e^-$	LF	< 2.7	$\times 10^{-8}$	CL=90%
Γ <sub>211</sub>	$e^- \mu^+ \mu^-$	LF	< 2.7	$\times 10^{-8}$	CL=90%
Γ <sub>212</sub>	$e^+ \mu^- \mu^-$	LF	< 1.7	$\times 10^{-8}$	CL=90%
Γ <sub>213</sub>	$\mu^- e^+ e^-$	LF	< 1.8	$\times 10^{-8}$	CL=90%
Γ <sub>214</sub>	$\mu^+ e^- e^-$	LF	< 1.5	$\times 10^{-8}$	CL=90%
Γ <sub>215</sub>	$\mu^- \mu^+ \mu^-$	LF	< 2.1	$\times 10^{-8}$	CL=90%
Γ <sub>216</sub>	$e^- \pi^+ \pi^-$	LF	< 2.3	$\times 10^{-8}$	CL=90%
Γ <sub>217</sub>	$e^+ \pi^- \pi^-$	L	< 2.0	$\times 10^{-8}$	CL=90%
Γ <sub>218</sub>	$\mu^- \pi^+ \pi^-$	LF	< 2.1	$\times 10^{-8}$	CL=90%
Γ <sub>219</sub>	$\mu^+ \pi^- \pi^-$	L	< 3.9	$\times 10^{-8}$	CL=90%
Γ <sub>220</sub>	$e^- \pi^+ K^-$	LF	< 3.7	$\times 10^{-8}$	CL=90%
Γ <sub>221</sub>	$e^- \pi^- K^+$	LF	< 3.1	$\times 10^{-8}$	CL=90%
Γ <sub>222</sub>	$e^+ \pi^- K^-$	L	< 3.2	$\times 10^{-8}$	CL=90%
Γ <sub>223</sub>	$e^- K_S^0 K_S^0$	LF	< 7.1	$\times 10^{-8}$	CL=90%
Γ <sub>224</sub>	$e^- K^+ K^-$	LF	< 3.4	$\times 10^{-8}$	CL=90%
Γ <sub>225</sub>	$e^+ K^- K^-$	L	< 3.3	$\times 10^{-8}$	CL=90%
Γ <sub>226</sub>	$\mu^- \pi^+ K^-$	LF	< 8.6	$\times 10^{-8}$	CL=90%
Γ <sub>227</sub>	$\mu^- \pi^- K^+$	LF	< 4.5	$\times 10^{-8}$	CL=90%
Γ <sub>228</sub>	$\mu^+ \pi^- K^-$	L	< 4.8	$\times 10^{-8}$	CL=90%
Γ <sub>229</sub>	$\mu^- K_S^0 K_S^0$	LF	< 8.0	$\times 10^{-8}$	CL=90%
Γ <sub>230</sub>	$\mu^- K^+ K^-$	LF	< 4.4	$\times 10^{-8}$	CL=90%
Γ <sub>231</sub>	$\mu^+ K^- K^-$	L	< 4.7	$\times 10^{-8}$	CL=90%
Γ <sub>232</sub>	$e^- \pi^0 \pi^0$	LF	< 6.5	$\times 10^{-6}$	CL=90%
Γ <sub>233</sub>	$\mu^- \pi^0 \pi^0$	LF	< 1.4	$\times 10^{-5}$	CL=90%
Γ <sub>234</sub>	$e^- \eta \eta$	LF	< 3.5	$\times 10^{-5}$	CL=90%
Γ <sub>235</sub>	$\mu^- \eta \eta$	LF	< 6.0	$\times 10^{-5}$	CL=90%
Γ <sub>236</sub>	$e^- \pi^0 \eta$	LF	< 2.4	$\times 10^{-5}$	CL=90%

$\Gamma_{237}$	$\mu^- \pi^0 \eta$	$LF$	$< 2.2$	$\times 10^{-5}$	CL=90%
$\Gamma_{238}$	$p e^- e^-$	$L,B$	$< 3.0$	$\times 10^{-8}$	CL=90%
$\Gamma_{239}$	$\bar{p} e^+ e^-$	$L,B$	$< 3.0$	$\times 10^{-8}$	CL=90%
$\Gamma_{240}$	$\bar{p} e^+ \mu^-$	$L,B$	$< 2.0$	$\times 10^{-8}$	CL=90%
$\Gamma_{241}$	$\bar{p} e^- \mu^+$	$L,B$	$< 1.8$	$\times 10^{-8}$	CL=90%
$\Gamma_{242}$	$p \mu^- \mu^-$	$L,B$	$< 4.0$	$\times 10^{-8}$	CL=90%
$\Gamma_{243}$	$\bar{p} \mu^+ \mu^-$	$L,B$	$< 1.8$	$\times 10^{-8}$	CL=90%
$\Gamma_{244}$	$\bar{p} \gamma$	$L,B$	$< 3.5$	$\times 10^{-6}$	CL=90%
$\Gamma_{245}$	$\bar{p} \pi^0$	$L,B$	$< 1.5$	$\times 10^{-5}$	CL=90%
$\Gamma_{246}$	$\bar{p} 2\pi^0$	$L,B$	$< 3.3$	$\times 10^{-5}$	CL=90%
$\Gamma_{247}$	$\bar{p} \eta$	$L,B$	$< 8.9$	$\times 10^{-6}$	CL=90%
$\Gamma_{248}$	$\bar{p} \pi^0 \eta$	$L,B$	$< 2.7$	$\times 10^{-5}$	CL=90%
$\Gamma_{249}$	$\Lambda \pi^-$	$L,B$	$< 7.2$	$\times 10^{-8}$	CL=90%
$\Gamma_{250}$	$\bar{\Lambda} \pi^-$	$L,B$	$< 1.4$	$\times 10^{-7}$	CL=90%
$\Gamma_{251}$	$e^-$ light boson	$LF$	$< 9$	$\times 10^{-4}$	CL=95%
$\Gamma_{252}$	$\mu^-$ light boson	$LF$	$< 6$	$\times 10^{-4}$	CL=95%

[a] Basis mode for the  $\tau$ .

[b] See the Particle Listings below for the energy limits used in this measurement.

### CONSTRAINED FIT INFORMATION

An overall fit to 85 branching ratios uses 170 measurements and one constraint to determine 46 parameters. The overall fit has a  $\chi^2 = 135.0$  for 125 degrees of freedom.

The following *off-diagonal* array elements are the correlation coefficients  $\langle \delta x_i \delta x_j \rangle / (\delta x_i \cdot \delta x_j)$ , in percent, from the fit to the branching fractions,  $x_i \equiv \Gamma_i / \Gamma_{\text{total}}$ . The fit constrains the  $x_i$  whose labels appear in this array to sum to one.

x <sub>5</sub>	18									
x <sub>9</sub>	2	-1								
x <sub>10</sub>	3	4	5							
x <sub>14</sub>	-18	-19	-17	-5						
x <sub>16</sub>	-1	-1	0	-2	-9					
x <sub>20</sub>	-11	-11	-14	-4	-46	-1				
x <sub>23</sub>	-1	0	-2	-3	-1	-14	-10			
x <sub>27</sub>	-6	-5	-10	-1	0	1	-39	1		
x <sub>28</sub>	0	-1	-1	-2	0	-13	-3	-23	-11	
x <sub>30</sub>	-4	-3	-11	-1	-9	0	7	-2	-44	2
x <sub>36</sub>	-2	-2	-3	-1	-1	-1	-2	0	-1	0
x <sub>38</sub>	-1	-1	1	0	0	0	0	-2	0	-2
x <sub>41</sub>	-2	-2	-2	-1	-1	0	-2	0	-1	0
x <sub>43</sub>	-1	-1	-1	-1	0	-3	0	-5	0	-5
x <sub>45</sub>	-5	-5	-5	-2	-3	-1	-4	-2	-1	-2
x <sub>48</sub>	-1	-1	2	0	-1	2	-1	-1	0	-1
x <sub>49</sub>	-5	-5	-5	-2	-3	-1	-5	-2	-1	-2
x <sub>52</sub>	0	0	0	0	0	0	0	-1	0	-1
x <sub>56</sub>	-2	-2	-2	-1	-1	-1	-2	-1	-1	-1
x <sub>61</sub>	-5	-5	-5	-2	-3	-1	-4	-2	-1	-2
x <sub>70</sub>	-7	-9	4	-2	-6	3	-12	-2	-7	-1
x <sub>78</sub>	-4	-4	-5	0	-9	0	1	1	-1	1
x <sub>85</sub>	0	0	-2	0	-2	0	0	0	2	0
x <sub>89</sub>	0	0	0	0	0	0	0	0	0	0
x <sub>97</sub>	-2	-2	-1	-1	-1	-1	-3	-1	-2	-1
x <sub>103</sub>	1	1	0	-1	1	-1	-1	-1	0	-1
x <sub>106</sub>	-1	-2	2	-1	-1	1	-2	-1	-1	-1
x <sub>107</sub>	0	0	0	0	0	0	0	0	0	0
x <sub>119</sub>	-1	-1	1	0	-1	1	-2	-1	-1	0
x <sub>120</sub>	0	0	0	0	0	0	0	0	0	0
x <sub>126</sub>	0	0	0	0	0	0	0	0	0	0
x <sub>127</sub>	0	0	0	0	0	0	0	0	0	0
x <sub>150</sub>	-1	-1	-1	0	-1	0	-2	-1	0	-1
x <sub>151</sub>	-1	-1	0	0	-1	0	-1	0	0	0
x <sub>152</sub>	0	0	0	0	0	0	0	-1	0	-1
x <sub>154</sub>	0	0	0	0	0	0	0	0	0	0
x <sub>156</sub>	0	0	0	0	0	0	0	0	0	0
x <sub>160</sub>	-1	-1	1	0	-1	1	-1	0	-1	0
x <sub>170</sub>	0	0	0	0	0	0	0	0	0	0
x <sub>173</sub>	0	-1	1	0	0	1	-1	0	0	0
x <sub>178</sub>	-3	-3	-3	-1	-4	0	-1	0	-1	0
x <sub>179</sub>	0	0	0	0	0	0	0	0	0	0
x <sub>180</sub>	-2	-2	-5	-1	-3	0	-2	-1	2	-1
x <sub>182</sub>	0	0	0	0	0	0	0	0	0	0
x <sub>185</sub>	-1	-1	0	0	-1	1	-1	0	0	0
	x <sub>3</sub>	x <sub>5</sub>	x <sub>9</sub>	x <sub>10</sub>	x <sub>14</sub>	x <sub>16</sub>	x <sub>20</sub>	x <sub>23</sub>	x <sub>27</sub>	x <sub>28</sub>

x36	0									
x38	0	-15								
x41	0	-13	2							
x43	0	-1	-14	-20						
x45	0	-3	0	-6	0					
x48	0	-2	3	-4	1	0				
x49	0	-5	0	-4	-1	-10	-1			
x52	0	1	5	-1	6	0	-7	0		
x56	0	-2	0	-2	-1	-4	0	-8	0	
x61	0	-2	0	-2	0	-4	0	-4	0	-2
x70	-5	-2	3	-2	-1	-4	5	-4	0	-2
x78	3	1	-1	1	0	2	-1	2	0	1
x85	2	0	0	0	0	0	0	0	0	0
x89	0	0	0	0	0	0	0	0	-1	0
x97	-1	-1	0	-1	0	-2	0	-2	0	-1
x103	-1	-1	0	-1	0	-1	0	-1	0	-1
x106	-1	-1	1	0	0	-1	2	-1	0	0
x107	0	0	0	0	0	0	0	0	0	0
x119	-1	-1	1	0	0	-1	2	-1	0	0
x120	0	0	0	0	0	0	0	0	0	0
x126	0	0	0	0	0	0	0	0	0	0
x127	0	0	0	0	0	0	0	0	0	0
x150	-2	-1	0	0	0	-1	0	-1	0	0
x151	0	0	0	0	0	-1	0	-1	0	0
x152	0	0	1	0	0	0	1	0	0	0
x154	0	0	0	0	0	0	0	0	0	0
x156	0	0	0	0	0	0	0	-1	0	0
x160	-1	0	1	0	0	-1	1	-1	0	0
x170	0	0	0	0	0	0	0	0	0	0
x173	-1	0	1	0	0	0	1	0	0	0
x178	1	0	0	0	0	-1	0	-1	0	0
x179	0	0	0	0	0	0	0	0	0	0
x180	2	-1	0	0	0	-1	0	-1	0	0
x182	0	0	0	0	0	0	0	0	0	0
x185	-1	0	1	0	0	0	1	-1	0	0
	x30	x36	x38	x41	x43	x45	x48	x49	x52	x56

x70	-4										
x78	2	-19									
x85	0	-1	-8								
x89	0	-1	-1	0							
x97	-2	19	-6	0	0						
x103	-1	-4	-14	-1	0	-1					
x106	-1	15	-4	0	0	0	-1				
x107	0	-1	-1	0	0	0	-3	0			
x119	-1	3	-1	0	-4	-1	0	1	0		
x120	0	0	0	0	0	0	0	0	0	-1	
x126	0	0	0	0	0	0	0	0	0	0	3
x127	0	0	0	0	0	0	0	0	0	0	-1
x150	-1	0	0	-5	0	0	0	0	0	0	0
x151	-1	0	0	0	-11	0	0	0	0	0	9
x152	0	2	0	0	0	0	-1	1	0	0	1
x154	0	0	0	-1	0	0	0	0	0	0	0
x156	0	0	0	0	-2	0	0	0	0	0	0
x160	-1	1	-1	0	-8	-1	0	1	0	0	46
x170	0	-1	0	0	0	1	0	1	0	0	0
x173	0	1	0	0	-2	0	0	1	0	0	34
x178	-1	-9	-67	-3	0	-2	10	-2	0	0	-1
x179	0	0	12	0	0	-2	-58	0	0	0	0
x180	-1	-2	-11	-64	-1	-1	-1	-1	0	0	0
x182	0	0	0	0	-16	0	0	0	0	0	7
x185	0	1	-1	0	-4	0	0	1	0	0	39
	x61	x70	x78	x85	x89	x97	x103	x106	x107	x119	

x126	0										
x127	0	-1									
x150	0	0	0								
x151	0	2	0	0							
x152	0	0	0	4	0						
x154	0	0	0	1	0	1					
x156	0	0	0	2	-1	1	0				
x160	-1	3	-1	0	25	0	0	0			
x170	0	0	0	0	0	0	0	0	0		
x173	-1	1	0	0	4	0	0	0	20	0	
x178	0	0	0	0	0	0	0	0	0	0	0
x179	0	0	0	0	0	0	0	0	0	0	0
x180	0	0	0	0	0	0	0	0	0	0	0
x182	0	2	0	0	10	0	0	-1	20	0	0
x185	-1	-2	-1	0	17	0	0	0	38	0	0
	x120	x126	x127	x150	x151	x152	x154	x156	x160	x170	



$x_{178}$	0				
$x_{179}$	0	-14			
$x_{180}$	0	-4	0		
$x_{182}$	3	0	0	0	
$x_{185}$	17	0	0	0	14

See the related review(s):  
[τ Branching Fractions](#)

$$(\Gamma(\tau^+) - \Gamma(\tau^-)) / (\Gamma(\tau^+) + \Gamma(\tau^-))$$

$\tau^\pm \rightarrow \pi^\pm K_S^0 \nu_\tau$  (RATE DIFFERENCE) / (RATE SUM)

VALUE (%)	DOCUMENT ID	TECN	COMMENT
$-0.36 \pm 0.23 \pm 0.11$	LEES	12M BABR	476 fb <sup>-1</sup> $E_{\text{cm}}^{ee} = 10.6$ GeV

### τ<sup>-</sup> BRANCHING RATIOS

$$\Gamma(\text{particle}^- \geq 0 \text{ neutrals} \geq 0 K^0 \nu_\tau \text{ ("1-prong")}) / \Gamma_{\text{total}} \quad \Gamma_1 / \Gamma$$

$$\Gamma_1 / \Gamma = (\Gamma_3 + \Gamma_5 + \Gamma_9 + \Gamma_{10} + \Gamma_{14} + \Gamma_{16} + \Gamma_{20} + \Gamma_{23} + \Gamma_{27} + \Gamma_{28} + \Gamma_{30} + \Gamma_{36} + \Gamma_{38} + \Gamma_{41} + \Gamma_{43} + \Gamma_{45} + \Gamma_{48} + \Gamma_{49} + \Gamma_{50} + \Gamma_{52} + \Gamma_{56} + \Gamma_{57} + 0.7196\Gamma_{150} + 0.7196\Gamma_{152} + 0.7196\Gamma_{154} + 0.7196\Gamma_{156} + 0.339\Gamma_{170} + 0.0835\Gamma_{178} + 0.0835\Gamma_{179} + 0.0835\Gamma_{180}) / \Gamma$$

The charged particle here can be  $e$ ,  $\mu$ , or hadron. In many analyses, the sum of the topological branching fractions (1, 3, and 5 prongs) is constrained to be unity. Since the 5-prong fraction is very small, the measured 1-prong and 3-prong fractions are highly correlated and cannot be treated as independent quantities in our overall fit. We arbitrarily choose to use the 3-prong fraction in our fit, and leave the 1-prong fraction out. We do, however, use these 1-prong measurements in our average below.

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>85.24 ± 0.06 OUR FIT</b>				(Produced by HFLAV)
<b>85.26 ± 0.13 OUR AVERAGE</b>				Error includes scale factor of 1.6. See the ideogram below.

• • • We use the following data for averages but not for fits. • • •

$85.316 \pm 0.093 \pm 0.049$	78k	<sup>1</sup> ABREU	01M DLPH	1992–1995 LEP runs
$85.274 \pm 0.105 \pm 0.073$		<sup>2</sup> ACHARD	01D L3	1992–1995 LEP runs
$84.48 \pm 0.27 \pm 0.23$		ACTON	92H OPAL	1990–1991 LEP runs

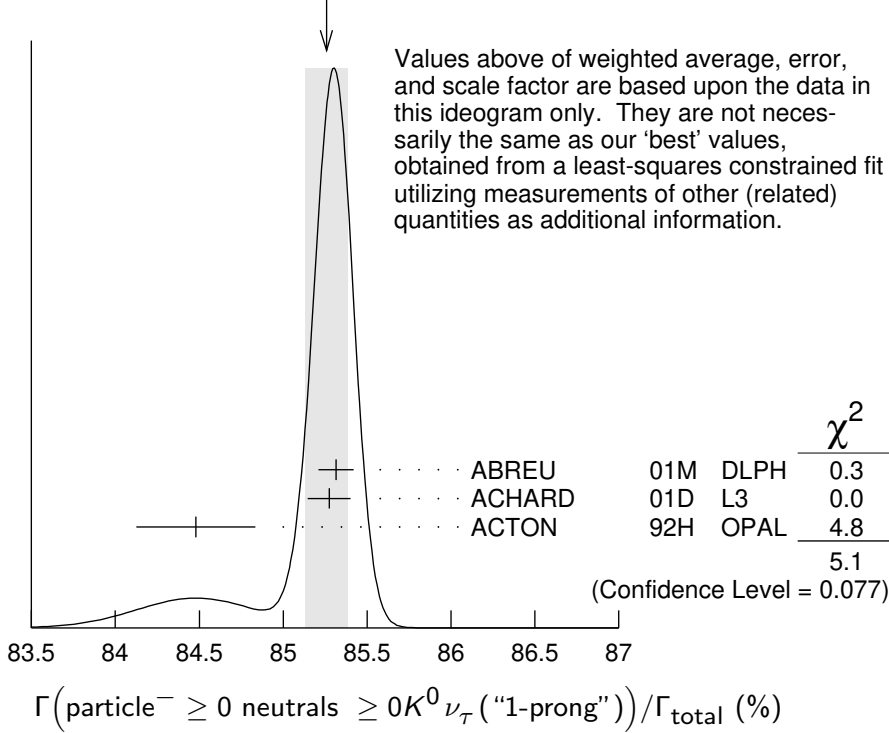
• • • We do not use the following data for averages, fits, limits, etc. • • •

$85.45 \begin{smallmatrix} +0.69 \\ -0.73 \end{smallmatrix} \pm 0.65$		DECAMP	92C ALEP	Repl. by SCHAELE 05c
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<sup>1</sup> The correlation coefficients between this measurement and the ABREU 01M measurements of  $B(\tau \rightarrow 3\text{-prong})$  and  $B(\tau \rightarrow 5\text{-prong})$  are  $-0.98$  and  $-0.08$  respectively.

<sup>2</sup> The correlation coefficients between this measurement and the ACHARD 01D measurements of  $B(\tau \rightarrow 3\text{-prong})$  and  $B(\tau \rightarrow 5\text{-prong})$  are  $-0.978$  and  $-0.082$  respectively.

WEIGHTED AVERAGE  
85.26±0.13 (Error scaled by 1.6)



$$\Gamma(\text{particle}^- \geq 0 \text{ neutrals} \geq 0K_L^0 \nu_\tau) / \Gamma_{\text{total}} \quad \Gamma_2 / \Gamma$$

$$\Gamma_2 / \Gamma = (\Gamma_3 + \Gamma_5 + \Gamma_9 + \Gamma_{10} + \Gamma_{14} + \Gamma_{16} + \Gamma_{20} + \Gamma_{23} + \Gamma_{27} + \Gamma_{28} + \Gamma_{30} + 0.6534\Gamma_{36} + 0.6534\Gamma_{38} + 0.6534\Gamma_{41} + 0.6534\Gamma_{43} + 0.6534\Gamma_{45} + 0.0942\Gamma_{48} + 0.3069\Gamma_{49} + \Gamma_{50} + 0.0942\Gamma_{52} + 0.3069\Gamma_{56} + \Gamma_{57} + 0.7196\Gamma_{150} + 0.7196\Gamma_{152} + 0.7196\Gamma_{154} + 0.4702\Gamma_{156} + 0.1049\Gamma_{170} + 0.0835\Gamma_{178} + 0.0835\Gamma_{179} + 0.0835\Gamma_{180}) / \Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
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**84.58±0.06 OUR FIT** (Produced by HFLAV)

**85.1 ±0.4 OUR AVERAGE**

• • • We use the following data for averages but not for fits. • • •

85.6 ±0.6 ±0.3	3300	<sup>1</sup> ADEVA	91F L3	$E_{\text{cm}}^{ee} = 88.3\text{--}94.3 \text{ GeV}$
84.9 ±0.4 ±0.3		BEHREND	89B CELL	$E_{\text{cm}}^{ee} = 14\text{--}47 \text{ GeV}$
84.7 ±0.8 ±0.6		<sup>2</sup> AIHARA	87B TPC	$E_{\text{cm}}^{ee} = 29 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

86.4 ±0.3 ±0.3		ABACHI	89B HRS	$E_{\text{cm}}^{ee} = 29 \text{ GeV}$
87.1 ±1.0 ±0.7		<sup>3</sup> BURCHAT	87 MRK2	$E_{\text{cm}}^{ee} = 29 \text{ GeV}$
87.2 ±0.5 ±0.8		SCHMIDKE	86 MRK2	$E_{\text{cm}}^{ee} = 29 \text{ GeV}$
84.7 ±1.1 $\begin{smallmatrix} +1.6 \\ -1.3 \end{smallmatrix}$	169	<sup>4</sup> ALTHOFF	85 TASS	$E_{\text{cm}}^{ee} = 34.5 \text{ GeV}$
86.1 ±0.5 ±0.9		BARTEL	85F JADE	$E_{\text{cm}}^{ee} = 34.6 \text{ GeV}$
87.8 ±1.3 ±3.9		<sup>5</sup> BERGER	85 PLUT	$E_{\text{cm}}^{ee} = 34.6 \text{ GeV}$
86.7 ±0.3 ±0.6		FERNANDEZ	85 MAC	$E_{\text{cm}}^{ee} = 29 \text{ GeV}$

<sup>1</sup> Not independent of ADEVA 91F  $\Gamma(h^- h^- h^+ \geq 0 \text{ neutrals} \geq 0K_L^0 \nu_\tau) / \Gamma_{\text{total}}$  value.

<sup>2</sup> Not independent of AIHARA 87B  $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau) / \Gamma_{\text{total}}$ ,  $\Gamma(e^- \bar{\nu}_e \nu_\tau) / \Gamma_{\text{total}}$ , and  $\Gamma(h^- \geq 0 \text{ neutrals} \geq 0K_L^0 \nu_\tau) / \Gamma_{\text{total}}$  values.

<sup>3</sup> Not independent of SCHMIDKE 86 value (also not independent of BURCHAT 87 value for  $\Gamma(h^- h^- h^+ \geq 0 \text{ neutrals} \geq 0 K_L^0 \nu_\tau)/\Gamma_{\text{total}}$ ).

<sup>4</sup> Not independent of ALTHOFF 85  $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau)/\Gamma_{\text{total}}$ ,  $\Gamma(e^- \bar{\nu}_e \nu_\tau)/\Gamma_{\text{total}}$ ,  $\Gamma(h^- \geq 0 \text{ neutrals} \geq 0 K_L^0 \nu_\tau)/\Gamma_{\text{total}}$ , and  $\Gamma(h^- h^- h^+ \geq 0 \text{ neutrals} \geq 0 K_L^0 \nu_\tau)/\Gamma_{\text{total}}$  values.

<sup>5</sup> Not independent of (1-prong +  $0\pi^0$ ) and (1-prong +  $\geq 1\pi^0$ ) values.

$\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau)/\Gamma_{\text{total}}$

$\Gamma_3/\Gamma$

To minimize the effect of experiments with large systematic errors, we exclude experiments which together would contribute 5% of the weight in the average.

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>17.39 ± 0.04</b>		<b>OUR FIT</b>		(Produced by HFLAV)
<b>17.33 ± 0.05</b>		<b>OUR AVERAGE</b>		
17.319 ± 0.070 ± 0.032	54k	<sup>1</sup> SCHAEL	05C ALEP	1991-1995 LEP runs
17.34 ± 0.09 ± 0.06	31.4k	ABBIENDI	03 OPAL	1990-1995 LEP runs
17.342 ± 0.110 ± 0.067	21.5k	<sup>2</sup> ACCIARRI	01F L3	1991-1995 LEP runs
17.325 ± 0.095 ± 0.077	27.7k	ABREU	99X DLPH	1991-1995 LEP runs
● ● ● We use the following data for averages but not for fits. ● ● ●				
17.37 ± 0.08 ± 0.18		<sup>3</sup> ANASTASSOV 97	CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
17.31 ± 0.11 ± 0.05	20.7k	BUSKULIC	96C ALEP	Repl. by SCHAEL 05C
17.02 ± 0.19 ± 0.24	6586	ABREU	95T DLPH	Repl. by ABREU 99X
17.36 ± 0.27	7941	AKERS	95I OPAL	Repl. by ABBIENDI 03
17.6 ± 0.4 ± 0.4	2148	ADRIANI	93M L3	Repl. by ACCIARRI 01F
17.4 ± 0.3 ± 0.5		<sup>4</sup> ALBRECHT	93G ARG	$E_{\text{cm}}^{ee} = 9.4\text{--}10.6 \text{ GeV}$
17.35 ± 0.41 ± 0.37		DECAMP	92C ALEP	1989-1990 LEP runs
17.7 ± 0.8 ± 0.4	568	BEHREND	90 CELL	$E_{\text{cm}}^{ee} = 35 \text{ GeV}$
17.4 ± 1.0	2197	ADEVA	88 MRKJ	$E_{\text{cm}}^{ee} = 14\text{--}16 \text{ GeV}$
17.7 ± 1.2 ± 0.7		AIHARA	87B TPC	$E_{\text{cm}}^{ee} = 29 \text{ GeV}$
18.3 ± 0.9 ± 0.8		BURCHAT	87 MRK2	$E_{\text{cm}}^{ee} = 29 \text{ GeV}$
18.6 ± 0.8 ± 0.7	558	<sup>5</sup> BARTEL	86D JADE	$E_{\text{cm}}^{ee} = 34.6 \text{ GeV}$
12.9 ± 1.7 $\begin{smallmatrix} +0.7 \\ -0.5 \end{smallmatrix}$		ALTHOFF	85 TASS	$E_{\text{cm}}^{ee} = 34.5 \text{ GeV}$
18.0 ± 0.9 ± 0.5	473	<sup>5</sup> ASH	85B MAC	$E_{\text{cm}}^{ee} = 29 \text{ GeV}$
18.0 ± 1.0 ± 0.6		<sup>6</sup> BALTRUSAITIS 85	MRK3	$E_{\text{cm}}^{ee} = 3.77 \text{ GeV}$
19.4 ± 1.6 ± 1.7	153	BERGER	85 PLUT	$E_{\text{cm}}^{ee} = 34.6 \text{ GeV}$
17.6 ± 2.6 ± 2.1	47	BEHREND	83C CELL	$E_{\text{cm}}^{ee} = 34 \text{ GeV}$
17.8 ± 2.0 ± 1.8		BERGER	81B PLUT	$E_{\text{cm}}^{ee} = 9\text{--}32 \text{ GeV}$

<sup>1</sup> See footnote to SCHAEL 05C  $\Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)/\Gamma_{\text{total}}$  measurement for correlations with other measurements.

<sup>2</sup> The correlation coefficient between this measurement and the ACCIARRI 01F measurement of  $B(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)$  is 0.08.

<sup>3</sup> The correlation coefficients between this measurement and the ANASTASSOV 97 measurements of  $B(e \bar{\nu}_e \nu_\tau)$ ,  $B(\mu \bar{\nu}_\mu \nu_\tau)/B(e \bar{\nu}_e \nu_\tau)$ ,  $B(h^- \nu_\tau)$ , and  $B(h^- \nu_\tau)/B(e \bar{\nu}_e \nu_\tau)$  are 0.50, 0.58, 0.50, and 0.08 respectively.

<sup>4</sup> Not independent of ALBRECHT 92D  $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau)/\Gamma(e^- \bar{\nu}_e \nu_\tau)$  and ALBRECHT 93G  $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau) \times \Gamma(e^- \bar{\nu}_e \nu_\tau)/\Gamma_{\text{total}}^2$  values.

<sup>5</sup> Modified using  $B(e^- \bar{\nu}_e \nu_\tau)/B(\text{"1 prong"})$  and  $B(\text{"1 prong"}) = 0.855$ .

<sup>6</sup> Error correlated with BALTRUSAITIS 85  $e \nu \bar{\nu}$  value.

$\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau \gamma) / \Gamma_{\text{total}}$   $\Gamma_4 / \Gamma$ 

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.367 ± 0.008 OUR AVERAGE</b>				
0.363 ± 0.002 ± 0.015	22k	<sup>1</sup> SHIMIZU	18A BELL	711 fb <sup>-1</sup> $E_{\text{cm}}^{ee} = 10.6$ GeV
0.369 ± 0.003 ± 0.010	16k	<sup>2</sup> LEES	15G BABR	431 fb <sup>-1</sup> $E_{\text{cm}}^{ee} = 10.6$ GeV
0.361 ± 0.016 ± 0.035		<sup>3</sup> BERGFELD	00 CLEO	$E_{\text{cm}}^{ee} = 10.6$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.30 ± 0.04 ± 0.05	116	<sup>4</sup> ALEXANDER	96S OPAL	1991–1994 LEP runs
0.23 ± 0.10	10	<sup>5</sup> WU	90 MRK2	$E_{\text{cm}}^{ee} = 29$ GeV

<sup>1</sup> SHIMIZU 18A impose requirements on detected  $\gamma$ 's corresponding to a  $\tau$ -rest-frame energy cutoff  $E_\gamma^* > 10$  MeV.

<sup>2</sup> LEES 15G impose requirements on detected  $\gamma$ 's corresponding to a  $\tau$ -rest-frame energy cutoff  $E_\gamma^* > 10$  MeV.

<sup>3</sup> BERGFELD 00 impose requirements on detected  $\gamma$ 's corresponding to a  $\tau$ -rest-frame energy cutoff  $E_\gamma^* > 10$  MeV. For  $E_\gamma^* > 20$  MeV, they quote  $(3.04 \pm 0.14 \pm 0.30) \times 10^{-3}$ .

<sup>4</sup> ALEXANDER 96S impose requirements on detected  $\gamma$ 's corresponding to a  $\tau$ -rest-frame energy cutoff  $E_\gamma > 20$  MeV.

<sup>5</sup> WU 90 reports  $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau \gamma) / \Gamma(\mu^- \bar{\nu}_\mu \nu_\tau) = 0.013 \pm 0.006$ , which is converted to  $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau \gamma) / \Gamma_{\text{total}}$  using  $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau \gamma) / \Gamma_{\text{total}} = 17.35\%$ . Requirements on detected  $\gamma$ 's correspond to a  $\tau$  rest frame energy cutoff  $E_\gamma > 37$  MeV.

 $\Gamma(e^- \bar{\nu}_e \nu_\tau) / \Gamma_{\text{total}}$   $\Gamma_5 / \Gamma$ 

To minimize the effect of experiments with large systematic errors, we exclude experiments which together would contribute 5% of the weight in the average.

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>17.82 ± 0.04 OUR FIT</b> (Produced by HFLAV)				
<b>17.82 ± 0.05 OUR AVERAGE</b>				
17.837 ± 0.072 ± 0.036	56k	<sup>1</sup> SCHAEEL	05C ALEP	1991–1995 LEP runs
17.806 ± 0.104 ± 0.076	24.7k	<sup>2</sup> ACCIARRI	01F L3	1991–1995 LEP runs
17.81 ± 0.09 ± 0.06	33.1k	ABBIENDI	99H OPAL	1991–1995 LEP runs
17.877 ± 0.109 ± 0.110	23.3k	ABREU	99X DLPH	1991–1995 LEP runs
17.76 ± 0.06 ± 0.17		<sup>3</sup> ANASTASSOV	97 CLEO	$E_{\text{cm}}^{ee} = 10.6$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
17.78 ± 0.10 ± 0.09	25.3k	ALEXANDER	96D OPAL	Repl. by ABBI- ENDI 99H
17.79 ± 0.12 ± 0.06	20.6k	BUSKULIC	96C ALEP	Repl. by SCHAEEL 05C
17.51 ± 0.23 ± 0.31	5059	ABREU	95T DLPH	Repl. by ABREU 99X
17.9 ± 0.4 ± 0.4	2892	ADRIANI	93M L3	Repl. by ACCIARRI 01F
17.5 ± 0.3 ± 0.5		<sup>4</sup> ALBRECHT	93G ARG	$E_{\text{cm}}^{ee} = 9.4$ – $10.6$ GeV
17.97 ± 0.14 ± 0.23	3970	AKERIB	92 CLEO	Repl. by ANAS- TASSOV 97
19.1 ± 0.4 ± 0.6	2960	<sup>5</sup> AMMAR	92 CLEO	$E_{\text{cm}}^{ee} = 10.5$ – $10.9$ GeV
18.09 ± 0.45 ± 0.45		DECAMP	92C ALEP	Repl. by SCHAEEL 05C
17.0 ± 0.5 ± 0.6	1.7k	ABACHI	90 HRS	$E_{\text{cm}}^{ee} = 29$ GeV
18.4 ± 0.8 ± 0.4	644	BEHREND	90 CELL	$E_{\text{cm}}^{ee} = 35$ GeV

16.3	$\pm 0.3$	$\pm 3.2$		JANSSEN	89	CBAL	$E_{\text{cm}}^{ee} = 9.4\text{--}10.6$ GeV
18.4	$\pm 1.2$	$\pm 1.0$		AIHARA	87B	TPC	$E_{\text{cm}}^{ee} = 29$ GeV
19.1	$\pm 0.8$	$\pm 1.1$		BURCHAT	87	MRK2	$E_{\text{cm}}^{ee} = 29$ GeV
16.8	$\pm 0.7$	$\pm 0.9$	515	<sup>5</sup> BARTEL	86D	JADE	$E_{\text{cm}}^{ee} = 34.6$ GeV
20.4	$\pm 3.0$	$+1.4$ $-0.9$		ALTHOFF	85	TASS	$E_{\text{cm}}^{ee} = 34.5$ GeV
17.8	$\pm 0.9$	$\pm 0.6$	390	<sup>5</sup> ASH	85B	MAC	$E_{\text{cm}}^{ee} = 29$ GeV
18.2	$\pm 0.7$	$\pm 0.5$		<sup>6</sup> BALTRUSAIT..	85	MRK3	$E_{\text{cm}}^{ee} = 3.77$ GeV
13.0	$\pm 1.9$	$\pm 2.9$		BERGER	85	PLUT	$E_{\text{cm}}^{ee} = 34.6$ GeV
18.3	$\pm 2.4$	$\pm 1.9$	60	BEHREND	83C	CELL	$E_{\text{cm}}^{ee} = 34$ GeV
16.0	$\pm 1.3$		459	<sup>7</sup> BACINO	78B	DLCO	$E_{\text{cm}}^{ee} = 3.1\text{--}7.4$ GeV

<sup>1</sup> Correlation matrix for SCHAEEL 05C branching fractions, in percent:

- (1)  $\Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau) / \Gamma_{\text{total}}$
- (2)  $\Gamma(\tau^- \rightarrow \mu^- \bar{\nu}_\mu \nu_\tau) / \Gamma_{\text{total}}$
- (3)  $\Gamma(\tau^- \rightarrow \pi^- \nu_\tau) / \Gamma_{\text{total}}$
- (4)  $\Gamma(\tau^- \rightarrow \pi^- \pi^0 \nu_\tau) / \Gamma_{\text{total}}$
- (5)  $\Gamma(\tau^- \rightarrow \pi^- 2\pi^0 \nu_\tau (\text{ex. } K^0)) / \Gamma_{\text{total}}$
- (6)  $\Gamma(\tau^- \rightarrow \pi^- 3\pi^0 \nu_\tau (\text{ex. } K^0)) / \Gamma_{\text{total}}$
- (7)  $\Gamma(\tau^- \rightarrow h^- 4\pi^0 \nu_\tau (\text{ex. } K^0, \eta)) / \Gamma_{\text{total}}$
- (8)  $\Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0, \omega)) / \Gamma_{\text{total}}$
- (9)  $\Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \pi^0 \nu_\tau (\text{ex. } K^0)) / \Gamma_{\text{total}}$
- (10)  $\Gamma(\tau^- \rightarrow h^- h^- h^+ 2\pi^0 \nu_\tau (\text{ex. } K^0)) / \Gamma_{\text{total}}$
- (11)  $\Gamma(\tau^- \rightarrow h^- h^- h^+ 3\pi^0 \nu_\tau) / \Gamma_{\text{total}}$
- (12)  $\Gamma(\tau^- \rightarrow 3h^- 2h^+ \nu_\tau (\text{ex. } K^0)) / \Gamma_{\text{total}}$
- (13)  $\Gamma(\tau^- \rightarrow 3h^- 2h^+ \pi^0 \nu_\tau (\text{ex. } K^0)) / \Gamma_{\text{total}}$

	(1)	(2)	(3)	(4)	(5)	(6)	(7)	(8)	(9)	(10)	(11)	(12)	
(2)		-20											
(3)		-9	-6										
(4)		-16	-12	2									
(5)		-5	-5	-17	-37								
(6)		0	-4	-15	2	-27							
(7)		-2	-4	-24	-15	20	-47						
(8)		-14	-9	15	-5	-17	-14	-8					
(9)		-13	-12	-25	-30	4	-2	16	-15				
(10)		0	-2	-23	-14	4	10	13	-6	-17			
(11)		1	0	-5	1	4	6	0	-9	-2	-11		
(12)		0	1	9	4	-8	-4	-6	9	-5	-4	-2	
(13)		1	-4	-3	-5	3	2	-4	-3	-1	4	1	-24

<sup>2</sup> The correlation coefficient between this measurement and the ACCIARRI 01F measurement of  $B(\tau^- \rightarrow \mu^- \bar{\nu}_\mu \nu_\tau)$  is 0.08.

<sup>3</sup> The correlation coefficients between this measurement and the ANASTASSOV 97 measurements of  $B(\mu^- \bar{\nu}_\mu \nu_\tau)$ ,  $B(\mu^- \bar{\nu}_\mu \nu_\tau) / B(e^- \bar{\nu}_e \nu_\tau)$ ,  $B(h^- \nu_\tau)$ , and  $B(h^- \nu_\tau) / B(e^- \bar{\nu}_e \nu_\tau)$  are 0.50,  $-0.42$ ,  $0.48$ , and  $-0.39$  respectively.

<sup>4</sup> Not independent of ALBRECHT 92D  $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau) / \Gamma(e^- \bar{\nu}_e \nu_\tau)$  and ALBRECHT 93G  $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau) \times \Gamma(e^- \bar{\nu}_e \nu_\tau) / \Gamma_{\text{total}}^2$  values.

<sup>5</sup> Modified using  $B(e^- \bar{\nu}_e \nu_\tau) / B(\text{"1 prong"})$  and  $B(\text{"1 prong"}) = 0.855$ .

<sup>6</sup> Error correlated with BALTRUSAITIS 85  $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau) / \Gamma_{\text{total}}$ .

<sup>7</sup> BACINO 78B value comes from fit to events with  $e^\pm$  and one other nonelectron charged prong.

### $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau) / \Gamma(e^- \bar{\nu}_e \nu_\tau)$

$\Gamma_3 / \Gamma_5$

Standard Model prediction including mass effects is 0.9726.

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>97.62 ± 0.28 OUR FIT</b>		(Produced by HFLAV)		
<b>97.9 ± 0.4 OUR AVERAGE</b>				
97.96 ± 0.16 ± 0.36	731k	<sup>1</sup> AUBERT	10F	BABR 467 fb <sup>-1</sup> $E_{\text{cm}}^{ee} = 10.6$ GeV
97.77 ± 0.63 ± 0.87		<sup>2</sup> ANASTASSOV	97	CLEO $E_{\text{cm}}^{ee} = 10.6$ GeV
99.7 ± 3.5 ± 4.0		ALBRECHT	92D	ARG $E_{\text{cm}}^{ee} = 9.4\text{--}10.6$ GeV

<sup>1</sup> Correlation matrix for AUBERT 10F branching fractions:

- (1)  $\Gamma(\tau^- \rightarrow \mu^- \bar{\nu}_\mu \nu_\tau) / \Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)$
- (2)  $\Gamma(\tau^- \rightarrow \pi^- \nu_\tau) / \Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)$
- (3)  $\Gamma(\tau^- \rightarrow K^- \nu_\tau) / \Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)$

(1)	(2)
(2)	0.25
(3)	0.12    0.33

<sup>2</sup> The correlation coefficients between this measurement and the ANASTASSOV 97 measurements of  $B(\mu^- \bar{\nu}_\mu \nu_\tau)$ ,  $B(e^- \bar{\nu}_e \nu_\tau)$ ,  $B(h^- \nu_\tau)$ , and  $B(h^- \nu_\tau) / B(e^- \bar{\nu}_e \nu_\tau)$  are 0.58, -0.42, 0.07, and 0.45 respectively.

### $\Gamma(e^- \bar{\nu}_e \nu_\tau \gamma) / \Gamma_{\text{total}}$

$\Gamma_6 / \Gamma$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.83 ± 0.05 OUR AVERAGE</b>				
1.79 ± 0.02 ± 0.10	12k	<sup>1</sup> SHIMIZU	18A	BELL 711 fb <sup>-1</sup> $E_{\text{cm}}^{ee} = 10.6$ GeV
1.847 ± 0.015 ± 0.052	18k	<sup>2</sup> LEES	15G	BABR 431 fb <sup>-1</sup> $E_{\text{cm}}^{ee} = 10.6$ GeV
1.75 ± 0.06 ± 0.17		<sup>3</sup> BERGFELD	00	CLEO $E_{\text{cm}}^{ee} = 10.6$ GeV

<sup>1</sup> SHIMIZU 18A impose requirements on detected  $\gamma$ 's corresponding to a  $\tau$ -rest-frame energy cutoff  $E_\gamma^* > 10$  MeV.

<sup>2</sup> LEES 15G impose requirements on detected  $\gamma$ 's corresponding to a  $\tau$ -rest-frame energy cutoff  $E_\gamma^* > 10$  MeV.

<sup>3</sup> BERGFELD 00 impose requirements on detected  $\gamma$ 's corresponding to a  $\tau$ -rest-frame energy cutoff  $E_\gamma^* > 10$  MeV.

### $\Gamma(h^- \geq 0K_L^0 \nu_\tau) / \Gamma_{\text{total}}$

$\Gamma_7 / \Gamma$

$$\Gamma_7 / \Gamma = (\Gamma_9 + \Gamma_{10} + \frac{1}{2}\Gamma_{36} + \frac{1}{2}\Gamma_{38} + \Gamma_{50}) / \Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>12.03 ± 0.05 OUR FIT</b>		(Produced by HFLAV)		
<b>12.2 ± 0.4 OUR AVERAGE</b>				
12.47 ± 0.26 ± 0.43	2967	<sup>1</sup> ACCIARRI	95	L3 1992 LEP run
12.4 ± 0.7 ± 0.7	283	<sup>2</sup> ABREU	92N	DLPH 1990 LEP run
12.1 ± 0.7 ± 0.5	309	ALEXANDER	91D	OPAL 1990 LEP run
• • • We use the following data for averages but not for fits. • • •				
11.3 ± 0.5 ± 0.8	798	<sup>3</sup> FORD	87	MAC $E_{\text{cm}}^{ee} = 29$ GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

12.44 ± 0.11 ± 0.11	15k	<sup>4</sup> BUSKULIC	96	ALEP	Repl. by SCHAEEL 05C
11.7 ± 0.6 ± 0.8		<sup>5</sup> ALBRECHT	92D	ARG	$E_{cm}^{ee} = 9.4\text{--}10.6$ GeV
12.98 ± 0.44 ± 0.33		<sup>6</sup> DECAMP	92C	ALEP	Repl. by SCHAEEL 05C
12.3 ± 0.9 ± 0.5	1338	BEHREND	90	CELL	$E_{cm}^{ee} = 35$ GeV
11.1 ± 1.1 ± 1.4		<sup>7</sup> BURCHAT	87	MRK2	$E_{cm}^{ee} = 29$ GeV
12.3 ± 0.6 ± 1.1	328	<sup>8</sup> BARTEL	86D	JADE	$E_{cm}^{ee} = 34.6$ GeV
13.0 ± 2.0 ± 4.0		BERGER	85	PLUT	$E_{cm}^{ee} = 34.6$ GeV
11.2 ± 1.7 ± 1.2	34	<sup>9</sup> BEHREND	83C	CELL	$E_{cm}^{ee} = 34$ GeV

- <sup>1</sup> ACCIARRI 95 with 0.65% added to remove their correction for  $\pi^- K_L^0$  backgrounds.  
<sup>2</sup> ABREU 92N with 0.5% added to remove their correction for  $K^*(892)^-$  backgrounds.  
<sup>3</sup> FORD 87 result for  $B(\pi^- \nu_\tau)$  with 0.67% added to remove their  $K^-$  correction and adjusted for 1992 B("1 prong").  
<sup>4</sup> BUSKULIC 96 quote  $11.78 \pm 0.11 \pm 0.13$  We add 0.66 to undo their correction for unseen  $K_L^0$  and modify the systematic error accordingly.  
<sup>5</sup> Not independent of ALBRECHT 92D  $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau)/\Gamma(e^- \bar{\nu}_e \nu_\tau)$ ,  $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau) \times \Gamma(e^- \bar{\nu}_e \nu_\tau)$ , and  $\Gamma(h^- \geq 0 K_L^0 \nu_\tau)/\Gamma(e^- \bar{\nu}_e \nu_\tau)$  values.  
<sup>6</sup> DECAMP 92C quote  $B(h^- \geq 0 K_L^0 \geq 0 (K_S^0 \rightarrow \pi^+ \pi^-) \nu_\tau) = 13.32 \pm 0.44 \pm 0.33$ . We subtract 0.35 to correct for their inclusion of the  $K_S^0$  decays.  
<sup>7</sup> BURCHAT 87 with 1.1% added to remove their correction for  $K^-$  and  $K^*(892)^-$  backgrounds.  
<sup>8</sup> BARTEL 86D result for  $B(\pi^- \nu_\tau)$  with 0.59% added to remove their  $K^-$  correction and adjusted for 1992 B("1 prong").  
<sup>9</sup> BEHREND 83C quote  $B(\pi^- \nu_\tau) = 9.9 \pm 1.7 \pm 1.3$  after subtracting  $1.3 \pm 0.5$  to correct for  $B(K^- \nu_\tau)$ .

$\Gamma(h^- \nu_\tau)/\Gamma_{total}$   $\Gamma_8/\Gamma = (\Gamma_9 + \Gamma_{10})/\Gamma$

<u>VALUE (%)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
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**11.51 ± 0.05 OUR FIT** (Produced by HFLAV)

**11.63 ± 0.12 OUR AVERAGE** Error includes scale factor of 1.4. See the ideogram below.

11.571 ± 0.120 ± 0.114	19k	<sup>1</sup> ABDALLAH	06A	DLPH	1992–1995 LEP runs
11.98 ± 0.13 ± 0.16		ACKERSTAFF	98M	OPAL	1991–1995 LEP runs
11.52 ± 0.05 ± 0.12		<sup>2</sup> ANASTASSOV	97	CLEO	$E_{cm}^{ee} = 10.6$ GeV

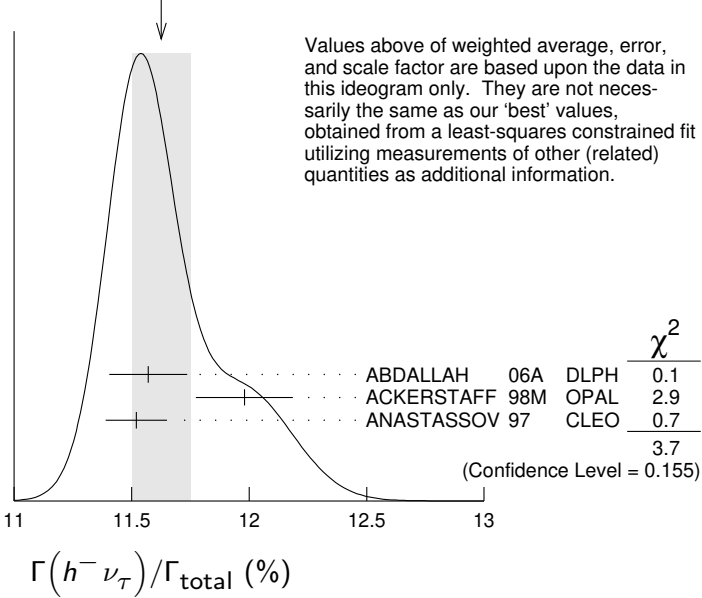
<sup>1</sup> Correlation matrix for ABDALLAH 06A branching fractions, in percent:

- (1)  $\Gamma(\tau^- \rightarrow h^- \nu_\tau)/\Gamma_{total}$
- (2)  $\Gamma(\tau^- \rightarrow h^- \pi^0 \nu_\tau)/\Gamma_{total}$
- (3)  $\Gamma(\tau^- \rightarrow h^- \geq 1 \pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{total}$
- (4)  $\Gamma(\tau^- \rightarrow h^- 2 \pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{total}$
- (5)  $\Gamma(\tau^- \rightarrow h^- \geq 3 \pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{total}$
- (6)  $\Gamma(\tau^- \rightarrow h^- h^- h^+ \nu_\tau (\text{ex. } K^0))/\Gamma_{total}$
- (7)  $\Gamma(\tau^- \rightarrow h^- h^- h^+ \pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{total}$
- (8)  $\Gamma(\tau^- \rightarrow h^- h^- h^+ \geq 1 \pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{total}$
- (9)  $\Gamma(\tau^- \rightarrow h^- h^- h^+ \geq 2 \pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{total}$
- (10)  $\Gamma(\tau^- \rightarrow 3 h^- 2 h^+ \nu_\tau (\text{ex. } K^0))/\Gamma_{total}$
- (11)  $\Gamma(\tau^- \rightarrow 3 h^- 2 h^+ \pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{total}$

	(1)	(2)	(3)	(4)	(5)	(6)	(7)	(8)	(9)	(10)
(2)	-34									
(3)	-47	56								
(4)	6	-66	15							
(5)	-6	38	11	-86						
(6)	-7	-8	15	0	-2					
(7)	-2	-1	-5	-3	3	-53				
(8)	-4	-4	-13	-4	-2	-56	75			
(9)	-1	-1	-4	3	-6	26	-78	-16		
(10)	-1	-1	1	0	0	-2	-3	-1	3	
(11)	0	0	0	0	0	1	0	-5	5	-57

<sup>2</sup> The correlation coefficients between this measurement and the ANASTASSOV 97 measurements of  $B(\mu\bar{\nu}_\mu\nu_\tau)$ ,  $B(e\bar{\nu}_e\nu_\tau)$ ,  $B(\mu\bar{\nu}_\mu\nu_\tau)/B(e\bar{\nu}_e\nu_\tau)$ , and  $B(h^-\nu_\tau)/B(e\bar{\nu}_e\nu_\tau)$  are 0.50, 0.48, 0.07, and 0.63 respectively.

WEIGHTED AVERAGE  
11.63±0.12 (Error scaled by 1.4)



$\Gamma(h^-\nu_\tau)/\Gamma(e^-\bar{\nu}_e\nu_\tau)$

$\Gamma_8/\Gamma_5 = (\Gamma_9 + \Gamma_{10})/\Gamma_5$

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
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**64.61±0.33 OUR FIT** (Produced by HFLAV)

**64.0 ±0.7 OUR AVERAGE** Error includes scale factor of 1.6.

• • • We use the following data for averages but not for fits. • • •

63.33±0.14±0.61	394k	<sup>1</sup> AUBERT	10F BABR	467 fb <sup>-1</sup> $E_{cm}^{ee}=10.6$ GeV
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64.84±0.41±0.60		<sup>2</sup> ANASTASSOV 97	CLEO	$E_{cm}^{ee}=10.6$ GeV
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<sup>1</sup> Not independent of AUBERT 10F  $\Gamma(\tau^- \rightarrow \pi^- \nu_\tau)/\Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)$  and  $\Gamma(\tau^- \rightarrow K^- \nu_\tau)/\Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)$ .

<sup>2</sup> The correlation coefficients between this measurement and the ANASTASSOV 97 measurements of  $B(\mu\bar{\nu}_\mu\nu_\tau)$ ,  $B(e\bar{\nu}_e\nu_\tau)$ ,  $B(\mu\bar{\nu}_\mu\nu_\tau)/B(e\bar{\nu}_e\nu_\tau)$ , and  $B(h^-\nu_\tau)$  are 0.08, -0.39, 0.45, and 0.63 respectively.



$\Gamma(\pi^- \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_9/\Gamma$ 

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>10.82 ± 0.05 OUR FIT</b>		(Produced by HFLAV)		
<b>10.828 ± 0.070 ± 0.078</b>	38k	<sup>1</sup> SCHAEL	05C ALEP	1991-1995 LEP runs
• • • We do not use the following data for averages, fits, limits, etc. • • •				
11.06 ± 0.11 ± 0.14		<sup>2</sup> BUSKULIC	96 ALEP	Repl. by SCHAEL 05C
11.7 ± 0.4 ± 1.8	1138	BLOCKER	82D MRK2	$E_{\text{cm}}^{ee} = 3.5\text{--}6.7$ GeV

<sup>1</sup> See footnote to SCHAEL 05C  $\Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)/\Gamma_{\text{total}}$  measurement for correlations with other measurements.

<sup>2</sup> Not independent of BUSKULIC 96  $B(h^- \nu_\tau)$  and  $B(K^- \nu_\tau)$  values.

 $\Gamma(\pi^- \nu_\tau)/\Gamma(e^- \bar{\nu}_e \nu_\tau)$   $\Gamma_9/\Gamma_5$ 

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>60.71 ± 0.32 OUR FIT</b>		(Produced by HFLAV)		
<b>59.45 ± 0.14 ± 0.61</b>	369k	<sup>1</sup> AUBERT	10F BABR	$467 \text{ fb}^{-1}$ $E_{\text{cm}}^{ee} = 10.6$ GeV

<sup>1</sup> See footnote to AUBERT 10F  $\Gamma(\tau^- \rightarrow \mu^- \bar{\nu}_\mu \nu_\tau)/\Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)$  for correlations with other measurements.

 $\Gamma(K^- \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{10}/\Gamma$ 

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.696 ± 0.010 OUR FIT</b>		(Produced by HFLAV)		
<b>0.685 ± 0.023 OUR AVERAGE</b>				
0.658 ± 0.027 ± 0.029		<sup>1</sup> ABBIENDI	01J OPAL	1990–1995 LEP runs
0.696 ± 0.025 ± 0.014	2032	BARATE	99K ALEP	1991–1995 LEP runs
0.85 ± 0.18	27	ABREU	94K DLPH	LEP 1992 Z data
0.66 ± 0.07 ± 0.09	99	BATTLE	94 CLEO	$E_{\text{cm}}^{ee} \approx 10.6$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.72 ± 0.04 ± 0.04	728	BUSKULIC	96 ALEP	Repl. by BARATE 99K
0.59 ± 0.18	16	MILLS	84 DLCO	$E_{\text{cm}}^{ee} = 29$ GeV
1.3 ± 0.5	15	BLOCKER	82B MRK2	$E_{\text{cm}}^{ee} = 3.9\text{--}6.7$ GeV

<sup>1</sup> The correlation coefficient between this measurement and the ABBIENDI 01J  $B(\tau^- \rightarrow K^- \geq 0\pi^0 \geq 0K^0 \geq 0\gamma \nu_\tau)$  is 0.60.

 $\Gamma(K^- \nu_\tau)/\Gamma(e^- \bar{\nu}_e \nu_\tau)$   $\Gamma_{10}/\Gamma_5$ 

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>3.91 ± 0.05 OUR FIT</b>		(Produced by HFLAV)		
<b>3.882 ± 0.032 ± 0.057</b>	25k	<sup>1</sup> AUBERT	10F BABR	$467 \text{ fb}^{-1}$ $E_{\text{cm}}^{ee} = 10.6$ GeV

<sup>1</sup> See footnote to AUBERT 10F  $\Gamma(\tau^- \rightarrow \mu^- \bar{\nu}_\mu \nu_\tau)/\Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)$  for correlations with other measurements.

 $\Gamma(K^- \nu_\tau)/\Gamma(\pi^- \nu_\tau)$   $\Gamma_{10}/\Gamma_9$ 

VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT
<b>6.44 ± 0.09 OUR FIT</b>	(Produced by HFLAV)		
• • • We use the following data for averages but not for fits. • • •			
<b>6.531 ± 0.056 ± 0.093</b>	<sup>1</sup> AUBERT	10F BABR	$467 \text{ fb}^{-1}$ $E_{\text{cm}}^{ee} = 10.6$ GeV

<sup>1</sup> Not independent of AUBERT 10F  $\Gamma(\tau^- \rightarrow \pi^- \nu_\tau)/\Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)$  and  $\Gamma(\tau^- \rightarrow K^- \nu_\tau)/\Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)$ .

$$\Gamma(h^- \geq 1 \text{ neutrals } \nu_\tau) / \Gamma_{\text{total}} \quad \Gamma_{11} / \Gamma$$

$$\Gamma_{11} / \Gamma = (\Gamma_{14} + \Gamma_{16} + \Gamma_{20} + \Gamma_{23} + \Gamma_{27} + \Gamma_{28} + \Gamma_{30} + 0.15344\Gamma_{36} + 0.15344\Gamma_{38} + 0.15344\Gamma_{41} + 0.15344\Gamma_{43} + 0.0942\Gamma_{48} + 0.0942\Gamma_{52} + 0.7196\Gamma_{150} + 0.7196\Gamma_{152} + 0.7196\Gamma_{154} + 0.1107\Gamma_{156} + 0.0835\Gamma_{178} + 0.0835\Gamma_{179} + 0.0835\Gamma_{180}) / \Gamma$$

VALUE (%)	DOCUMENT ID	TECN	COMMENT
<b>37.00 ± 0.09 OUR FIT</b>	(Produced by HFLAV)		

• • • We do not use the following data for averages, fits, limits, etc. • • •

36.14 ± 0.33 ± 0.58	<sup>1</sup> AKERS	94E	OPAL 1991–1992 LEP runs
38.4 ± 1.2 ± 1.0	<sup>2</sup> BURCHAT	87	MRK2 $E_{\text{cm}}^{ee} = 29$ GeV
42.7 ± 2.0 ± 2.9	BERGER	85	PLUT $E_{\text{cm}}^{ee} = 34.6$ GeV

<sup>1</sup> Not independent of ACKERSTAFF 98M  $B(h^- \pi^0 \nu_\tau)$  and  $B(h^- \geq 2\pi^0 \nu_\tau)$  values.

<sup>2</sup> BURCHAT 87 quote for  $B(\pi^\pm \geq 1 \text{ neutral } \nu_\tau) = 0.378 \pm 0.012 \pm 0.010$ . We add 0.006 to account for contribution from  $(K^{*-} \nu_\tau)$  which they fixed at BR = 0.013.

$$\Gamma(h^- \geq 1\pi^0 \nu_\tau \text{ (ex. } K^0)) / \Gamma_{\text{total}} \quad \Gamma_{12} / \Gamma$$

$$\Gamma_{12} / \Gamma = (\Gamma_{14} + \Gamma_{16} + \Gamma_{20} + \Gamma_{23} + \Gamma_{27} + \Gamma_{28} + \Gamma_{30} + 0.3257\Gamma_{150} + 0.3257\Gamma_{152} + 0.3257\Gamma_{154}) / \Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>36.50 ± 0.09 OUR FIT</b>	(Produced by HFLAV)			

• • • We use the following data for averages but not for fits. • • •

<b>36.641 ± 0.155 ± 0.127</b>	45k	<sup>1</sup> ABDALLAH	06A	DLPH 1992–1995 LEP runs
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<sup>1</sup> See footnote to ABDALLAH 06A  $\Gamma(\tau^- \rightarrow h^- \nu_\tau) / \Gamma_{\text{total}}$  measurement for correlations with other measurements.

$$\Gamma(h^- \pi^0 \nu_\tau) / \Gamma_{\text{total}} \quad \Gamma_{13} / \Gamma = (\Gamma_{14} + \Gamma_{16}) / \Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>25.93 ± 0.09 OUR FIT</b>	(Produced by HFLAV)			

**25.73 ± 0.16 OUR AVERAGE**

25.67 ± 0.01 ± 0.39	5.4M	FUJIKAWA	08	BELL $72 \text{ fb}^{-1}$ $E_{\text{cm}}^{ee} = 10.6$ GeV
25.740 ± 0.201 ± 0.138	35k	<sup>1</sup> ABDALLAH	06A	DLPH 1992–1995 LEP runs
25.89 ± 0.17 ± 0.29		ACKERSTAFF	98M	OPAL 1991–1995 LEP runs
25.05 ± 0.35 ± 0.50	6613	ACCIARRI	95	L3 1992 LEP run
25.87 ± 0.12 ± 0.42	51k	<sup>2</sup> ARTUSO	94	CLEO $E_{\text{cm}}^{ee} = 10.6$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
25.76 ± 0.15 ± 0.13	31k	BUSKULIC	96	ALEP Repl. by SCHAEEL 05C
25.98 ± 0.36 ± 0.52		<sup>3</sup> AKERS	94E	OPAL Repl. by ACKER-STAFF 98M
22.9 ± 0.8 ± 1.3	283	<sup>4</sup> ABREU	92N	DLPH $E_{\text{cm}}^{ee} = 88.2\text{--}94.2$ GeV
23.1 ± 0.4 ± 0.9	1249	<sup>5</sup> ALBRECHT	92Q	ARG $E_{\text{cm}}^{ee} = 10$ GeV
25.02 ± 0.64 ± 0.88	1849	DECAMP	92C	ALEP 1989–1990 LEP runs
22.0 ± 0.8 ± 1.9	779	ANTREASYAN	91	CBAL $E_{\text{cm}}^{ee} = 9.4\text{--}10.6$ GeV
22.6 ± 1.5 ± 0.7	1101	BEHREND	90	CELL $E_{\text{cm}}^{ee} = 35$ GeV
23.1 ± 1.9 ± 1.6		BEHREND	84	CELL $E_{\text{cm}}^{ee} = 14,22$ GeV

<sup>1</sup> See footnote to ABDALLAH 06A  $\Gamma(\tau^- \rightarrow h^- \nu_\tau) / \Gamma_{\text{total}}$  measurement for correlations with other measurements.

<sup>2</sup> ARTUSO 94 reports the combined result from three independent methods, one of which (23% of the  $\tau^- \rightarrow h^- \pi^0 \nu_\tau$ ) is normalized to the inclusive one-prong branching fraction, taken as  $0.854 \pm 0.004$ . Renormalization to the present value causes negligible change.

<sup>3</sup> AKERS 94E quote  $(26.25 \pm 0.36 \pm 0.52) \times 10^{-2}$ ; we subtract 0.27% from their number to correct for  $\tau^- \rightarrow h^- K_L^0 \nu_\tau$ .

<sup>4</sup> ABREU 92N with 0.5% added to remove their correction for  $K^*(892)^-$  backgrounds.

<sup>5</sup> ALBRECHT 92Q with 0.5% added to remove their correction for  $\tau^- \rightarrow K^*(892)^- \nu_\tau$  background.

$\Gamma(\pi^- \pi^0 \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{14}/\Gamma$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>25.49 ± 0.09 OUR FIT</b>				(Produced by HFLAV)
<b>25.46 ± 0.12 OUR AVERAGE</b>				
25.471 ± 0.097 ± 0.085	81k	<sup>1</sup> SCHAEL	05C ALEP	1991-1995 LEP runs
• • • We use the following data for averages but not for fits. • • •				
25.36 ± 0.44		<sup>2</sup> ARTUSO	94 CLEO	$E_{\text{cm}}^{ee} = 10.6$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
25.30 ± 0.15 ± 0.13		<sup>3</sup> BUSKULIC	96 ALEP	Repl. by SCHAEL 05C
21.5 ± 0.4 ± 1.9	4400	<sup>4,5</sup> ALBRECHT	88L ARG	$E_{\text{cm}}^{ee} = 10$ GeV
23.0 ± 1.3 ± 1.7	582	ADLER	87B MRK3	$E_{\text{cm}}^{ee} = 3.77$ GeV
25.8 ± 1.7 ± 2.5		<sup>6</sup> BURCHAT	87 MRK2	$E_{\text{cm}}^{ee} = 29$ GeV
22.3 ± 0.6 ± 1.4	629	<sup>5</sup> YELTON	86 MRK2	$E_{\text{cm}}^{ee} = 29$ GeV

<sup>1</sup> See footnote to SCHAEL 05C  $\Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)/\Gamma_{\text{total}}$  measurement for correlations with other measurements.

<sup>2</sup> Not independent of ARTUSO 94  $B(h^- \pi^0 \nu_\tau)$  and BATTLE 94  $B(K^- \pi^0 \nu_\tau)$  values.

<sup>3</sup> Not independent of BUSKULIC 96  $B(h^- \pi^0 \nu_\tau)$  and  $B(K^- \pi^0 \nu_\tau)$  values.

<sup>4</sup> The authors divide by  $(\Gamma_3 + \Gamma_5 + \Gamma_9 + \Gamma_{10})/\Gamma = 0.467$  to obtain this result.

<sup>5</sup> Experiment had no hadron identification. Kaon corrections were made, but insufficient information is given to permit their removal.

<sup>6</sup> BURCHAT 87 value is not independent of YELTON 86 value. Nonresonant decays included.

$\Gamma(\pi^- \pi^0 \text{non-}\rho(770) \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{15}/\Gamma$

VALUE (%)	DOCUMENT ID	TECN	COMMENT
<b>0.3 ± 0.1 ± 0.3</b>	<sup>1</sup> BEHREND	84 CELL	$E_{\text{cm}}^{ee} = 14.22$ GeV

<sup>1</sup> BEHREND 84 assume a flat nonresonant mass distribution down to the  $\rho(770)$  mass, using events with mass above 1300 to set the level.

$\Gamma(K^- \pi^0 \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{16}/\Gamma$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.433 ± 0.015 OUR FIT</b>				(Produced by HFLAV)
<b>0.426 ± 0.016 OUR AVERAGE</b>				
0.416 ± 0.003 ± 0.018	78k	AUBERT	07AP BABR	230 fb <sup>-1</sup> $E_{\text{cm}}^{ee} = 10.6$ GeV
0.471 ± 0.059 ± 0.023	360	ABBIENDI	04J OPAL	1991-1995 LEP runs
0.444 ± 0.026 ± 0.024	923	BARATE	99K ALEP	1991-1995 LEP runs
0.51 ± 0.10 ± 0.07	37	BATTLE	94 CLEO	$E_{\text{cm}}^{ee} \approx 10.6$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.52 ± 0.04 ± 0.05	395	BUSKULIC	96 ALEP	Repl. by BARATE 99K

$$\Gamma(h^- \geq 2\pi^0 \nu_\tau) / \Gamma_{\text{total}} \quad \Gamma_{17} / \Gamma$$

$$\Gamma_{17} / \Gamma = (\Gamma_{20} + \Gamma_{23} + \Gamma_{27} + \Gamma_{28} + \Gamma_{30} + 0.15344\Gamma_{36} + 0.15344\Gamma_{38} + 0.15344\Gamma_{41} + 0.15344\Gamma_{43} + 0.09419\Gamma_{48} + 0.0942\Gamma_{52} + 0.3257\Gamma_{150} + 0.3257\Gamma_{152} + 0.3257\Gamma_{154}) / \Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
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**10.81 ± 0.09 OUR FIT** (Produced by HFLAV)

**9.91 ± 0.31 ± 0.27** ACKERSTAFF 98M OPAL 1991–1995 LEP runs

• • • We do not use the following data for averages, fits, limits, etc. • • •

9.89 ± 0.34 ± 0.55		<sup>1</sup> AKERS	94E	OPAL Repl. by ACKER-STAFF 98M
14.0 ± 1.2 ± 0.6	938	<sup>2</sup> BEHREND	90	CELL $E_{\text{cm}}^{ee} = 35$ GeV
12.0 ± 1.4 ± 2.5		<sup>3</sup> BURCHAT	87	MRK2 $E_{\text{cm}}^{ee} = 29$ GeV
13.9 ± 2.0 $\begin{smallmatrix} +1.9 \\ -2.2 \end{smallmatrix}$		<sup>4</sup> AIHARA	86E	TPC $E_{\text{cm}}^{ee} = 29$ GeV

<sup>1</sup> AKERS 94E not independent of AKERS 94E  $B(h^- \geq 1\pi^0 \nu_\tau)$  and  $B(h^- \pi^0 \nu_\tau)$  measurements.

<sup>2</sup> No independent of BEHREND 90  $\Gamma(h^- 2\pi^0 \nu_\tau (\text{exp. } K^0))$  and  $\Gamma(h^- \geq 3\pi^0 \nu_\tau)$ .

<sup>3</sup> Error correlated with BURCHAT 87  $\Gamma(\rho^- \nu_e) / \Gamma(\text{total})$  value.

<sup>4</sup> AIHARA 86E (TPC) quote  $B(2\pi^0 \pi^- \nu_\tau) + 1.6B(3\pi^0 \pi^- \nu_\tau) + 1.1B(\pi^0 \eta \pi^- \nu_\tau)$ .

$$\Gamma(h^- 2\pi^0 \nu_\tau) / \Gamma_{\text{total}} \quad \Gamma_{18} / \Gamma$$

$$\Gamma_{18} / \Gamma = (\Gamma_{20} + \Gamma_{23} + 0.15344\Gamma_{36} + 0.15344\Gamma_{38}) / \Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
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**9.48 ± 0.10 OUR FIT** (Produced by HFLAV)

• • • We do not use the following data for averages, fits, limits, etc. • • •

9.48 ± 0.13 ± 0.10 12k <sup>1</sup> BUSKULIC 96 ALEP Repl. by SCHAELE 05C

<sup>1</sup> BUSKULIC 96 quote  $9.29 \pm 0.13 \pm 0.10$ . We add 0.19 to undo their correction for  $\tau^- \rightarrow h^- K^0 \nu_\tau$ .

$$\Gamma(h^- 2\pi^0 \nu_\tau (\text{ex. } K^0)) / \Gamma_{\text{total}} \quad \Gamma_{19} / \Gamma$$

$$\Gamma_{19} / \Gamma = (\Gamma_{20} + \Gamma_{23}) / \Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
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**9.32 ± 0.10 OUR FIT** (Produced by HFLAV)

**9.17 ± 0.27 OUR AVERAGE**

9.498 ± 0.320 ± 0.275 9.5k <sup>1</sup> ABDALLAH 06A DLPH 1992–1995 LEP runs

8.88 ± 0.37 ± 0.42 1060 ACCIARRI 95 L3 1992 LEP run

• • • We use the following data for averages but not for fits. • • •

8.96 ± 0.16 ± 0.44 <sup>2</sup> PROCARIO 93 CLEO  $E_{\text{cm}}^{ee} \approx 10.6$  GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

10.38 ± 0.66 ± 0.82 809 <sup>3</sup> DECAMP 92C ALEP Repl. by SCHAELE 05C

5.7 ± 0.5  $\begin{smallmatrix} +1.7 \\ -1.0 \end{smallmatrix}$  133 <sup>4</sup> ANTREASYAN 91 CBAL  $E_{\text{cm}}^{ee} = 9.4\text{--}10.6$  GeV

10.0 ± 1.5 ± 1.1 333 <sup>5</sup> BEHREND 90 CELL  $E_{\text{cm}}^{ee} = 35$  GeV

8.7 ± 0.4 ± 1.1 815 <sup>6</sup> BAND 87 MAC  $E_{\text{cm}}^{ee} = 29$  GeV

6.2 ± 0.6 ± 1.2 <sup>7</sup> GAN 87 MRK2  $E_{\text{cm}}^{ee} = 29$  GeV

6.0 ± 3.0 ± 1.8 BEHREND 84 CELL  $E_{\text{cm}}^{ee} = 14,22$  GeV

<sup>1</sup> See footnote to ABDALLAH 06A  $\Gamma(\tau^- \rightarrow h^- \nu_\tau) / \Gamma_{\text{total}}$  measurement for correlations with other measurements.

<sup>2</sup> PROCARIO 93 entry is obtained from  $B(h^- 2\pi^0 \nu_\tau) / B(h^- \pi^0 \nu_\tau)$  using ARTUSO 94 result for  $B(h^- \pi^0 \nu_\tau)$ .

<sup>3</sup> We subtract 0.0015 to account for  $\tau^- \rightarrow K^*(892)^- \nu_\tau$  contribution.

<sup>4</sup> ANTREASYAN 91 subtract 0.001 to account for the  $\tau^- \rightarrow K^*(892)^- \nu_\tau$  contribution.

<sup>5</sup> BEHREND 90 subtract 0.002 to account for the  $\tau^- \rightarrow K^*(892)^- \nu_\tau$  contribution.

<sup>6</sup> BAND 87 assume  $B(\pi^- 3\pi^0 \nu_\tau) = 0.01$  and  $B(\pi^- \pi^0 \eta \nu_\tau) = 0.005$ .

<sup>7</sup> GAN 87 analysis use photon multiplicity distribution.

### $\Gamma(h^- 2\pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma(h^- \pi^0 \nu_\tau)$ $\Gamma_{19}/\Gamma_{13}$

$\Gamma_{19}/\Gamma_{13} = (\Gamma_{20} + \Gamma_{23})/(\Gamma_{14} + \Gamma_{16})$

VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT
<b>36.0 ± 0.4 OUR FIT</b>	(Produced by HFLAV)		
<b>34.2 ± 0.6 ± 1.6</b>	<sup>1</sup> PROCARIO	93	CLEO $E_{\text{cm}}^{ee} \approx 10.6$ GeV

<sup>1</sup> PROCARIO 93 quote  $0.345 \pm 0.006 \pm 0.016$  after correction for 2 kaon backgrounds assuming  $B(K^{*-} \nu_\tau) = 1.42 \pm 0.18\%$  and  $B(h^- K^0 \pi^0 \nu_\tau) = 0.48 \pm 0.48\%$ . We multiply by  $0.990 \pm 0.010$  to remove these corrections to  $B(h^- \pi^0 \nu_\tau)$ .

### $\Gamma(\pi^- 2\pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$ $\Gamma_{20}/\Gamma$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>9.26 ± 0.10 OUR FIT</b>	(Produced by HFLAV)			
<b>9.239 ± 0.086 ± 0.090</b>	31k	<sup>1</sup> SCHAEL	05C	ALEP 1991-1995 LEP runs
• • • We do not use the following data for averages, fits, limits, etc. • • •				

9.21 ± 0.13 ± 0.11 <sup>2</sup> BUSKULIC 96 ALEP Repl. by SCHAEL 05C

<sup>1</sup> See footnote to SCHAEL 05C  $\Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)/\Gamma_{\text{total}}$  measurement for correlations with other measurements.

<sup>2</sup> Not independent of BUSKULIC 96  $B(h^- 2\pi^0 \nu_\tau (\text{ex. } K^0))$  and  $B(K^- 2\pi^0 \nu_\tau (\text{ex. } K^0))$  values.

### $\Gamma(\pi^- 2\pi^0 \nu_\tau (\text{ex. } K^0), \text{scalar})/\Gamma(\pi^- 2\pi^0 \nu_\tau (\text{ex. } K^0))$ $\Gamma_{21}/\Gamma_{20}$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.094</b>	95	<sup>1</sup> BROWDER	00	CLEO $4.7 \text{ fb}^{-1}$ $E_{\text{cm}}^{ee} = 10.6$ GeV

<sup>1</sup> Model-independent limit from structure function analysis on contribution to  $B(\tau^- \rightarrow \pi^- 2\pi^0 \nu_\tau (\text{ex. } K^0))$  from scalars.

### $\Gamma(\pi^- 2\pi^0 \nu_\tau (\text{ex. } K^0), \text{vector})/\Gamma(\pi^- 2\pi^0 \nu_\tau (\text{ex. } K^0))$ $\Gamma_{22}/\Gamma_{20}$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.073</b>	95	<sup>1</sup> BROWDER	00	CLEO $4.7 \text{ fb}^{-1}$ $E_{\text{cm}}^{ee} = 10.6$ GeV

<sup>1</sup> Model-independent limit from structure function analysis on contribution to  $B(\tau^- \rightarrow \pi^- 2\pi^0 \nu_\tau (\text{ex. } K^0))$  from vectors.

### $\Gamma(K^- 2\pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$ $\Gamma_{23}/\Gamma$

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>6.5 ± 2.2 OUR FIT</b>	(Produced by HFLAV)			
<b>5.8 ± 2.4 OUR AVERAGE</b>				

5.6 ± 2.0 ± 1.5 131 BARATE 99K ALEP 1991–1995 LEP runs

9 ± 10 ± 3 3 <sup>1</sup> BATTLE 94 CLEO  $E_{\text{cm}}^{ee} \approx 10.6$  GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

8 ± 2 ± 2 59 BUSKULIC 96 ALEP Repl. by BARATE 99K

<sup>1</sup> BATTLE 94 quote  $(14 \pm 10 \pm 3) \times 10^{-4}$  or  $< 30 \times 10^{-4}$  at 90% CL. We subtract  $(5 \pm 2) \times 10^{-4}$  to account for  $\tau^- \rightarrow K^- (K^0 \rightarrow \pi^0 \pi^0) \nu_\tau$  background.

$$\Gamma(h^- \geq 3\pi^0 \nu_\tau) / \Gamma_{\text{total}} \quad \Gamma_{24} / \Gamma$$

$$\Gamma_{24} / \Gamma = (\Gamma_{27} + \Gamma_{28} + \Gamma_{30} + 0.15344\Gamma_{41} + 0.15344\Gamma_{43} + 0.0942\Gamma_{48} + 0.0942\Gamma_{52} + 0.3257\Gamma_{150} + 0.3257\Gamma_{152} + 0.3257\Gamma_{154} + 0.0501\Gamma_{156}) / \Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.34 ± 0.07 OUR FIT</b>	(Produced by HFLAV)			

• • • We do not use the following data for averages, fits, limits, etc. • • •

1.53 ± 0.40 ± 0.46	186	DECAMP	92C	ALEP Repl. by SCHAEEL 05C
3.2 ± 1.0 ± 1.0		BEHREND	90	CELL $E_{\text{cm}}^{\text{ee}} = 35$ GeV

$$\Gamma(h^- \geq 3\pi^0 \nu_\tau (\text{ex. } K^0)) / \Gamma_{\text{total}} \quad \Gamma_{25} / \Gamma$$

$$\Gamma_{25} / \Gamma = (\Gamma_{27} + \Gamma_{28} + \Gamma_{30} + 0.3257\Gamma_{150} + 0.3257\Gamma_{152} + 0.3257\Gamma_{154}) / \Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.25 ± 0.07 OUR FIT</b>	(Produced by HFLAV)			

**1.403 ± 0.214 ± 0.224** 1.1k <sup>1</sup> ABDALLAH 06A DLPH 1992–1995 LEP runs

<sup>1</sup> See footnote to ABDALLAH 06A  $\Gamma(\tau^- \rightarrow h^- \nu_\tau) / \Gamma_{\text{total}}$  measurement for correlations with other measurements.

$$\Gamma(h^- 3\pi^0 \nu_\tau) / \Gamma_{\text{total}} \quad \Gamma_{26} / \Gamma$$

$$\Gamma_{26} / \Gamma = (\Gamma_{27} + \Gamma_{28} + 0.15344\Gamma_{41} + 0.15344\Gamma_{43} + 0.3257\Gamma_{152}) / \Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.18 ± 0.07 OUR FIT</b>	(Produced by HFLAV)			

**1.21 ± 0.17 OUR AVERAGE** Error includes scale factor of 1.2.

1.70 ± 0.24 ± 0.38	293	ACCIARRI	95	L3 1992 LEP run
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• • • We use the following data for averages but not for fits. • • •

1.15 ± 0.08 ± 0.13		<sup>1</sup> PROCARIO	93	CLEO $E_{\text{cm}}^{\text{ee}} \approx 10.6$ GeV
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• • • We do not use the following data for averages, fits, limits, etc. • • •

1.24 ± 0.09 ± 0.11	2.3k	<sup>2</sup> BUSKULIC	96	ALEP Repl. by SCHAEEL 05C
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0.0 <sup>+1.4</sup> <sub>-0.1</sub> <sup>+1.1</sup> <sub>-0.1</sub>		<sup>3</sup> GAN	87	MRK2 $E_{\text{cm}}^{\text{ee}} = 29$ GeV
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<sup>1</sup> PROCARIO 93 entry is obtained from  $B(h^- 3\pi^0 \nu_\tau) / B(h^- \pi^0 \nu_\tau)$  using ARTUSO 94 result for  $B(h^- \pi^0 \nu_\tau)$ .

<sup>2</sup> BUSKULIC 96 quote  $B(h^- 3\pi^0 \nu_\tau (\text{ex. } K^0)) = 1.17 \pm 0.09 \pm 0.11$ . We add 0.07 to remove their correction for  $K^0$  backgrounds.

<sup>3</sup> Highly correlated with GAN 87  $\Gamma(\eta \pi^- \pi^0 \nu_\tau) / \Gamma_{\text{total}}$  value. Authors quote  $B(\pi^\pm 3\pi^0 \nu_\tau) + 0.67B(\pi^\pm \eta \pi^0 \nu_\tau) = 0.047 \pm 0.010 \pm 0.011$ .

$$\Gamma(h^- 3\pi^0 \nu_\tau) / \Gamma(h^- \pi^0 \nu_\tau) \quad \Gamma_{26} / \Gamma_{13}$$

$$\Gamma_{26} / \Gamma_{13} = (\Gamma_{27} + \Gamma_{28} + 0.15344\Gamma_{41} + 0.15344\Gamma_{43} + 0.3257\Gamma_{152}) / (\Gamma_{14} + \Gamma_{16})$$

VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT
<b>4.54 ± 0.28 OUR FIT</b>	(Produced by HFLAV)		

**4.4 ± 0.3 ± 0.5** <sup>1</sup> PROCARIO 93 CLEO  $E_{\text{cm}}^{\text{ee}} \approx 10.6$  GeV

<sup>1</sup> PROCARIO 93 quote  $0.041 \pm 0.003 \pm 0.005$  after correction for 2 kaon backgrounds assuming  $B(K^{*-} \nu_\tau) = 1.42 \pm 0.18\%$  and  $B(h^- K^0 \pi^0 \nu_\tau) = 0.48 \pm 0.48\%$ . We add  $0.003 \pm 0.003$  and multiply the sum by  $0.990 \pm 0.010$  to remove these corrections.

$$\Gamma(\pi^- 3\pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}} \quad \Gamma_{27}/\Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.04 ± 0.07 OUR FIT</b>		(Produced by HFLAV)		
<b>0.977 ± 0.069 ± 0.058</b>	6.1k	<sup>1</sup> SCHAEL	05C ALEP	1991-1995 LEP runs

<sup>1</sup> See footnote to SCHAEL 05C  $\Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)/\Gamma_{\text{total}}$  measurement for correlations with other measurements.

$$\Gamma(K^- 3\pi^0 \nu_\tau (\text{ex. } K^0, \eta))/\Gamma_{\text{total}} \quad \Gamma_{28}/\Gamma$$

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>4.8 ± 2.1 OUR FIT</b>		(Produced by HFLAV)		
<b>3.7 ± 2.1 ± 1.1</b>	22	BARATE	99K ALEP	1991–1995 LEP runs

• • • We do not use the following data for averages, fits, limits, etc. • • •

5 ± 13 <sup>1</sup> BUSKULIC 94E ALEP Repl. by BARATE 99K

<sup>1</sup> BUSKULIC 94E quote  $B(K^- \geq 0\pi^0 \geq 0K^0 \nu_\tau) - [B(K^- \nu_\tau) + B(K^- \pi^0 \nu_\tau) + B(K^- K^0 \nu_\tau) + B(K^- \pi^0 \pi^0 \nu_\tau) + B(K^- \pi^0 K^0 \nu_\tau)] = (5 \pm 13) \times 10^{-4}$  accounting for common systematic errors in BUSKULIC 94E and BUSKULIC 94F measurements of these modes. We assume  $B(K^- \geq 2K^0 \nu_\tau)$  and  $B(K^- \geq 4\pi^0 \nu_\tau)$  are negligible.

$$\Gamma(h^- 4\pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}} \quad \Gamma_{29}/\Gamma$$

$$\Gamma_{29}/\Gamma = (\Gamma_{30} + 0.3257\Gamma_{150} + 0.3257\Gamma_{154})/\Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.16 ± 0.04 OUR FIT</b>		(Produced by HFLAV)		
<b>0.16 ± 0.05 ± 0.05</b>		<sup>1</sup> PROCARIO	93 CLEO	$E_{\text{cm}}^{ee} \approx 10.6$ GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.16 ± 0.04 ± 0.09 232 <sup>2</sup> BUSKULIC 96 ALEP Repl. by SCHAEL 05C

<sup>1</sup> PROCARIO 93 quotes  $B(h^- 4\pi^0 \nu_\tau)/B(h^- \pi^0 \nu_\tau) = 0.006 \pm 0.002 \pm 0.002$ . We multiply by the ARTUSO 94 result for  $B(h^- \pi^0 \nu_\tau)$  to obtain  $B(h^- 4\pi^0 \nu_\tau)$ . PROCARIO 93 assume  $B(h^- \geq 5 \pi^0 \nu_\tau)$  is small and do not correct for it.

<sup>2</sup> BUSKULIC 96 quote result for  $\tau^- \rightarrow h^- \geq 4\pi^0 \nu_\tau$ . We assume  $B(h^- \geq 5\pi^0 \nu_\tau)$  is negligible.

$$\Gamma(h^- 4\pi^0 \nu_\tau (\text{ex. } K^0, \eta))/\Gamma_{\text{total}} \quad \Gamma_{30}/\Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.11 ± 0.04 OUR FIT</b>		(Produced by HFLAV)		
<b>0.112 ± 0.037 ± 0.035</b>	957	<sup>1</sup> SCHAEL	05C ALEP	1991-1995 LEP runs

<sup>1</sup> See footnote to SCHAEL 05C  $\Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)/\Gamma_{\text{total}}$  measurement for correlations with other measurements.

$$\Gamma(a_1(1260) \nu_\tau \rightarrow \pi^- \gamma \nu_\tau)/\Gamma_{\text{total}} \quad \Gamma_{31}/\Gamma = (0.0022\Gamma_{20} + 0.0022\Gamma_{70})/\Gamma$$

In the fit,  $\Gamma(\tau^- \rightarrow a_1(1260) \nu_\tau \rightarrow \pi^- \gamma \nu_\tau)/\Gamma_{\text{total}}$  is set equal to  $(\Gamma(\tau^- \rightarrow \pi^- 2\pi^0 \nu_\tau (\text{ex. } K^0)) + \Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0, \omega)))/\Gamma_{\text{total}} \times B(a_1(1260) \rightarrow \pi \gamma)/(1 - B(a_1(1260) \rightarrow \pi \gamma))$  and  $B(a_1(1260) \rightarrow \pi \gamma) = \Gamma(a_1(1260) \rightarrow \pi \gamma)/\Gamma_{\text{total}}^{a_1(1260)}$  is a nuisance fit variable with a  $\chi^2$  term corresponding to its estimate in reference SCHAEL 05C.

VALUE (units $10^{-4}$ )	DOCUMENT ID
<b>4.0 ± 1.5 OUR FIT</b>	(Produced by HFLAV)

$$\Gamma(K^- \geq 0\pi^0 \geq 0K^0 \geq 0\gamma \nu_\tau) / \Gamma_{\text{total}} \quad \Gamma_{32} / \Gamma$$

$$\Gamma_{32} / \Gamma = (\Gamma_{10} + \Gamma_{16} + \Gamma_{23} + \Gamma_{28} + \Gamma_{38} + \Gamma_{43} + 0.7196\Gamma_{152} + 0.1049\Gamma_{170}) / \Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.552 ± 0.029 OUR FIT</b>				(Produced by HFLAV)
<b>1.53 ± 0.04 OUR AVERAGE</b>				

1.528 ± 0.039 ± 0.040		<sup>1</sup> ABBIENDI	01J	OPAL	1990–1995 LEP runs
1.54 ± 0.24		ABREU	94K	DLPH	LEP 1992 Z data
1.70 ± 0.12 ± 0.19	202	<sup>2</sup> BATTLE	94	CLEO	$E_{\text{cm}}^{\text{ee}} \approx 10.6$ GeV

• • • We use the following data for averages but not for fits. • • •

1.520 ± 0.040 ± 0.041	4006	<sup>3</sup> BARATE	99K	ALEP	1991–1995 LEP runs
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• • • We do not use the following data for averages, fits, limits, etc. • • •

1.70 ± 0.05 ± 0.06	1610	<sup>4</sup> BUSKULIC	96	ALEP	Repl. by BARATE 99K
1.6 ± 0.4 ± 0.2	35	AIHARA	87B	TPC	$E_{\text{cm}}^{\text{ee}} = 29$ GeV
1.71 ± 0.29	53	MILLS	84	DLCO	$E_{\text{cm}}^{\text{ee}} = 29$ GeV

<sup>1</sup> The correlation coefficient between this measurement and the ABBIENDI 01J  $B(\tau^- \rightarrow K^- \nu_\tau)$  is 0.60.

<sup>2</sup> BATTLE 94 quote  $1.60 \pm 0.12 \pm 0.19$ . We add  $0.10 \pm 0.02$  to correct for their rejection of  $K_S^0 \rightarrow \pi^+ \pi^-$  decays.

<sup>3</sup> Not independent of BARATE 99K  $B(K^- \nu_\tau)$ ,  $B(K^- \pi^0 \nu_\tau)$ ,  $B(K^- 2\pi^0 \nu_\tau)$  (ex.  $K^0$ ),  $B(K^- 3\pi^0 \nu_\tau)$  (ex.  $K^0$ ),  $B(K^- K^0 \nu_\tau)$ , and  $B(K^- K^0 \pi^0 \nu_\tau)$  values.

<sup>4</sup> Not independent of BUSKULIC 96  $B(K^- \nu_\tau)$ ,  $B(K^- \pi^0 \nu_\tau)$ ,  $B(K^- 2\pi^0 \nu_\tau)$ ,  $B(K^- K^0 \nu_\tau)$ , and  $B(K^- K^0 \pi^0 \nu_\tau)$  values.

$$\Gamma(K^- \geq 1(\pi^0 \text{ or } K^0 \text{ or } \gamma) \nu_\tau) / \Gamma_{\text{total}} \quad \Gamma_{33} / \Gamma$$

$$\Gamma_{33} / \Gamma = (\Gamma_{16} + \Gamma_{23} + \Gamma_{28} + \Gamma_{38} + \Gamma_{43} + 0.7196\Gamma_{152} + 0.7196\Gamma_{154} + 0.1049\Gamma_{170}) / \Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.859 ± 0.028 OUR FIT</b>				(Produced by HFLAV)
<b>0.86 ± 0.05 OUR AVERAGE</b>				

• • • We use the following data for averages but not for fits. • • •

0.869 ± 0.031 ± 0.034		<sup>1</sup> ABBIENDI	01J	OPAL	1990–1995 LEP runs
0.69 ± 0.25		<sup>2</sup> ABREU	94K	DLPH	LEP 1992 Z data

• • • We do not use the following data for averages, fits, limits, etc. • • •

1.2 ± 0.5 $\begin{smallmatrix} +0.2 \\ -0.4 \end{smallmatrix}$	9	AIHARA	87B	TPC	$E_{\text{cm}}^{\text{ee}} = 29$ GeV
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<sup>1</sup> Not independent of ABBIENDI 01J  $B(\tau^- \rightarrow K^- \nu_\tau)$  and  $B(\tau^- \rightarrow K^- \geq 0\pi^0 \geq 0K^0 \geq 0\gamma \nu_\tau)$  values.

<sup>2</sup> Not independent of ABREU 94K  $B(K^- \nu_\tau)$  and  $B(K^- \geq 0 \text{ neutrals } \nu_\tau)$  measurements.

$$\Gamma(K_S^0(\text{particles})^- \nu_\tau) / \Gamma_{\text{total}} \quad \Gamma_{34} / \Gamma$$

$$\Gamma_{34} / \Gamma = (\frac{1}{2}\Gamma_{36} + \frac{1}{2}\Gamma_{38} + \frac{1}{2}\Gamma_{41} + \frac{1}{2}\Gamma_{43} + \frac{1}{2}\Gamma_{45} + \Gamma_{48} + \Gamma_{49} + \Gamma_{52} + \Gamma_{56} + 0.3598\Gamma_{156} + 0.339\Gamma_{170}) / \Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.943 ± 0.028 OUR FIT</b>				(Produced by HFLAV)
<b>0.918 ± 0.015 OUR AVERAGE</b>				

0.970 ± 0.058 ± 0.062	929	BARATE	98E	ALEP	1991–1995 LEP runs
0.97 ± 0.09 ± 0.06	141	AKERS	94G	OPAL	$E_{\text{cm}}^{\text{ee}} = 88$ –94 GeV



• • • We use the following data for averages but not for fits. • • •

0.915 ± 0.001 ± 0.015 398k <sup>1</sup>RUYU 14 BELL 669 fb<sup>-1</sup> E<sub>cm</sub><sup>ee</sup> = 10.6 GeV

<sup>1</sup>Not independent of RUYU 14 measurements of B(τ<sup>-</sup> → π<sup>-</sup> K<sup>0</sup> ν<sub>τ</sub>), B(τ<sup>-</sup> → K<sup>-</sup> K<sup>0</sup> ν<sub>τ</sub>), B(τ<sup>-</sup> → π<sup>-</sup> K<sup>0</sup> π<sup>0</sup> ν<sub>τ</sub>), B(τ<sup>-</sup> → K<sup>-</sup> K<sup>0</sup> π<sup>0</sup> ν<sub>τ</sub>), B(τ<sup>-</sup> → π<sup>-</sup> K<sub>S</sub><sup>0</sup> K<sub>S</sub><sup>0</sup> ν<sub>τ</sub>), and B(τ<sup>-</sup> → π<sup>-</sup> K<sub>S</sub><sup>0</sup> K<sub>S</sub><sup>0</sup> π<sup>0</sup> ν<sub>τ</sub>).

**Γ(h<sup>-</sup> K<sup>0</sup> ν<sub>τ</sub>)/Γ<sub>total</sub>**

**Γ<sub>35</sub>/Γ = (Γ<sub>36</sub> + Γ<sub>38</sub>)/Γ**

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.987 ± 0.014 OUR FIT</b>		(Produced by HFLAV)		
<b>0.90 ± 0.07 OUR AVERAGE</b>				

0.855 ± 0.036 ± 0.073 1242 COAN 96 CLEO E<sub>cm</sub><sup>ee</sup> ≈ 10.6 GeV

• • • We use the following data for averages but not for fits. • • •

1.01 ± 0.11 ± 0.07 555 <sup>1</sup>BARATE 98E ALEP 1991–1995 LEP runs

<sup>1</sup>Not independent of BARATE 98E B(τ<sup>-</sup> → π<sup>-</sup> K<sup>0</sup> ν<sub>τ</sub>) and B(τ<sup>-</sup> → K<sup>-</sup> K<sup>0</sup> ν<sub>τ</sub>) values.

**Γ(π<sup>-</sup> K<sup>0</sup> ν<sub>τ</sub>)/Γ<sub>total</sub>**

**Γ<sub>36</sub>/Γ**

VALUE (units 10 <sup>-3</sup> )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>8.38 ± 0.14 OUR FIT</b>		(Produced by HFLAV)		
<b>8.39 ± 0.22 OUR AVERAGE</b>		Error includes scale factor of 1.5. See the ideogram below.		

8.32 ± 0.02 ± 0.16 158k <sup>1</sup>RUYU 14 BELL 669 fb<sup>-1</sup> E<sub>cm</sub><sup>ee</sup> = 10.6 GeV

9.33 ± 0.68 ± 0.49 377 ABBIENDI 00C OPAL 1991–1995 LEP runs

9.28 ± 0.45 ± 0.34 937 <sup>2</sup>BARATE 99K ALEP 1991–1995 LEP runs

9.5 ± 1.5 ± 0.6 <sup>3</sup>ACCIARRI 95F L3 1991–1993 LEP runs

• • • We use the following data for averages but not for fits. • • •

8.55 ± 1.17 ± 0.66 509 <sup>4</sup>BARATE 98E ALEP 1991–1995 LEP runs

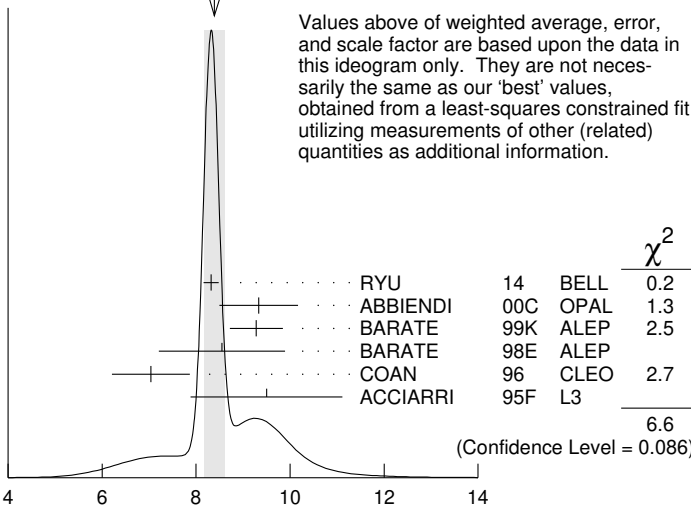
7.04 ± 0.41 ± 0.72 <sup>5</sup>COAN 96 CLEO E<sub>cm</sub><sup>ee</sup> ≈ 10.6 GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

8.08 ± 0.04 ± 0.26 53k EPIFANOV 07 BELL Repl. by RUYU 14

7.9 ± 1.0 ± 0.9 98 <sup>6</sup>BUSKULIC 96 ALEP Repl. by BARATE 99K

WEIGHTED AVERAGE  
8.39 ± 0.22 (Error scaled by 1.5)



Γ(π<sup>-</sup> K<sup>0</sup> ν<sub>τ</sub>)/Γ<sub>total</sub> (units 10<sup>-3</sup>)

<sup>1</sup>RUYU 14 reconstruct K<sup>0</sup>'s using K<sub>S</sub><sup>0</sup> → π<sup>+</sup> π<sup>-</sup> decays.

- <sup>2</sup> BARATE 99K measure  $K^0$ 's by detecting  $K_L^0$ 's in their hadron calorimeter.  
<sup>3</sup> ACCIARRI 95F do not identify  $\pi^-/K^-$  and assume  $B(K^- K^0 \nu_\tau) = (0.29 \pm 0.12)\%$ .  
<sup>4</sup> BARATE 98E reconstruct  $K^0$ 's using  $K_S^0 \rightarrow \pi^+ \pi^-$  decays. Not independent of BARATE 98E  $B(K^0 \text{ particles}^- \nu_\tau)$  value.  
<sup>5</sup> Not independent of COAN 96  $B(h^- K^0 \nu_\tau)$  and  $B(K^- K^0 \nu_\tau)$  measurements.  
<sup>6</sup> BUSKULIC 96 measure  $K^0$ 's by detecting  $K_L^0$ 's in their hadron calorimeter.

**$\Gamma(\pi^- \bar{K}^0 (\text{non-}K^*(892)^- ) \nu_\tau) / \Gamma_{\text{total}}$   $\Gamma_{37}/\Gamma$**

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>5.4±2.1</b>		<sup>1</sup> EPIFANOV 07	BELL	351 fb <sup>-1</sup> $E_{\text{cm}}^{ee} = 10.6$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<17	95	ACCIARRI 95F	L3	1991–1993 LEP runs
<sup>1</sup> EPIFANOV 07 quote $B(\tau^- \rightarrow K^*(892)^- \nu_\tau) B(K^*(892)^- \rightarrow K_S^0 \pi^-) / B(\tau^- \rightarrow K_S^0 \pi^- \nu_\tau) = 0.933 \pm 0.027$ . We multiply their $B(\tau^- \rightarrow \bar{K}^0 \pi^- \nu_\tau)$ by $[1 - (0.933 \pm 0.027)]$ to obtain this result.				

**$\Gamma(K^- K^0 \nu_\tau) / \Gamma_{\text{total}}$   $\Gamma_{38}/\Gamma$**

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>14.86±0.34 OUR FIT</b>		(Produced by HFLAV)		
<b>14.83±0.35 OUR AVERAGE</b>				
14.78±0.22±0.40	29k	<sup>1</sup> LEES 18B	BABR	468 fb <sup>-1</sup> $E_{\text{cm}}^{ee} = 10.6$ GeV
14.80±0.14±0.54	33k	<sup>2</sup> RYU 14	BELL	669 fb <sup>-1</sup> $E_{\text{cm}}^{ee} = 10.6$ GeV
16.2 ±2.1 ±1.1	150	<sup>3</sup> BARATE 99K	ALEP	1991–1995 LEP runs
15.8 ±4.2 ±1.7	46	<sup>4</sup> BARATE 98E	ALEP	1991–1995 LEP runs
15.1 ±2.1 ±2.2	111	COAN 96	CLEO	$E_{\text{cm}}^{ee} \approx 10.6$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
26 ±9 ±2	13	<sup>5</sup> BUSKULIC 96	ALEP	Repl. by BARATE 99K
<sup>1</sup> LEES 18B reconstructs $K_S^0$ 's using $K_S^0 \rightarrow \pi^+ \pi^-$ decays.				
<sup>2</sup> RYU 14 reconstruct $K^0$ 's using $K_S^0 \rightarrow \pi^+ \pi^-$ decays.				
<sup>3</sup> BARATE 99K measure $K^0$ 's by detecting $K_L^0$ 's in their hadron calorimeter.				
<sup>4</sup> BARATE 98E reconstruct $K^0$ 's using $K_S^0 \rightarrow \pi^+ \pi^-$ decays.				
<sup>5</sup> BUSKULIC 96 measure $K^0$ 's by detecting $K_L^0$ 's in their hadron calorimeter.				

**$\Gamma(K^- K^0 \geq 0 \pi^0 \nu_\tau) / \Gamma_{\text{total}}$   $\Gamma_{39}/\Gamma = (\Gamma_{38} + \Gamma_{43})/\Gamma$**

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.299±0.007 OUR FIT</b>		(Produced by HFLAV)		
<b>0.330±0.055±0.039</b>	124	ABBIENDI 00C	OPAL	1991–1995 LEP runs

**$\Gamma(h^- \bar{K}^0 \pi^0 \nu_\tau) / \Gamma_{\text{total}}$   $\Gamma_{40}/\Gamma = (\Gamma_{41} + \Gamma_{43})/\Gamma$**

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.532±0.013 OUR FIT</b>		(Produced by HFLAV)		
<b>0.50 ±0.06 OUR AVERAGE</b>		Error includes scale factor of 1.2.		
0.562±0.050±0.048	264	COAN 96	CLEO	$E_{\text{cm}}^{ee} \approx 10.6$ GeV
• • • We use the following data for averages but not for fits. • • •				
0.446±0.052±0.046	157	<sup>1</sup> BARATE 98E	ALEP	1991–1995 LEP runs
<sup>1</sup> Not independent of BARATE 98E $B(\tau^- \rightarrow \pi^- \bar{K}^0 \pi^0 \nu_\tau)$ and $B(\tau^- \rightarrow K^- K^0 \pi^0 \nu_\tau)$ values.				

$\Gamma(\pi^- \bar{K}^0 \pi^0 \nu_\tau) / \Gamma_{\text{total}}$   $\Gamma_{41} / \Gamma$ 

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
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**0.382 ± 0.013 OUR FIT** (Produced by HFLAV)**0.383 ± 0.014 OUR AVERAGE**

0.386 ± 0.004 ± 0.014	27k	<sup>1</sup> RYU	14	BELL	669 fb <sup>-1</sup> E <sub>cm</sub> <sup>ee</sup> = 10.6 GeV
0.347 ± 0.053 ± 0.037	299	<sup>2</sup> BARATE	99K	ALEP	1991–1995 LEP runs
0.294 ± 0.073 ± 0.037	142	<sup>3</sup> BARATE	98E	ALEP	1991–1995 LEP runs
0.41 ± 0.12 ± 0.03		<sup>4</sup> ACCIARRI	95F	L3	1991–1993 LEP runs

• • • We use the following data for averages but not for fits. • • •

0.417 ± 0.058 ± 0.044		<sup>5</sup> COAN	96	CLEO	E <sub>cm</sub> <sup>ee</sup> ≈ 10.6 GeV
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• • • We do not use the following data for averages, fits, limits, etc. • • •

0.32 ± 0.11 ± 0.05	23	<sup>6</sup> BUSKULIC	96	ALEP	Repl. by BARATE 99K
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<sup>1</sup> RYU 14 reconstruct  $K^0$ 's using  $K_S^0 \rightarrow \pi^+ \pi^-$  decays.<sup>2</sup> BARATE 99K measure  $K^0$ 's by detecting  $K_L^0$ 's in their hadron calorimeter.<sup>3</sup> BARATE 98E reconstruct  $K^0$ 's using  $K_S^0 \rightarrow \pi^+ \pi^-$  decays.<sup>4</sup> ACCIARRI 95F do not identify  $\pi^- / K^-$  and assume  $B(K^- K^0 \pi^0 \nu_\tau) = (0.05 \pm 0.05)\%$ .<sup>5</sup> Not independent of COAN 96  $B(h^- K^0 \pi^0 \nu_\tau)$  and  $B(K^- K^0 \pi^0 \nu_\tau)$  measurements.<sup>6</sup> BUSKULIC 96 measure  $K^0$ 's by detecting  $K_L^0$ 's in their hadron calorimeter. $\Gamma(\bar{K}^0 \rho^- \nu_\tau) / \Gamma_{\text{total}}$   $\Gamma_{42} / \Gamma$ 

VALUE (%)	DOCUMENT ID	TECN	COMMENT
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**0.22 ± 0.05 OUR AVERAGE**

0.250 ± 0.057 ± 0.044	<sup>1</sup> BARATE	99K	ALEP	1991–1995 LEP runs
0.188 ± 0.054 ± 0.038	<sup>2</sup> BARATE	98E	ALEP	1991–1995 LEP runs

<sup>1</sup> BARATE 99K measure  $K^0$ 's by detecting  $K_L^0$ 's in hadron calorimeter. They determine the  $\bar{K}^0 \rho^-$  fraction in  $\tau^- \rightarrow \pi^- \bar{K}^0 \pi^0 \nu_\tau$  decays to be  $(0.72 \pm 0.12 \pm 0.10)$  and multiply their  $B(\pi^- \bar{K}^0 \pi^0 \nu_\tau)$  measurement by this fraction to obtain the quoted result.<sup>2</sup> BARATE 98E reconstruct  $K^0$ 's using  $K_S^0 \rightarrow \pi^+ \pi^-$  decays. They determine the  $\bar{K}^0 \rho^-$  fraction in  $\tau^- \rightarrow \pi^- \bar{K}^0 \pi^0 \nu_\tau$  decays to be  $(0.64 \pm 0.09 \pm 0.10)$  and multiply their  $B(\pi^- \bar{K}^0 \pi^0 \nu_\tau)$  measurement by this fraction to obtain the quoted result. $\Gamma(K^- K^0 \pi^0 \nu_\tau) / \Gamma_{\text{total}}$   $\Gamma_{43} / \Gamma$ 

VALUE (units 10 <sup>-4</sup> )	EVTS	DOCUMENT ID	TECN	COMMENT
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**15.0 ± 0.7 OUR FIT** (Produced by HFLAV)**14.9 ± 0.7 OUR AVERAGE**

14.96 ± 0.20 ± 0.74	8.3k	<sup>1</sup> RYU	14	BELL	669 fb <sup>-1</sup> E <sub>cm</sub> <sup>ee</sup> = 10.6 GeV
14.3 ± 2.5 ± 1.5	78	<sup>2</sup> BARATE	99K	ALEP	1991–1995 LEP runs
15.2 ± 7.6 ± 2.1	15	<sup>3</sup> BARATE	98E	ALEP	1991–1995 LEP runs
14.5 ± 3.6 ± 2.0	32	COAN	96	CLEO	E <sub>cm</sub> <sup>ee</sup> ≈ 10.6 GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

10 ± 5 ± 3	5	<sup>4</sup> BUSKULIC	96	ALEP	Repl. by BARATE 99K
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<sup>1</sup> RYU 14 reconstruct  $K^0$ 's using  $K_S^0 \rightarrow \pi^+ \pi^-$  decays.<sup>2</sup> BARATE 99K measure  $K^0$ 's by detecting  $K_L^0$ 's in their hadron calorimeter.<sup>3</sup> BARATE 98E reconstruct  $K^0$ 's using  $K_S^0 \rightarrow \pi^+ \pi^-$  decays.<sup>4</sup> BUSKULIC 96 measure  $K^0$ 's by detecting  $K_L^0$ 's in their hadron calorimeter.

$$\Gamma(\pi^- \bar{K}^0 \geq 1\pi^0 \nu_\tau) / \Gamma_{\text{total}} \qquad \Gamma_{44} / \Gamma = (\Gamma_{41} + \Gamma_{45}) / \Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.408 ± 0.025 OUR FIT</b>	(Produced by HFLAV)			
<b>0.324 ± 0.074 ± 0.066</b>	148	ABBIENDI	00C OPAL	1991–1995 LEP runs

$$\Gamma(\pi^- \bar{K}^0 \pi^0 \pi^0 \nu_\tau (\text{ex. } K^0)) / \Gamma_{\text{total}} \qquad \Gamma_{45} / \Gamma$$

VALUE (units $10^{-3}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.26 ± 0.23 OUR FIT</b>	(Produced by HFLAV)				
<b>0.26 ± 0.24</b>			<sup>1</sup> BARATE	99R ALEP	1991–1995 LEP runs
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<0.66	95	17	<sup>2</sup> BARATE	99K ALEP	1991–1995 LEP runs
0.58 ± 0.33 ± 0.14		5	<sup>3</sup> BARATE	98E ALEP	1991–1995 LEP runs

<sup>1</sup> BARATE 99R combine the BARATE 98E and BARATE 99K measurements to obtain this value.

<sup>2</sup> BARATE 99K measure  $K^0$ 's by detecting  $K_L^0$ 's in their hadron calorimeter.

<sup>3</sup> BARATE 98E reconstruct  $K^0$ 's using  $K_S^0 \rightarrow \pi^+ \pi^-$  decays.

$$\Gamma(K^- K^0 \pi^0 \pi^0 \nu_\tau) / \Gamma_{\text{total}} \qquad \Gamma_{46} / \Gamma$$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.16 × 10<sup>-3</sup></b>	95	<sup>1</sup> BARATE	99R ALEP	1991–1995 LEP runs
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<0.18 × 10 <sup>-3</sup>	95	<sup>2</sup> BARATE	99K ALEP	1991–1995 LEP runs
<0.39 × 10 <sup>-3</sup>	95	<sup>3</sup> BARATE	98E ALEP	1991–1995 LEP runs

<sup>1</sup> BARATE 99R combine the BARATE 98E and BARATE 99K bounds to obtain this value.

<sup>2</sup> BARATE 99K measure  $K^0$ 's by detecting  $K_L^0$ 's in hadron calorimeter.

<sup>3</sup> BARATE 98E reconstruct  $K^0$ 's by using  $K_S^0 \rightarrow \pi^+ \pi^-$  decays.

$$\Gamma(\pi^- K^0 \bar{K}^0 \nu_\tau) / \Gamma_{\text{total}} \qquad \Gamma_{47} / \Gamma = (\Gamma_{48} + \Gamma_{49} + \Gamma_{50}) / \Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.155 ± 0.024 OUR FIT</b>	(Produced by HFLAV)			

• • • We use the following data for averages but not for fits. • • •

**0.153 ± 0.030 ± 0.016** 74 <sup>1</sup> BARATE 98E ALEP 1991–1995 LEP runs

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.31 ± 0.12 ± 0.04 <sup>2</sup> ACCIARRI 95F L3 1991–1993 LEP runs

<sup>1</sup> BARATE 98E obtain this value by adding twice their  $B(\pi^- K_S^0 K_S^0 \nu_\tau)$  value to their  $B(\pi^- K_S^0 K_L^0 \nu_\tau)$  value.

<sup>2</sup> ACCIARRI 95F assume  $B(\pi^- K_S^0 K_S^0 \nu) = B(\pi^- K_S^0 K_L^0 \nu) = 1/2 B(\pi^- K_S^0 K_L^0 \nu)$ .

$$\Gamma(\pi^- K_S^0 K_S^0 \nu_\tau) / \Gamma_{\text{total}} \qquad \Gamma_{48} / \Gamma$$

Bose-Einstein correlations might make the mixing fraction different than 1/4.

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2.35 ± 0.06 OUR FIT</b>	(Produced by HFLAV)			
<b>2.32 ± 0.06 OUR AVERAGE</b>				
2.33 ± 0.03 ± 0.09	6.7k	RYU	14 BELL	669 fb <sup>-1</sup> $E_{\text{cm}}^{ee} = 10.6$ GeV
2.31 ± 0.04 ± 0.08	5.0k	<sup>1</sup> LEES	12Y BABR	468 fb <sup>-1</sup> $E_{\text{cm}}^{ee} = 10.6$ GeV
2.6 ± 1.0 ± 0.5	6	BARATE	98E ALEP	1991–1995 LEP runs
2.3 ± 0.5 ± 0.3	42	COAN	96 CLEO	$E_{\text{cm}}^{ee} \approx 10.6$ GeV

<sup>1</sup> The correlation coefficient between this measurement and the LEES 12Y  $\Gamma(\tau^- \rightarrow \pi^- K_S^0 K_S^0 \pi^0 \nu_\tau) / \Gamma_{\text{total}}$  one is 0.0828.

$$\Gamma(\pi^- K_S^0 K_L^0 \nu_\tau)/\Gamma_{\text{total}} \quad \Gamma_{49}/\Gamma$$

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>10.8±2.4 OUR FIT</b>		(Produced by HFLAV)		
<b>10.1±2.3±1.3</b>	68	BARATE	98E	ALEP 1991–1995 LEP runs

$$\Gamma(\pi^- K_L^0 K_L^0 \nu_\tau)/\Gamma_{\text{total}} \quad \Gamma_{50}/\Gamma = \Gamma_{48}/\Gamma$$

VALUE (units $10^{-4}$ )	DOCUMENT ID
<b>2.35±0.06 OUR FIT</b>	(Produced by HFLAV)

$$\Gamma(\pi^- K^0 \bar{K}^0 \pi^0 \nu_\tau)/\Gamma_{\text{total}} \quad \Gamma_{51}/\Gamma = (\Gamma_{52} + \Gamma_{56} + \Gamma_{57})/\Gamma$$

VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT
<b>3.6±1.2 OUR FIT</b>	(Produced by HFLAV)		

• • • We use the following data for averages but not for fits. • • •

<b>3.1±2.3</b>	<sup>1</sup> BARATE	99R	ALEP	1991–1995 LEP runs
	<sup>1</sup> BARATE	99R	combine	BARATE 98E $\Gamma(\pi^- K_S^0 K_S^0 \pi^0 \nu_\tau)/\Gamma_{\text{total}}$ and $\Gamma(\pi^- K_S^0 K_L^0 \pi^0 \nu_\tau)/\Gamma_{\text{total}}$ measurements to obtain this value.

$$\Gamma(\pi^- K_S^0 K_S^0 \pi^0 \nu_\tau)/\Gamma_{\text{total}} \quad \Gamma_{52}/\Gamma$$

VALUE (units $10^{-5}$ )	CL% EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.82±0.21 OUR FIT</b>		(Produced by HFLAV)		
<b>1.80±0.21 OUR AVERAGE</b>				
2.00±0.22±0.20	303	RYU	14	BELL 669 fb <sup>-1</sup> $E_{\text{cm}}^{ee}=10.6$ GeV
1.60±0.20±0.22	409	<sup>1</sup> LEES	12Y	BABR 468 fb <sup>-1</sup> $E_{\text{cm}}^{ee}=10.6$ GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

<20	95	BARATE	98E	ALEP 1991–1995 LEP runs
		<sup>1</sup>		The correlation coefficient between this measurement and the LEES 12Y $\Gamma(\tau^- \rightarrow \pi^- K_S^0 K_S^0 \nu_\tau)/\Gamma_{\text{total}}$ one is 0.0828.

$$\Gamma(K^{*-} K^0 \pi^0 \nu_\tau \rightarrow \pi^- K_S^0 K_S^0 \pi^0 \nu_\tau)/\Gamma_{\text{total}} \quad \Gamma_{53}/\Gamma$$

VALUE (units $10^{-6}$ )	DOCUMENT ID	TECN	COMMENT
<b>10.8±1.4±1.5</b>	RYU	14	BELL 669 fb <sup>-1</sup> $E_{\text{cm}}^{ee}=10.6$ GeV

$$\Gamma(f_1(1285)\pi^- \nu_\tau \rightarrow \pi^- K_S^0 K_S^0 \pi^0 \nu_\tau)/\Gamma_{\text{total}} \quad \Gamma_{54}/\Gamma$$

VALUE (units $10^{-6}$ )	DOCUMENT ID	TECN	COMMENT
<b>6.8±1.3±0.7</b>	RYU	14	BELL 669 fb <sup>-1</sup> $E_{\text{cm}}^{ee}=10.6$ GeV

$$\Gamma(f_1(1420)\pi^- \nu_\tau \rightarrow \pi^- K_S^0 K_S^0 \pi^0 \nu_\tau)/\Gamma_{\text{total}} \quad \Gamma_{55}/\Gamma$$

VALUE (units $10^{-6}$ )	DOCUMENT ID	TECN	COMMENT
<b>2.4±0.5±0.6</b>	RYU	14	BELL 669 fb <sup>-1</sup> $E_{\text{cm}}^{ee}=10.6$ GeV

$$\Gamma(\pi^- K_S^0 K_L^0 \pi^0 \nu_\tau)/\Gamma_{\text{total}} \quad \Gamma_{56}/\Gamma$$

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>3.2±1.2 OUR FIT</b>		(Produced by HFLAV)		
<b>3.1±1.1±0.5</b>	11	BARATE	98E	ALEP 1991–1995 LEP runs

$$\Gamma(\pi^- K_L^0 K_L^0 \pi^0 \nu_\tau) / \Gamma_{\text{total}} \quad \Gamma_{57} / \Gamma = \Gamma_{52} / \Gamma$$

VALUE (units $10^{-5}$ )	DOCUMENT ID
<b>1.82 ± 0.21 OUR FIT</b>	(Produced by HFLAV)

$$\Gamma(K^- K_S^0 K_S^0 \nu_\tau) / \Gamma_{\text{total}} \quad \Gamma_{58} / \Gamma$$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt; 6.3 × 10<sup>-7</sup></b>	90	LEES	12Y BABR	468 fb <sup>-1</sup> $E_{\text{cm}}^{ee} = 10.6$ GeV

$$\Gamma(K^- K_S^0 K_S^0 \pi^0 \nu_\tau) / \Gamma_{\text{total}} \quad \Gamma_{59} / \Gamma$$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt; 4.0 × 10<sup>-7</sup></b>	90	LEES	12Y BABR	468 fb <sup>-1</sup> $E_{\text{cm}}^{ee} = 10.6$ GeV

$$\Gamma(K^0 h^+ h^- h^- \geq 0 \text{ neutrals } \nu_\tau) / \Gamma_{\text{total}} \quad \Gamma_{60} / \Gamma$$

VALUE (%)	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt; 0.17</b>	95	TSCHIRHART	88 HRS	$E_{\text{cm}}^{ee} = 29$ GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

< 0.27	90	BELTRAMI	85 HRS	$E_{\text{cm}}^{ee} = 29$ GeV
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$$\Gamma(K^0 h^+ h^- h^- \nu_\tau) / \Gamma_{\text{total}} \quad \Gamma_{61} / \Gamma$$

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2.5 ± 2.0 OUR FIT</b>		(Produced by HFLAV)		

<b>2.3 ± 1.9 ± 0.7</b>	6	<sup>1</sup> BARATE	98E ALEP	1991–1995 LEP runs
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<sup>1</sup>BARATE 98E reconstruct  $K^0$ 's using  $K_S^0 \rightarrow \pi^+ \pi^-$  decays.

$$\Gamma(h^- h^- h^+ \geq 0 \text{ neutrals } \geq 0 K_L^0 \nu_\tau) / \Gamma_{\text{total}} \quad \Gamma_{62} / \Gamma$$

$$\Gamma_{62} / \Gamma = (0.34598\Gamma_{36} + 0.34598\Gamma_{38} + 0.34598\Gamma_{41} + 0.34598\Gamma_{43} + 0.4247\Gamma_{48} + 0.6920\Gamma_{49} + 0.4247\Gamma_{52} + 0.6920\Gamma_{56} + 0.6534\Gamma_{61} + \Gamma_{70} + \Gamma_{78} + \Gamma_{85} + \Gamma_{86} + \Gamma_{97} + \Gamma_{103} + \Gamma_{106} + \Gamma_{107} + 0.2804\Gamma_{150} + 0.2804\Gamma_{152} + 0.2804\Gamma_{154} + 0.2628\Gamma_{156} + 0.7259\Gamma_{170} + 0.9078\Gamma_{178} + 0.9078\Gamma_{179} + 0.9078\Gamma_{180}) / \Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>15.20 ± 0.06 OUR FIT</b>		(Produced by HFLAV)		

**14.8 ± 0.4 OUR AVERAGE**

14.4 ± 0.6 ± 0.3		ADEVA	91F L3	$E_{\text{cm}}^{ee} = 88.3\text{--}94.3$ GeV
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15.0 ± 0.4 ± 0.3		BEHREND	89B CELL	$E_{\text{cm}}^{ee} = 14\text{--}47$ GeV
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15.1 ± 0.8 ± 0.6		AIHARA	87B TPC	$E_{\text{cm}}^{ee} = 29$ GeV
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• • • We do not use the following data for averages, fits, limits, etc. • • •

13.5 ± 0.3 ± 0.3		ABACHI	89B HRS	$E_{\text{cm}}^{ee} = 29$ GeV
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12.8 ± 1.0 ± 0.7		<sup>1</sup> BURCHAT	87 MRK2	$E_{\text{cm}}^{ee} = 29$ GeV
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12.1 ± 0.5 ± 1.2		RUCKSTUHL	86 DLCO	$E_{\text{cm}}^{ee} = 29$ GeV
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12.8 ± 0.5 ± 0.8	1420	SCHMIDKE	86 MRK2	$E_{\text{cm}}^{ee} = 29$ GeV
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15.3 ± 1.1 $\begin{smallmatrix} +1.3 \\ -1.6 \end{smallmatrix}$	367	ALTHOFF	85 TASS	$E_{\text{cm}}^{ee} = 34.5$ GeV
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13.6 ± 0.5 ± 0.8		BARTEL	85F JADE	$E_{\text{cm}}^{ee} = 34.6$ GeV
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12.2 ± 1.3 ± 3.9		<sup>2</sup> BERGER	85 PLUT	$E_{\text{cm}}^{ee} = 34.6$ GeV
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13.3 ± 0.3 ± 0.6		FERNANDEZ	85 MAC	$E_{\text{cm}}^{ee} = 29$ GeV
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24 ± 6	35	BRANDELIK	80 TASS	$E_{\text{cm}}^{ee} = 30$ GeV
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32 ± 5	692	<sup>3</sup> BACINO	78B	DLCO	$E_{\text{cm}}^{ee} = 3.1\text{--}7.4$ GeV
35 ± 11		<sup>3</sup> BRANDELIK	78	DASP	Assumes $V\text{--}A$ decay
18 ± 6.5	33	<sup>3</sup> JAROS	78	LGW	$E_{\text{cm}}^{ee} > 6$ GeV

<sup>1</sup> BURCHAT 87 value is not independent of SCHMIDKE 86 value.

<sup>2</sup> Not independent of BERGER 85  $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau)/\Gamma_{\text{total}}$ ,  $\Gamma(e^- \bar{\nu}_e \nu_\tau)/\Gamma_{\text{total}}$ ,  $\Gamma(h^- \geq 1 \text{ neutrals } \nu_\tau)/\Gamma_{\text{total}}$ , and  $\Gamma(h^- \geq 0 K_L^0 \nu_\tau)/\Gamma_{\text{total}}$ , and therefore not used in the fit.

<sup>3</sup> Low energy experiments are not in average or fit because the systematic errors in background subtraction are judged to be large.

$$\Gamma(h^- h^- h^+ \geq 0 \text{ neutrals } \nu_\tau \text{ (ex. } K_S^0 \rightarrow \pi^+ \pi^-) \text{ ("3-prong")})/\Gamma_{\text{total}} \quad \Gamma_{63}/\Gamma$$

$$\Gamma_{63}/\Gamma = (\Gamma_{70} + \Gamma_{78} + \Gamma_{85} + \Gamma_{86} + \Gamma_{97} + \Gamma_{103} + \Gamma_{106} + \Gamma_{107} + 0.2804\Gamma_{150} + 0.2804\Gamma_{152} + 0.2804\Gamma_{154} + 0.491\Gamma_{170} + 0.9078\Gamma_{178} + 0.9078\Gamma_{179} + 0.9078\Gamma_{180})/\Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
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**14.55 ± 0.06 OUR FIT** (Produced by HFLAV)

**14.61 ± 0.06 OUR AVERAGE**

14.556 ± 0.105 ± 0.076		<sup>1</sup> ACHARD	01D	L3	1992–1995 LEP runs
14.96 ± 0.09 ± 0.22	10.4k	AKERS	95Y	OPAL	1991–1994 LEP runs

• • • We use the following data for averages but not for fits. • • •

14.652 ± 0.067 ± 0.086		SCHAEL	05C	ALEP	1991–1995 LEP runs
14.569 ± 0.093 ± 0.048	23k	<sup>2</sup> ABREU	01M	DLPH	1992–1995 LEP runs
14.22 ± 0.10 ± 0.37		<sup>3</sup> BALEST	95C	CLEO	$E_{\text{cm}}^{ee} \approx 10.6$ GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

15.26 ± 0.26 ± 0.22		ACTON	92H	OPAL	Repl. by AKERS 95Y
13.3 ± 0.3 ± 0.8		<sup>4</sup> ALBRECHT	92D	ARG	$E_{\text{cm}}^{ee} = 9.4\text{--}10.6$ GeV
14.35 $^{+0.40}_{-0.45}$ ± 0.24		DECAMP	92C	ALEP	1989–1990 LEP runs

<sup>1</sup> The correlation coefficients between this measurement and the ACHARD 01D measurements of  $B(\tau \rightarrow \text{"1-prong"})$  and  $B(\tau \rightarrow \text{"5-prong"})$  are  $-0.978$  and  $-0.19$  respectively.

<sup>2</sup> The correlation coefficients between this measurement and the ABREU 01M measurements of  $B(\tau \rightarrow \text{1-prong})$  and  $B(\tau \rightarrow \text{5-prong})$  are  $-0.98$  and  $-0.08$  respectively.

<sup>3</sup> Not independent of BALEST 95C  $B(h^- h^- h^+ \nu_\tau)$  and  $B(h^- h^- h^+ \pi^0 \nu_\tau)$  values, and BORTOLETTO 93  $B(h^- h^- h^+ 2\pi^0 \nu_\tau)/B(h^- h^- h^+ \geq 0 \text{ neutrals } \nu_\tau)$  value.

<sup>4</sup> This ALBRECHT 92D value is not independent of their  $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau)\Gamma(e^- \bar{\nu}_e \nu_\tau)/\Gamma_{\text{total}}^2$  value.

$$\Gamma(h^- h^- h^+ \nu_\tau)/\Gamma_{\text{total}} \quad \Gamma_{64}/\Gamma$$

$$\Gamma_{64}/\Gamma = (0.34598\Gamma_{36} + 0.34598\Gamma_{38} + \Gamma_{70} + \Gamma_{97} + \Gamma_{106} + 0.491\Gamma_{170} + 0.0153\Gamma_{178} + 0.0153\Gamma_{179})/\Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
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**9.80 ± 0.05 OUR FIT** (Produced by HFLAV)

• • • We use the following data for averages but not for fits. • • •

<b>7.6 ± 0.1 ± 0.5</b>	7.5k	<sup>1</sup> ALBRECHT	96E	ARG	$E_{\text{cm}}^{ee} = 9.4\text{--}10.6$ GeV
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• • • We do not use the following data for averages, fits, limits, etc. • • •

9.92 ± 0.10 ± 0.09	11.2k	<sup>2</sup> BUSKULIC	96	ALEP	Repl. by SCHAEL 05C
9.49 ± 0.36 ± 0.63		DECAMP	92C	ALEP	Repl. by SCHAEL 05C
8.7 ± 0.7 ± 0.3	694	<sup>3</sup> BEHREND	90	CELL	$E_{\text{cm}}^{ee} = 35$ GeV
7.0 ± 0.3 ± 0.7	1566	<sup>4</sup> BAND	87	MAC	$E_{\text{cm}}^{ee} = 29$ GeV
6.7 ± 0.8 ± 0.9		<sup>5</sup> BURCHAT	87	MRK2	$E_{\text{cm}}^{ee} = 29$ GeV

6.4 ±0.4 ±0.9		<sup>6</sup> RUCKSTUHL	86	DLCO	$E_{cm}^{ee} = 29$ GeV
7.8 ±0.5 ±0.8	890	SCHMIDKE	86	MRK2	$E_{cm}^{ee} = 29$ GeV
8.4 ±0.4 ±0.7	1255	<sup>6</sup> FERNANDEZ	85	MAC	$E_{cm}^{ee} = 29$ GeV
9.7 ±2.0 ±1.3		BEHREND	84	CELL	$E_{cm}^{ee} = 14,22$ GeV

<sup>1</sup> ALBRECHT 96E not independent of ALBRECHT 93C  $\Gamma(h^- h^- h^+ \nu_\tau(\text{ex. } K^0)) \times \Gamma(\text{particle}^- \geq 0 \text{ neutrals} \geq 0 K_L^0 \nu_\tau) / \Gamma_{\text{total}}^2$  value.

<sup>2</sup> BUSKULIC 96 quote  $B(h^- h^- h^+ \nu_\tau(\text{ex. } K^0)) = 9.50 \pm 0.10 \pm 0.11$ . We add 0.42 to remove their  $K^0$  correction and reduce the systematic error accordingly.

<sup>3</sup> BEHREND 90 subtract 0.3% to account for the  $\tau^- \rightarrow K^*(892)^- \nu_\tau$  contribution to measured events.

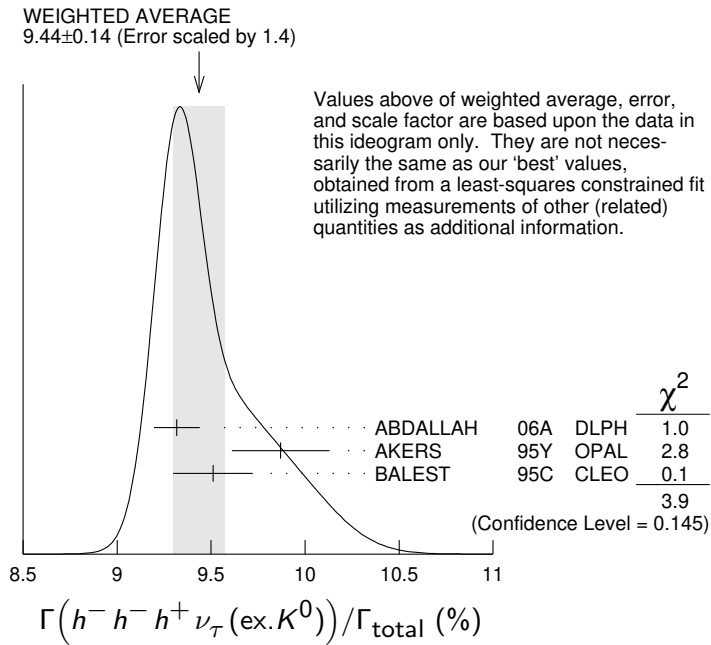
<sup>4</sup> BAND 87 subtract for charged kaon modes; not independent of FERNANDEZ 85 value.

<sup>5</sup> BURCHAT 87 value is not independent of SCHMIDKE 86 value.

<sup>6</sup> Value obtained by multiplying paper's  $R = B(h^- h^- h^+ \nu_\tau) / B(3\text{-prong})$  by  $B(3\text{-prong}) = 0.143$  and subtracting 0.3% for  $K^*(892)$  background.

$$\frac{\Gamma(h^- h^- h^+ \nu_\tau(\text{ex. } K^0))}{\Gamma_{\text{total}}} \quad \Gamma_{65}/\Gamma$$

$\Gamma_{65}/\Gamma = (\Gamma_{70} + \Gamma_{97} + \Gamma_{106} + 0.491\Gamma_{170} + 0.0153\Gamma_{178} + 0.0153\Gamma_{179}) / \Gamma$					
<u>VALUE (%)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	
<b>9.46 ±0.05 OUR FIT</b>		(Produced by HFLAV)			
<b>9.44 ±0.14 OUR AVERAGE</b>		Error includes scale factor of 1.4.		See the ideogram below.	
9.317 ±0.090 ±0.082	12.2k	<sup>1</sup> ABDALLAH	06A	DLPH	1992–1995 LEP runs
9.51 ±0.07 ±0.20	37.7k	BALEST	95C	CLEO	$E_{cm}^{ee} \approx 10.6$ GeV
● ● ● We use the following data for averages but not for fits. ● ● ●					
9.87 ±0.10 ±0.24		<sup>2</sup> AKERS	95Y	OPAL	1991–1994 LEP runs
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●					
9.50 ±0.10 ±0.11	11.2k	<sup>3</sup> BUSKULIC	96	ALEP	Repl. by SCHAEEL 05C



<sup>1</sup> See footnote to ABDALLAH 06A  $\Gamma(\tau^- \rightarrow h^- \nu_\tau) / \Gamma_{\text{total}}$  measurement for correlations with other measurements.

<sup>2</sup> Not independent of AKERS 95Y  $B(h^- h^- h^+ \geq 0 \text{ neutrals } \nu_\tau(\text{ex. } K_S^0 \rightarrow \pi^+ \pi^-))$  and  $B(h^- h^- h^+ \nu_\tau(\text{ex. } K^0)) / B(h^- h^- h^+ \geq 0 \text{ neutrals } \nu_\tau(\text{ex. } K_S^0 \rightarrow \pi^+ \pi^-))$  values.

<sup>3</sup> Not independent of BUSKULIC 96  $B(h^- h^- h^+ \nu_\tau)$  value.



$$\Gamma(h^- h^- h^+ \nu_\tau (\text{ex. } K^0)) / \Gamma(h^- h^- h^+ \geq 0 \text{ neutrals } \nu_\tau (\text{ex. } K_S^0 \rightarrow \pi^+ \pi^-)) \quad \Gamma_{65} / \Gamma_{63}$$

$$\Gamma_{65} / \Gamma_{63} = (\Gamma_{70} + \Gamma_{97} + \Gamma_{106} + 0.491\Gamma_{170} + 0.0153\Gamma_{178} + 0.0153\Gamma_{179}) / (0.4247\Gamma_{52} + \Gamma_{70} + \Gamma_{78} + \Gamma_{85} + \Gamma_{89} + \Gamma_{97} + \Gamma_{103} + \Gamma_{106} + \Gamma_{107} + 0.2804\Gamma_{150} + 0.2302\Gamma_{151} + 0.2804\Gamma_{152} + 0.2804\Gamma_{154} + 0.1131\Gamma_{156} + 0.3257\Gamma_{160} + 0.491\Gamma_{170} + 0.9078\Gamma_{178} + 0.9078\Gamma_{179} + 0.9078\Gamma_{180} + 0.892\Gamma_{182})$$

VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT
<b>64.98 ± 0.31 OUR FIT</b> (Produced by HFLAV)			
<b>66.0 ± 0.4 ± 1.4</b>	AKERS	95Y OPAL	1991–1994 LEP runs

$$\Gamma(h^- h^- h^+ \nu_\tau (\text{ex. } K^0, \omega)) / \Gamma_{\text{total}} \quad \Gamma_{66} / \Gamma$$

$$\Gamma_{66} / \Gamma = (\Gamma_{70} + \Gamma_{97} + \Gamma_{106} + 0.491\Gamma_{170}) / \Gamma$$

VALUE (%)	DOCUMENT ID
<b>9.43 ± 0.05 OUR FIT</b> (Produced by HFLAV)	

$$\Gamma(\pi^- \pi^+ \pi^- \nu_\tau) / \Gamma_{\text{total}} \quad \Gamma_{67} / \Gamma = (0.34598\Gamma_{36} + \Gamma_{70} + 0.0153\Gamma_{178}) / \Gamma$$

VALUE (%)	DOCUMENT ID
<b>9.31 ± 0.05 OUR FIT</b> (Produced by HFLAV)	

$$\Gamma(\pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0)) / \Gamma_{\text{total}} \quad \Gamma_{68} / \Gamma = (\Gamma_{70} + 0.0153\Gamma_{178}) / \Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>9.02 ± 0.05 OUR FIT</b> (Produced by HFLAV)				
<b>8.77 ± 0.13 OUR AVERAGE</b>		Error includes scale factor of 1.1.		

8.42 ± 0.00 <sup>+0.26</sup> <sub>-0.25</sub>	8.9M	<sup>1</sup> LEE	10	BELL	666 fb <sup>-1</sup>	$E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
8.83 ± 0.01 ± 0.13	1.6M	<sup>2</sup> AUBERT	08	BABR	342 fb <sup>-1</sup>	$E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
9.13 ± 0.05 ± 0.46	43k	<sup>3</sup> BRIERE	03	CLE3		$E_{\text{cm}}^{\text{ee}} = 10.6$ GeV

<sup>1</sup> Quoted statistical error is 0.003%. Correlation matrix for LEE 10 branching fractions:

- (1)  $\Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0)) / \Gamma_{\text{total}}$
- (2)  $\Gamma(\tau^- \rightarrow K^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0)) / \Gamma_{\text{total}}$
- (3)  $\Gamma(\tau^- \rightarrow K^- K^+ \pi^- \nu_\tau) / \Gamma_{\text{total}}$
- (4)  $\Gamma(\tau^- \rightarrow K^- K^+ K^- \nu_\tau) / \Gamma_{\text{total}}$

	(1)	(2)	(3)
(2)	0.175		
(3)	0.049	0.080	
(4)	-0.053	0.035	-0.008

<sup>2</sup> Correlation matrix for AUBERT 08 branching fractions:

- (1)  $\Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0)) / \Gamma_{\text{total}}$
- (2)  $\Gamma(\tau^- \rightarrow K^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0)) / \Gamma_{\text{total}}$
- (3)  $\Gamma(\tau^- \rightarrow K^- K^+ \pi^- \nu_\tau) / \Gamma_{\text{total}}$
- (4)  $\Gamma(\tau^- \rightarrow K^- K^+ K^- \nu_\tau) / \Gamma_{\text{total}}$

	(1)	(2)	(3)
(2)	0.544		
(3)	0.390	0.177	
(4)	0.031	0.093	0.087

<sup>3</sup> 47% correlated with BRIERE 03  $\tau^- \rightarrow K^- \pi^+ \pi^- \nu_\tau$  and 71% correlated with  $\tau^- \rightarrow K^- K^+ \pi^- \nu_\tau$  because of a common 5% normalization error.

$$\Gamma(\pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0), \text{ non-axial vector}) / \Gamma(\pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0)) \quad \Gamma_{69}/\Gamma_{68}$$

$$\Gamma_{69}/\Gamma_{68} = \Gamma_{69}/(\Gamma_{70} + 0.0153\Gamma_{177})$$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<0.261	95	<sup>1</sup> ACKERSTAFF 97R	OPAL	1992–1994 LEP runs

<sup>1</sup> Model-independent limit from structure function analysis on contribution to  $B(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0))$  from non-axial vectors.

$$\Gamma(\pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0, \omega)) / \Gamma_{\text{total}} \quad \Gamma_{70}/\Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>8.99 ± 0.05 OUR FIT</b>	(Produced by HFLAV)			
<b>9.041 ± 0.060 ± 0.076</b>	29k	<sup>1</sup> SCHAEL	05C ALEP	1991-1995 LEP runs

<sup>1</sup> See footnote to SCHAEL 05C  $\Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau) / \Gamma_{\text{total}}$  measurement for correlations with other measurements.

$$\Gamma(h^- h^- h^+ \geq 1 \text{ neutrals } \nu_\tau) / \Gamma_{\text{total}} \quad \Gamma_{71}/\Gamma$$

$$\Gamma_{71}/\Gamma = (0.34598\Gamma_{41} + 0.34598\Gamma_{43} + 0.4247\Gamma_{48} + 0.4247\Gamma_{52} + \Gamma_{78} + \Gamma_{85} + \Gamma_{86} + \Gamma_{103} + \Gamma_{107} + 0.2804\Gamma_{150} + 0.2804\Gamma_{152} + 0.2804\Gamma_{154} + 0.2926\Gamma_{156} + 0.892\Gamma_{178} + 0.892\Gamma_{179} + 0.9078\Gamma_{180}) / \Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>5.29 ± 0.05 OUR FIT</b>	(Produced by HFLAV)			

• • • We do not use the following data for averages, fits, limits, etc. • • •

5.6 ± 0.7 ± 0.3	352	<sup>1</sup> BEHREND	90	CELL	$E_{\text{cm}}^{ee} = 35 \text{ GeV}$
4.2 ± 0.5 ± 0.9	203	<sup>2</sup> ALBRECHT	87L	ARG	$E_{\text{cm}}^{ee} = 10 \text{ GeV}$
6.1 ± 0.8 ± 0.9		<sup>3</sup> BURCHAT	87	MRK2	$E_{\text{cm}}^{ee} = 29 \text{ GeV}$
7.6 ± 0.4 ± 0.9		<sup>4,5</sup> RUCKSTUHL	86	DLCO	$E_{\text{cm}}^{ee} = 29 \text{ GeV}$
4.7 ± 0.5 ± 0.8	530	<sup>6</sup> SCHMIDKE	86	MRK2	$E_{\text{cm}}^{ee} = 29 \text{ GeV}$
5.6 ± 0.4 ± 0.7		<sup>5</sup> FERNANDEZ	85	MAC	$E_{\text{cm}}^{ee} = 29 \text{ GeV}$
6.2 ± 2.3 ± 1.7		BEHREND	84	CELL	$E_{\text{cm}}^{ee} = 14,22 \text{ GeV}$

<sup>1</sup> BEHREND 90 value is not independent of BEHREND 90  $B(3h\nu_\tau \geq 1 \text{ neutrals}) + B(5\text{-prong})$ .

<sup>2</sup> ALBRECHT 87L measure the product of branching ratios  $B(3\pi^\pm \pi^0 \nu_\tau) B((e\bar{\nu} \text{ or } \mu\bar{\nu} \text{ or } \pi \text{ or } K \text{ or } \rho)\nu_\tau) = 0.029$  and use the PDG 86 values for the second branching ratio which sum to  $0.69 \pm 0.03$  to get the quoted value.

<sup>3</sup> BURCHAT 87 value is not independent of SCHMIDKE 86 value.

<sup>4</sup> Contributions from kaons and from  $>1\pi^0$  are subtracted. Not independent of  $(3\text{-prong} + 0\pi^0)$  and  $(3\text{-prong} + \geq 0\pi^0)$  values.

<sup>5</sup> Value obtained using paper's  $R = B(h^- h^- h^+ \nu_\tau) / B(3\text{-prong})$  and current  $B(3\text{-prong}) = 0.143$ .

<sup>6</sup> Not independent of SCHMIDKE 86  $h^- h^- h^+ \nu_\tau$  and  $h^- h^- h^+ (\geq 0\pi^0)\nu_\tau$  values.

$$\Gamma(h^- h^- h^+ \geq 1\pi^0 \nu_\tau (\text{ex. } K^0)) / \Gamma_{\text{total}} \quad \Gamma_{72}/\Gamma$$

$$\Gamma_{72}/\Gamma = (\Gamma_{78} + \Gamma_{85} + \Gamma_{86} + \Gamma_{103} + \Gamma_{107} + 0.2302\Gamma_{150} + 0.2302\Gamma_{152} + 0.2302\Gamma_{154} + 0.892\Gamma_{178} + 0.892\Gamma_{179} + 0.9078\Gamma_{180}) / \Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>5.09 ± 0.05 OUR FIT</b>	(Produced by HFLAV)			
<b>5.10 ± 0.12 OUR AVERAGE</b>				

• • • We use the following data for averages but not for fits. • • •

5.106 ± 0.083 ± 0.103    10.1k    <sup>1</sup> ABDALLAH    06A    DLPH    1992–1995 LEP runs  
 5.09 ± 0.10 ± 0.23    <sup>2</sup> AKERS    95Y    OPAL    1991–1994 LEP runs

• • • We do not use the following data for averages, fits, limits, etc. • • •

4.95 ± 0.29 ± 0.65    570    DECAMP    92C    ALEP    Repl. by SCHAEEL 05C

<sup>1</sup> See footnote to ABDALLAH 06A  $\Gamma(\tau^- \rightarrow h^- \nu_\tau)/\Gamma_{\text{total}}$  measurement for correlations with other measurements.

<sup>2</sup> Not independent of AKERS 95Y  $B(h^- h^- h^+ \geq 0 \text{ neutrals } \nu_\tau \text{ (ex. } K_S^0 \rightarrow \pi^+ \pi^-))$  and  $B(h^- h^- h^+ \geq 0 \text{ neutrals } \nu_\tau \text{ (ex. } K^0)) / B(h^- h^- h^+ \geq 0 \text{ neutrals } \nu_\tau \text{ (ex. } K_S^0 \rightarrow \pi^+ \pi^-))$  values.

$$\Gamma(h^- h^- h^+ \pi^0 \nu_\tau) / \Gamma_{\text{total}} \quad \Gamma_{73} / \Gamma$$

$$\Gamma_{73} / \Gamma = (0.34598\Gamma_{41} + 0.34598\Gamma_{43} + \Gamma_{78} + \Gamma_{103} + \Gamma_{107} + 0.2302\Gamma_{152} + 0.892\Gamma_{178} + 0.892\Gamma_{179} + 0.0153\Gamma_{180}) / \Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>4.76 ± 0.05 OUR FIT</b>				(Produced by HFLAV)

• • • We do not use the following data for averages, fits, limits, etc. • • •

4.45 ± 0.09 ± 0.07    6.1k    <sup>1</sup> BUSKULIC    96    ALEP    Repl. by SCHAEEL 05C

<sup>1</sup> BUSKULIC 96 quote  $B(h^- h^- h^+ \pi^0 \nu_\tau \text{ (ex. } K^0)) = 4.30 \pm 0.09 \pm 0.09$ . We add 0.15 to remove their  $K^0$  correction and reduce the systematic error accordingly.

$$\Gamma(h^- h^- h^+ \pi^0 \nu_\tau \text{ (ex. } K^0)) / \Gamma_{\text{total}} \quad \Gamma_{74} / \Gamma$$

$$\Gamma_{74} / \Gamma = (\Gamma_{78} + \Gamma_{103} + \Gamma_{107} + 0.2302\Gamma_{152} + 0.892\Gamma_{178} + 0.892\Gamma_{179} + 0.0153\Gamma_{180}) / \Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>4.57 ± 0.05 OUR FIT</b>				(Produced by HFLAV)

**4.45 ± 0.14 OUR AVERAGE** Error includes scale factor of 1.2.

4.545 ± 0.106 ± 0.103    8.9k    <sup>1</sup> ABDALLAH    06A    DLPH    1992–1995 LEP runs  
 4.23 ± 0.06 ± 0.22    7.2k    BALEST    95C    CLEO     $E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$

<sup>1</sup> See footnote to ABDALLAH 06A  $\Gamma(\tau^- \rightarrow h^- \nu_\tau)/\Gamma_{\text{total}}$  measurement for correlations with other measurements.

$$\Gamma(h^- h^- h^+ \pi^0 \nu_\tau \text{ (ex. } K^0, \omega)) / \Gamma_{\text{total}} \quad \Gamma_{75} / \Gamma = (\Gamma_{78} + \Gamma_{103} + \Gamma_{107} + 0.2302\Gamma_{152}) / \Gamma$$

VALUE (%)	DOCUMENT ID
<b>2.79 ± 0.07 OUR FIT</b>	(Produced by HFLAV)

$$\Gamma(\pi^- \pi^+ \pi^- \pi^0 \nu_\tau) / \Gamma_{\text{total}} \quad \Gamma_{76} / \Gamma$$

$$\Gamma_{76} / \Gamma = (0.34598\Gamma_{41} + \Gamma_{78} + 0.892\Gamma_{178} + 0.0153\Gamma_{180}) / \Gamma$$

VALUE (%)	DOCUMENT ID
<b>4.62 ± 0.05 OUR FIT</b>	(Produced by HFLAV)

$$\Gamma(\pi^- \pi^+ \pi^- \pi^0 \nu_\tau \text{ (ex. } K^0)) / \Gamma_{\text{total}} \quad \Gamma_{77} / \Gamma$$

$$\Gamma_{77} / \Gamma = (\Gamma_{78} + 0.892\Gamma_{178} + 0.0153\Gamma_{180}) / \Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>4.49 ± 0.05 OUR FIT</b>				(Produced by HFLAV)

**4.55 ± 0.13 OUR AVERAGE** Error includes scale factor of 1.6.

4.598±0.057±0.064 16k <sup>1</sup> SCHAEEL 05C ALEP 1991-1995 LEP runs  
 4.19 ±0.10 ±0.21 <sup>2</sup> EDWARDS 00A CLEO 4.7 fb<sup>-1</sup> E<sub>cm</sub><sup>ee</sup> = 10.6 GeV

<sup>1</sup> SCHAEEL 05C quote (4.590±0.057±0.064)%. We add 0.008% to remove their correction for  $\tau^- \rightarrow \pi^- \pi^0 \omega \nu_\tau \rightarrow \pi^- \pi^0 \pi^+ \pi^- \nu_\tau$  decays. See footnote to SCHAEEL 05C

$\Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)/\Gamma_{\text{total}}$  measurement for correlations with other measurements.

<sup>2</sup> EDWARDS 00A quote (4.19 ± 0.10) × 10<sup>-2</sup> with a 5% systematic error.

$\Gamma(\pi^- \pi^+ \pi^- \pi^0 \nu_\tau (\text{ex. } K^0, \omega))/\Gamma_{\text{total}}$   $\Gamma_{78}/\Gamma$   
VALUE (%) DOCUMENT ID  
**2.74±0.07 OUR FIT** (Produced by HFLAV)

$\Gamma(h^- \rho \pi^0 \nu_\tau)/\Gamma(h^- h^- h^+ \pi^0 \nu_\tau)$   $\Gamma_{79}/\Gamma_{73}$   
VALUE EVTS DOCUMENT ID TECN COMMENT

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.30±0.04±0.02 393 ALBRECHT 91D ARG E<sub>cm</sub><sup>ee</sup> = 9.4–10.6 GeV

$\Gamma(h^- \rho^+ h^- \nu_\tau)/\Gamma(h^- h^- h^+ \pi^0 \nu_\tau)$   $\Gamma_{80}/\Gamma_{73}$   
VALUE EVTS DOCUMENT ID TECN COMMENT

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.10±0.03±0.04 142 ALBRECHT 91D ARG E<sub>cm</sub><sup>ee</sup> = 9.4–10.6 GeV

$\Gamma(h^- \rho^- h^+ \nu_\tau)/\Gamma(h^- h^- h^+ \pi^0 \nu_\tau)$   $\Gamma_{81}/\Gamma_{73}$   
VALUE EVTS DOCUMENT ID TECN COMMENT

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.26±0.05±0.01 370 ALBRECHT 91D ARG E<sub>cm</sub><sup>ee</sup> = 9.4–10.6 GeV

$\Gamma(h^- h^- h^+ \geq 2\pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$   $\Gamma_{82}/\Gamma$   
 $\Gamma_{82}/\Gamma = (\Gamma_{85} + \Gamma_{86} + 0.2302\Gamma_{150} + 0.2302\Gamma_{154} + 0.892\Gamma_{180})/\Gamma$

VALUE (%) EVTS DOCUMENT ID TECN COMMENT  
**0.517±0.031 OUR FIT** (Produced by HFLAV)

**0.561±0.068±0.095** 1.3k <sup>1</sup> ABDALLAH 06A DLPH 1992–1995 LEP runs

<sup>1</sup> See footnote to ABDALLAH 06A  $\Gamma(\tau^- \rightarrow h^- \nu_\tau)/\Gamma_{\text{total}}$  measurement for correlations with other measurements.

$\Gamma(h^- h^- h^+ 2\pi^0 \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{83}/\Gamma$   
 $\Gamma_{83}/\Gamma = (0.4247\Gamma_{48} + \Gamma_{85} + 0.2302\Gamma_{150} + 0.2302\Gamma_{154} + 0.892\Gamma_{180})/\Gamma$

VALUE (%) DOCUMENT ID  
**0.505±0.031 OUR FIT** (Produced by HFLAV)

$\Gamma(h^- h^- h^+ 2\pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$   $\Gamma_{84}/\Gamma$   
 $\Gamma_{84}/\Gamma = (\Gamma_{85} + 0.2302\Gamma_{150} + 0.2302\Gamma_{154} + 0.892\Gamma_{180})/\Gamma$

VALUE (%) EVTS DOCUMENT ID TECN COMMENT  
**0.495±0.031 OUR FIT** (Produced by HFLAV)

**0.435±0.030±0.035** 2.6k <sup>1</sup> SCHAEEL 05C ALEP 1991-1995 LEP runs

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.50 ± 0.07 ± 0.07      1.8k      BUSKULIC      96      ALEP      Repl. by SCHAEEL 05C

<sup>1</sup>SCHAEEL 05C quote (0.392 ± 0.030 ± 0.035)%. We add 0.043% to remove their correction for  $\tau^- \rightarrow \pi^- \eta \pi^0 \nu_\tau \rightarrow \pi^- \pi^+ \pi^- 2\pi^0 \nu_\tau$  and  $\tau^- \rightarrow K^*(892)^- \eta \nu_\tau \rightarrow K^- \pi^+ \pi^- 2\pi^0 \nu_\tau$  decays. See footnote to SCHAEEL 05C  $\Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)/\Gamma_{\text{total}}$  measurement for correlations with other measurements.

$\Gamma(h^- h^- h^+ 2\pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma(h^- h^- h^+ \geq 0 \text{ neutrals} \geq 0 K_L^0 \nu_\tau)$        $\Gamma_{84}/\Gamma_{62}$

$$\Gamma_{84}/\Gamma_{62} = (\Gamma_{85} + 0.2302\Gamma_{150} + 0.2302\Gamma_{154} + 0.892\Gamma_{180}) / (0.34598\Gamma_{36} + 0.34598\Gamma_{38} + 0.34598\Gamma_{41} + 0.34598\Gamma_{43} + 0.4247\Gamma_{48} + 0.6920\Gamma_{49} + 0.8494\Gamma_{52} + 0.6920\Gamma_{56} + 0.6534\Gamma_{61} + \Gamma_{70} + \Gamma_{78} + \Gamma_{85} + \Gamma_{89} + \Gamma_{97} + \Gamma_{103} + \Gamma_{106} + \Gamma_{107} + 0.2804\Gamma_{150} + 0.2302\Gamma_{151} + 0.2804\Gamma_{152} + 0.2804\Gamma_{154} + 0.3759\Gamma_{156} + 0.3257\Gamma_{160} + 0.7259\Gamma_{170} + 0.9078\Gamma_{178} + 0.9078\Gamma_{179} + 0.9078\Gamma_{180} + 0.892\Gamma_{182})$$

VALUE (units 10 <sup>-2</sup> )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>3.26 ± 0.20 OUR FIT</b>		(Produced by HFLAV)		
<b>3.4 ± 0.2 ± 0.3</b>	668	BORTOLETTO93	CLEO	$E_{\text{cm}}^{ee} \approx 10.6 \text{ GeV}$

$\Gamma(h^- h^- h^+ 2\pi^0 \nu_\tau (\text{ex. } K^0, \omega, \eta))/\Gamma_{\text{total}}$        $\Gamma_{85}/\Gamma$

VALUE (units 10 <sup>-4</sup> )	DOCUMENT ID
<b>10 ± 4 OUR FIT</b>	(Produced by HFLAV)

$\Gamma(h^- h^- h^+ 3\pi^0 \nu_\tau)/\Gamma_{\text{total}}$        $\Gamma_{86}/\Gamma = (0.4247\Gamma_{52} + \Gamma_{87} + 0.1131\Gamma_{156})/\Gamma$

VALUE (units 10 <sup>-4</sup> )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2.13 ± 0.30 OUR FIT</b>			(Produced by HFLAV)		
<b>2.2 ± 0.3 ± 0.4</b>		139	ANASTASSOV 01	CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

< 4.9      95      SCHAEEL      05C      ALEP      1991-1995 LEP runs  
 2.85 ± 0.56 ± 0.51      57      ANDERSON      97      CLEO      Repl. by ANAS-TASSOV 01  
 11 ± 4 ± 5      440      <sup>1</sup>BUSKULIC      96      ALEP      Repl. by SCHAEEL 05C

<sup>1</sup>BUSKULIC 96 state their measurement is for  $B(h^- h^- h^+ \geq 3\pi^0 \nu_\tau)$ . We assume that  $B(h^- h^- h^+ \geq 4\pi^0 \nu_\tau)$  is very small.

$\Gamma(2\pi^- \pi^+ 3\pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$        $\Gamma_{87}/\Gamma$

$$\Gamma_{87}/\Gamma = (\Gamma_{89} + 0.2302\Gamma_{151} + 0.3257\Gamma_{160} + 0.892\Gamma_{182})/\Gamma$$

VALUE (units 10 <sup>-4</sup> )	DOCUMENT ID	TECN	COMMENT
<b>1.94 ± 0.30 OUR FIT</b>	(Produced by HFLAV)		

• • • We use the following data for averages but not for fits. • • •

**2.07 ± 0.18 ± 0.37**      <sup>1</sup>LEES      12X      BABR      468 fb<sup>-1</sup>       $E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

<sup>1</sup>Not independent of LEES 12X  $\Gamma(\tau^- \rightarrow \eta \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0))/\Gamma$ ,  $\Gamma(\tau^- \rightarrow \eta \pi^- \pi^0 \pi^0 \nu_\tau)/\Gamma$ ,  $\Gamma(\tau^- \rightarrow \pi^- \omega 2\pi^0 \nu_\tau)/\Gamma$ , and  $\Gamma(\tau^- \rightarrow f_1(1285) \pi^- \nu_\tau \rightarrow \eta \pi^- \pi^+ \pi^- \nu_\tau)/\Gamma$  values.

$\Gamma(2\pi^- \pi^+ 3\pi^0 \nu_\tau (\text{ex. } K^0, \eta, f_1(1285)))/\Gamma_{\text{total}}$        $\Gamma_{88}/\Gamma$

VALUE (units 10 <sup>-4</sup> )	DOCUMENT ID	TECN	COMMENT
<b>1.69 ± 0.08 ± 0.43</b>	LEES	12X	BABR 468 fb <sup>-1</sup> $E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

$$\Gamma(2\pi^- \pi^+ 3\pi^0 \nu_\tau (\text{ex. } K^0, \eta, \omega, f_1(1285))) / \Gamma_{\text{total}} \quad \Gamma_{89} / \Gamma$$

VALUE (units $10^{-5}$ )	DOCUMENT ID	TECN	COMMENT
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**1.4 ± 2.7 OUR FIT** (Produced by HFLAV)

**1.0 ± 0.8 ± 3.0** <sup>1</sup> LEES 12X BABR 468 fb<sup>-1</sup>  $E_{\text{cm}}^{\text{ee}} = 10.6$  GeV

<sup>1</sup> LEES 12X measurement corresponds to the lower limit of  $< 5.8 \times 10^{-5}$  at 90% CL.

$$\Gamma(K^- h^+ h^- \geq 0 \text{ neutrals } \nu_\tau) / \Gamma_{\text{total}} \quad \Gamma_{90} / \Gamma$$

$$\Gamma_{90} / \Gamma = (0.34598\Gamma_{38} + 0.34598\Gamma_{43} + \Gamma_{97} + \Gamma_{103} + \Gamma_{106} + \Gamma_{107} + 0.2804\Gamma_{152} + 0.491\Gamma_{170} + 0.9078\Gamma_{179}) / \Gamma$$

VALUE (%)	CL%	DOCUMENT ID	TECN	COMMENT
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**0.629 ± 0.014 OUR FIT** (Produced by HFLAV)

**< 0.6** 90 AIHARA 84c TPC  $E_{\text{cm}}^{\text{ee}} = 29$  GeV

$$\Gamma(K^- h^+ \pi^- \nu_\tau (\text{ex. } K^0)) / \Gamma_{\text{total}} \quad \Gamma_{91} / \Gamma = (\Gamma_{97} + \Gamma_{106} + 0.0153\Gamma_{179}) / \Gamma$$

VALUE (%)	DOCUMENT ID
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**0.437 ± 0.007 OUR FIT** (Produced by HFLAV)

$$\Gamma(K^- h^+ \pi^- \nu_\tau (\text{ex. } K^0)) / \Gamma(\pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0)) \quad \Gamma_{91} / \Gamma_{68}$$

$$\Gamma_{91} / \Gamma_{68} = (\Gamma_{97} + \Gamma_{106} + 0.0153\Gamma_{179}) / (\Gamma_{70} + 0.0153\Gamma_{178})$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
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**4.84 ± 0.08 OUR FIT** (Produced by HFLAV)

**5.44 ± 0.21 ± 0.53** 7.9k RICHICHI 99 CLEO  $E_{\text{cm}}^{\text{ee}} = 10.6$  GeV

$$\Gamma(K^- h^+ \pi^- \pi^0 \nu_\tau (\text{ex. } K^0)) / \Gamma_{\text{total}} \quad \Gamma_{92} / \Gamma$$

$$\Gamma_{92} / \Gamma = (\Gamma_{103} + \Gamma_{107} + 0.2302\Gamma_{152} + 0.892\Gamma_{179}) / \Gamma$$

VALUE (units $10^{-4}$ )	DOCUMENT ID
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**8.6 ± 1.2 OUR FIT** (Produced by HFLAV)

$$\Gamma(K^- h^+ \pi^- \pi^0 \nu_\tau (\text{ex. } K^0)) / \Gamma(\pi^- \pi^+ \pi^- \pi^0 \nu_\tau (\text{ex. } K^0)) \quad \Gamma_{92} / \Gamma_{77}$$

$$\Gamma_{92} / \Gamma_{77} = (\Gamma_{103} + \Gamma_{107} + 0.2302\Gamma_{152} + 0.892\Gamma_{179}) / (\Gamma_{78} + 0.892\Gamma_{178} + 0.0153\Gamma_{180})$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
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**1.91 ± 0.26 OUR FIT** (Produced by HFLAV)

**2.61 ± 0.45 ± 0.42** 719 RICHICHI 99 CLEO  $E_{\text{cm}}^{\text{ee}} = 10.6$  GeV

$$\Gamma(K^- \pi^+ \pi^- \geq 0 \text{ neutrals } \nu_\tau) / \Gamma_{\text{total}} \quad \Gamma_{93} / \Gamma$$

$$\Gamma_{93} / \Gamma = (0.34598\Gamma_{38} + 0.34598\Gamma_{43} + \Gamma_{97} + \Gamma_{103} + 0.2804\Gamma_{152} + 0.9078\Gamma_{179}) / \Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
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**0.477 ± 0.014 OUR FIT** (Produced by HFLAV)

**0.58 <sup>+0.15</sup> <sub>-0.13</sub> ± 0.12** 20 <sup>1</sup> BAUER 94 TPC  $E_{\text{cm}}^{\text{ee}} = 29$  GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.22 <sup>+0.16</sup> <sub>-0.13</sub> ± 0.05 9 <sup>2</sup> MILLS 85 DLCO  $E_{\text{cm}}^{\text{ee}} = 29$  GeV

<sup>1</sup> We multiply 0.58% by 0.20, the relative systematic error quoted by BAUER 94, to obtain the systematic error.

<sup>2</sup> Error correlated with MILLS 85 ( $K K \pi \nu$ ) value. We multiply 0.22% by 0.23, the relative systematic error quoted by MILLS 85, to obtain the systematic error.

$$\Gamma(K^- \pi^+ \pi^- \geq 0 \pi^0 \nu_\tau \text{ (ex. } K^0)) / \Gamma_{\text{total}} \quad \Gamma_{94} / \Gamma$$

$$\Gamma_{94} / \Gamma = (\Gamma_{97} + \Gamma_{103} + 0.2302 \Gamma_{152} + 0.9078 \Gamma_{179}) / \Gamma$$

VALUE (%)	DOCUMENT ID	TECN	COMMENT
<b>0.373 ± 0.013 OUR FIT</b>	(Produced by HFLAV)		
<b>0.30 ± 0.05 OUR AVERAGE</b>			

• • • We use the following data for averages but not for fits. • • •

0.343 ± 0.073 ± 0.031	ABBIENDI	00D	OPAL	1990–1995 LEP runs
0.275 ± 0.064	<sup>1</sup> BARATE	98	ALEP	1991–1995 LEP runs

<sup>1</sup> Not independent of BARATE 98  $\Gamma(\tau^- \rightarrow K^- \pi^+ \pi^- \nu_\tau) / \Gamma_{\text{total}}$  and  $\Gamma(\tau^- \rightarrow K^- \pi^+ \pi^- \pi^0 \nu_\tau) / \Gamma_{\text{total}}$  values.

$$\Gamma(K^- \pi^+ \pi^- \nu_\tau) / \Gamma_{\text{total}} \quad \Gamma_{95} / \Gamma = (0.34598 \Gamma_{38} + \Gamma_{97} + 0.0153 \Gamma_{179}) / \Gamma$$

VALUE (%)	DOCUMENT ID
<b>0.345 ± 0.007 OUR FIT</b>	(Produced by HFLAV)

$$\Gamma(K^- \pi^+ \pi^- \nu_\tau \text{ (ex. } K^0)) / \Gamma_{\text{total}} \quad \Gamma_{96} / \Gamma = (\Gamma_{97} + 0.0153 \Gamma_{179}) / \Gamma$$

VALUE (%)	EVTs	DOCUMENT ID	TECN	COMMENT
<b>0.293 ± 0.007 OUR FIT</b>	(Produced by HFLAV)			
<b>0.290 ± 0.018 OUR AVERAGE</b>	Error includes scale factor of 2.4. See the ideogram below.			

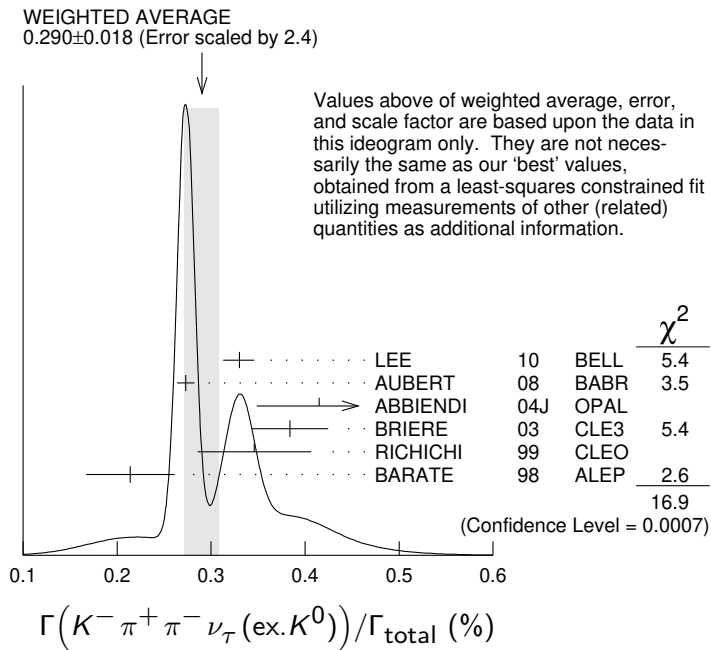
0.330 ± 0.001 <sup>+0.016</sup> <sub>-0.017</sub>	794k	<sup>1</sup> LEE	10	BELL	666 fb <sup>-1</sup>	$E_{\text{cm}}^{ee} = 10.6$ GeV
0.273 ± 0.002 ± 0.009	70k	<sup>2</sup> AUBERT	08	BABR	342 fb <sup>-1</sup>	$E_{\text{cm}}^{ee} = 10.6$ GeV
0.415 ± 0.053 ± 0.040	269	ABBIENDI	04J	OPAL	1991–1995 LEP runs	
0.384 ± 0.014 ± 0.038	3.5k	<sup>3</sup> BRIERE	03	CLE3	$E_{\text{cm}}^{ee} = 10.6$ GeV	
0.214 ± 0.037 ± 0.029		BARATE	98	ALEP	1991–1995 LEP runs	

• • • We use the following data for averages but not for fits. • • •

0.346 ± 0.023 ± 0.056	158	<sup>4</sup> RICHICHI	99	CLEO	$E_{\text{cm}}^{ee} = 10.6$ GeV	
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• • • We do not use the following data for averages, fits, limits, etc. • • •

0.360 ± 0.082 ± 0.048	ABBIENDI	00D	OPAL	1990–1995 LEP runs		
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<sup>1</sup> See footnote to LEE 10  $\Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$  measurement for correlations with other measurements. Not independent of LEE 10  $\Gamma(\tau^- \rightarrow K^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0))/\Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0))$  value.

<sup>2</sup> See footnote to AUBERT 08  $\Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$  measurement for correlations with other measurements.

<sup>3</sup> 47% correlated with BRIERE 03  $\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau$  and 34% correlated with  $\tau^- \rightarrow K^- K^+ \pi^- \nu_\tau$  because of a common 5% normalization error.

<sup>4</sup> Not independent of RICHICHI 99

$\Gamma(\tau^- \rightarrow K^- h^+ \pi^- \nu_\tau (\text{ex. } K^0))/\Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0))$ ,  $\Gamma(\tau^- \rightarrow K^- K^+ \pi^- \nu_\tau)/\Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0))$  and BALEST 95c  $\Gamma(\tau^- \rightarrow h^- h^- h^+ \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$  values.

$$\Gamma(K^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0))/\Gamma(\pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0)) \quad \Gamma_{96}/\Gamma_{68}$$

$$\Gamma_{96}/\Gamma_{68} = (\Gamma_{97} + 0.0153\Gamma_{179})/(\Gamma_{70} + 0.0153\Gamma_{178})$$

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>3.25 ± 0.07 OUR FIT</b>		(Produced by HFLAV)		

• • • We use the following data for averages but not for fits. • • •

<b>3.92 ± 0.02<sup>+0.15</sup><sub>-0.16</sub></b>	794k	<sup>1</sup> LEE	10 BELL	666 fb <sup>-1</sup> $E_{\text{cm}}^{ee} = 10.6$ GeV
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<sup>1</sup> Not independent of LEE 10  $\Gamma(\tau^- \rightarrow K^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$  and  $\Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$  values.

$$\Gamma(K^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0, \omega))/\Gamma_{\text{total}} \quad \Gamma_{97}/\Gamma$$

VALUE (units $10^{-3}$ )	DOCUMENT ID
<b>2.93 ± 0.07 OUR FIT</b>	(Produced by HFLAV)

$$\Gamma(K^- \rho^0 \nu_\tau \rightarrow K^- \pi^+ \pi^- \nu_\tau)/\Gamma(K^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0)) \quad \Gamma_{98}/\Gamma_{96}$$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.48 ± 0.14 ± 0.10</b>	<sup>1</sup> ASNER	00B CLEO	$E_{\text{cm}}^{ee} = 10.6$ GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.39 ± 0.14	<sup>2</sup> BARATE	99R ALEP	1991–1995 LEP runs
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<sup>1</sup> ASNER 00B assume  $\tau^- \rightarrow K^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0)$  decays proceed only through  $K\rho$  and  $K^* \pi$  intermediate states. They assume the resonance structure of  $\tau^- \rightarrow K^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0)$  decays is dominated by  $K_1(1270)^-$  and  $K_1(1400)^-$  resonances, and assume  $B(K_1(1270) \rightarrow K^*(892)\pi) = (16 \pm 5)\%$ ,  $B(K_1(1270) \rightarrow K\rho) = (42 \pm 6)\%$ , and  $B(K_1(1400) \rightarrow K\rho) = 0$ .

<sup>2</sup> BARATE 99R assume  $\tau^- \rightarrow K^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0)$  decays proceed only through  $K\rho$  and  $K^* \pi$  intermediate states. The quoted error is statistical only.

$$\Gamma(K^- \pi^+ \pi^- \pi^0 \nu_\tau)/\Gamma_{\text{total}} \quad \Gamma_{99}/\Gamma$$

$$\Gamma_{99}/\Gamma = (0.34598\Gamma_{43} + \Gamma_{103} + 0.2302\Gamma_{152} + 0.892\Gamma_{179})/\Gamma$$

VALUE (units $10^{-4}$ )	DOCUMENT ID
<b>13.1 ± 1.2 OUR FIT</b>	(Produced by HFLAV)



$$\Gamma(K^- \pi^+ \pi^- \pi^0 \nu_\tau (\text{ex. } K^0)) / \Gamma_{\text{total}} \quad \Gamma_{100} / \Gamma$$

$$\Gamma_{100} / \Gamma = (\Gamma_{103} + 0.2302 \Gamma_{152} + 0.892 \Gamma_{179}) / \Gamma$$

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
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**7.9 ± 1.2 OUR FIT** (Produced by HFLAV)

**7.3 ± 1.2 OUR AVERAGE**

7.4 ± 0.8 ± 1.1		<sup>1</sup> ARMS	05	CLE3	7.6 fb <sup>-1</sup> , $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
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6.1 ± 3.9 ± 1.8		BARATE	98	ALEP	1991–1995 LEP runs
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• • • We use the following data for averages but not for fits. • • •

7.5 ± 2.6 ± 1.8		<sup>2</sup> RICHICHI	99	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
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• • • We do not use the following data for averages, fits, limits, etc. • • •

<17	95	ABBIENDI	00D	OPAL	1990–1995 LEP runs
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<sup>1</sup> Not independent of ARMS 05  $\Gamma(\tau^- \rightarrow K^- \pi^+ \pi^- \pi^0 \nu_\tau (\text{ex. } K^0, \omega)) / \Gamma_{\text{total}}$  and  $\Gamma(\tau^- \rightarrow K^- \omega \nu_\tau) / \Gamma_{\text{total}}$  values.

<sup>2</sup> Not independent of RICHICHI 99

$\Gamma(\tau^- \rightarrow K^- h^+ \pi^- \nu_\tau (\text{ex. } K^0)) / \Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0))$ ,  $\Gamma(\tau^- \rightarrow K^- K^+ \pi^- \nu_\tau) / \Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0))$  and BALEST 95C  $\Gamma(\tau^- \rightarrow h^- h^- h^+ \nu_\tau (\text{ex. } K^0)) / \Gamma_{\text{total}}$  values.

$$\Gamma(K^- \pi^+ \pi^- \pi^0 \nu_\tau (\text{ex. } K^0, \eta)) / \Gamma_{\text{total}} \quad \Gamma_{101} / \Gamma = (\Gamma_{103} + 0.892 \Gamma_{179}) / \Gamma$$

VALUE (units $10^{-4}$ )	DOCUMENT ID
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**7.6 ± 1.2 OUR FIT** (Produced by HFLAV)

$$\Gamma(K^- \pi^+ \pi^- \pi^0 \nu_\tau (\text{ex. } K^0, \omega)) / \Gamma_{\text{total}} \quad \Gamma_{102} / \Gamma$$

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
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<b>3.7 ± 0.5 ± 0.8</b>	833	ARMS	05	CLE3	7.6 fb <sup>-1</sup> , $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
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$$\Gamma(K^- \pi^+ \pi^- \pi^0 \nu_\tau (\text{ex. } K^0, \omega, \eta)) / \Gamma_{\text{total}} \quad \Gamma_{103} / \Gamma$$

VALUE (units $10^{-4}$ )	DOCUMENT ID
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**3.9 ± 1.4 OUR FIT** (Produced by HFLAV)

$$\Gamma(K^- \pi^+ K^- \geq 0 \text{ neut. } \nu_\tau) / \Gamma_{\text{total}} \quad \Gamma_{104} / \Gamma$$

VALUE (%)	CL%	DOCUMENT ID	TECN	COMMENT
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<b>&lt;0.09</b>	95	BAUER	94	TPC	$E_{\text{cm}}^{\text{ee}} = 29$ GeV
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$$\Gamma(K^- K^+ \pi^- \geq 0 \text{ neut. } \nu_\tau) / \Gamma_{\text{total}} \quad \Gamma_{105} / \Gamma = (\Gamma_{106} + \Gamma_{107}) / \Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
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**0.1496 ± 0.0033 OUR FIT** (Produced by HFLAV)

**0.203 ± 0.031 OUR AVERAGE**

0.159 ± 0.053 ± 0.020		ABBIENDI	00D	OPAL	1990–1995 LEP runs
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0.15 $\begin{smallmatrix} +0.09 \\ -0.07 \end{smallmatrix}$ ± 0.03	4	<sup>1</sup> BAUER	94	TPC	$E_{\text{cm}}^{\text{ee}} = 29$ GeV
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• • • We use the following data for averages but not for fits. • • •

0.238 ± 0.042		<sup>2</sup> BARATE	98	ALEP	1991–1995 LEP runs
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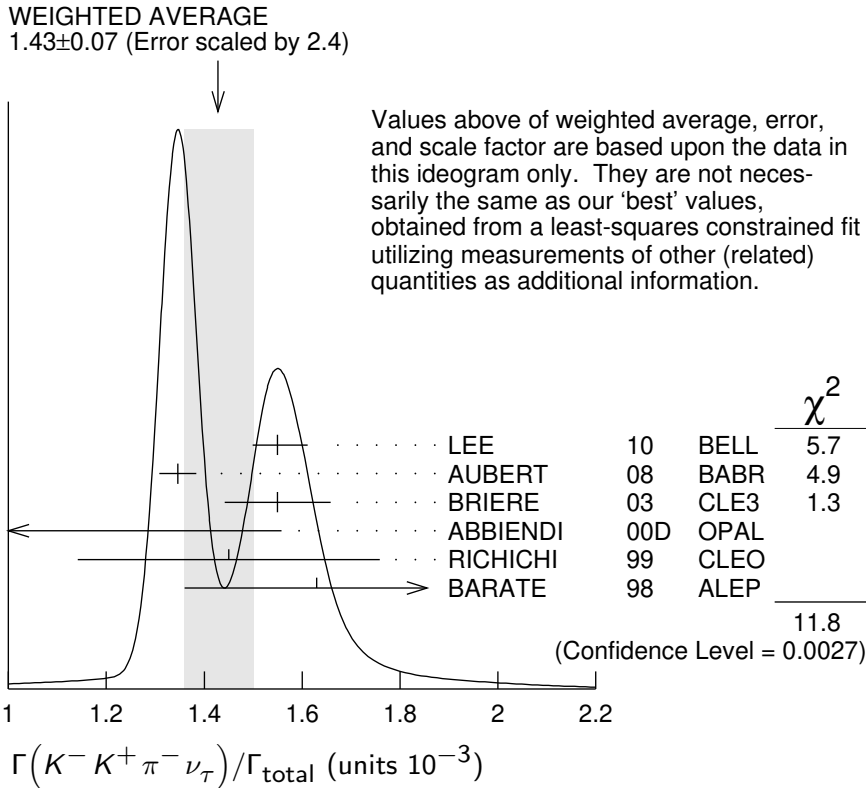
<sup>1</sup> We multiply 0.15% by 0.20, the relative systematic error quoted by BAUER 94, to obtain the systematic error.

<sup>2</sup> Not independent of BARATE 98  $\Gamma(\tau^- \rightarrow K^- K^+ \pi^- \nu_\tau) / \Gamma_{\text{total}}$  and  $\Gamma(\tau^- \rightarrow K^- K^+ \pi^- \pi^0 \nu_\tau) / \Gamma_{\text{total}}$  values.

$\Gamma(K^- K^+ \pi^- \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{106}/\Gamma$

VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.435 ± 0.027 OUR FIT</b>		(Produced by HFLAV)		
<b>1.43 ± 0.07 OUR AVERAGE</b>		Error includes scale factor of 2.4. See the ideogram below.		
1.55 ± 0.01 <sup>+0.06</sup> / <sub>-0.05</sub>	108k	<sup>1</sup> LEE	10	BELL 666 fb <sup>-1</sup> E <sub>cm</sub> <sup>ee</sup> = 10.6 GeV
1.346 ± 0.010 ± 0.036	18k	<sup>2</sup> AUBERT	08	BABR 342 fb <sup>-1</sup> E <sub>cm</sub> <sup>ee</sup> = 10.6 GeV
1.55 ± 0.06 ± 0.09	932	<sup>3</sup> BRIERE	03	CLE3 E <sub>cm</sub> <sup>ee</sup> = 10.6 GeV
1.63 ± 0.21 ± 0.17		BARATE	98	ALEP 1991–1995 LEP runs
● ● ● We use the following data for averages but not for fits. ● ● ●				
0.87 ± 0.56 ± 0.40		ABBIENDI	00D	OPAL 1990–1995 LEP runs
1.45 ± 0.13 ± 0.28	2.3k	<sup>4</sup> RICHICHI	99	CLEO E <sub>cm</sub> <sup>ee</sup> = 10.6 GeV
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
2.2 <sup>+1.7</sup> / <sub>-1.1</sub> ± 0.5	9	<sup>5</sup> MILLS	85	DLCO E <sub>cm</sub> <sup>ee</sup> = 29 GeV

<sup>1</sup> See footnote to LEE 10  $\Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$  measurement for correlations with other measurements. Not independent of LEE 10  $\Gamma(\tau^- \rightarrow K^- K^+ \pi^- \nu_\tau)/\Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0))$  value.  
<sup>2</sup> See footnote to AUBERT 08  $\Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$  measurement for correlations with other measurements.  
<sup>3</sup> 71% correlated with BRIERE 03  $\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau$  and 34% correlated with  $\tau \rightarrow K^- \pi^+ \pi^- \nu_\tau$  because of a common 5% normalization error.  
<sup>4</sup> Not independent of RICHICHI 99  $\Gamma(\tau^- \rightarrow K^- K^+ \pi^- \nu_\tau)/\Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0))$  and BALEST 95c  $\Gamma(\tau^- \rightarrow h^- h^- h^+ \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$  values.  
<sup>5</sup> Error correlated with MILLS 85 ( $K \pi \pi \pi^0 \nu$ ) value. We multiply 0.22% by 0.23, the relative systematic error quoted by MILLS 85, to obtain the systematic error.



$$\Gamma(K^- K^+ \pi^- \nu_\tau) / \Gamma(\pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0)) \quad \Gamma_{106} / \Gamma_{68}$$

$$\Gamma_{106} / \Gamma_{68} = \Gamma_{106} / (\Gamma_{70} + 0.0153 \Gamma_{178})$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.592 ± 0.030 OUR FIT</b>				(Produced by HFLAV)

**1.83 ± 0.05 OUR AVERAGE**

1.60 ± 0.15 ± 0.30 2.3k RICHICHI 99 CLEO  $E_{\text{cm}}^{ee} = 10.6$  GeV

• • • We use the following data for averages but not for fits. • • •

1.84 ± 0.01 ± 0.05 108k <sup>1</sup> LEE 10 BELL 666 fb<sup>-1</sup>  $E_{\text{cm}}^{ee} = 10.6$  GeV

<sup>1</sup>Not independent of LEE 10  $\Gamma(\tau^- \rightarrow K^- K^+ \pi^- \nu_\tau) / \Gamma_{\text{total}}$  and  $\Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0)) / \Gamma_{\text{total}}$  values.

$$\Gamma(K^- K^+ \pi^- \pi^0 \nu_\tau) / \Gamma_{\text{total}} \quad \Gamma_{107} / \Gamma$$

VALUE (units 10 <sup>-4</sup> )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.61 ± 0.18 OUR FIT</b>					(Produced by HFLAV)

**0.60 ± 0.18 OUR AVERAGE**

0.55 ± 0.14 ± 0.12 48 ARMS 05 CLE3 7.6 fb<sup>-1</sup>,  $E_{\text{cm}}^{ee} = 10.6$  GeV

7.5 ± 2.9 ± 1.5 BARATE 98 ALEP 1991–1995 LEP runs

• • • We use the following data for averages but not for fits. • • •

3.3 ± 1.8 ± 0.7 158 <sup>1</sup> RICHICHI 99 CLEO  $E_{\text{cm}}^{ee} = 10.6$  GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

<27 95 ABBIENDI 00D OPAL 1990–1995 LEP runs

<sup>1</sup>Not independent of RICHICHI 99

$\Gamma(\tau^- \rightarrow K^- K^+ \pi^- \nu_\tau) / \Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0))$  and BALEST 95C  $\Gamma(\tau^- \rightarrow h^- h^- h^+ \nu_\tau (\text{ex. } K^0)) / \Gamma_{\text{total}}$  values.

$$\Gamma(K^- K^+ \pi^- \pi^0 \nu_\tau) / \Gamma(\pi^- \pi^+ \pi^- \pi^0 \nu_\tau (\text{ex. } K^0)) \quad \Gamma_{107} / \Gamma_{77}$$

$$\Gamma_{107} / \Gamma_{77} = \Gamma_{107} / (\Gamma_{78} + 0.892 \Gamma_{178} + 0.0153 \Gamma_{180})$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.14 ± 0.04 OUR FIT</b>				(Produced by HFLAV)

**0.79 ± 0.44 ± 0.16** 158 <sup>1</sup> RICHICHI 99 CLEO  $E_{\text{cm}}^{ee} = 10.6$  GeV

<sup>1</sup>RICHICHI 99 also quote a 95%CL upper limit of 0.0157 for this measurement.

$$\Gamma(K^- K^+ K^- \nu_\tau) / \Gamma_{\text{total}} \quad \Gamma_{108} / \Gamma = 0.491 \Gamma_{170} / \Gamma$$

VALUE (units 10 <sup>-5</sup> )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2.2 ± 0.8 OUR FIT</b>					Error includes scale factor of 5.4. (Produced by HFLAV)

**2.1 ± 0.8 OUR AVERAGE** Error includes scale factor of 5.4.

3.29 ± 0.17 <sup>+0.19</sup>/<sub>-0.20</sub> 3.2k <sup>1</sup> LEE 10 BELL 666 fb<sup>-1</sup>  $E_{\text{cm}}^{ee} = 10.6$  GeV

1.58 ± 0.13 ± 0.12 275 <sup>2</sup> AUBERT 08 BABR 342 fb<sup>-1</sup>  $E_{\text{cm}}^{ee} = 10.6$  GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

< 3.7 90 BRIERE 03 CLE3  $E_{\text{cm}}^{ee} = 10.6$  GeV

< 19 90 BARATE 98 ALEP 1991–1995 LEP runs

<sup>1</sup>See footnote to LEE 10  $\Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0)) / \Gamma_{\text{total}}$  measurement for correlations with other measurements. Not independent of LEE 10  $\Gamma(\tau^- \rightarrow K^- K^+ K^- \nu_\tau) / \Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0))$  value.

<sup>2</sup>See footnote to AUBERT 08  $\Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0)) / \Gamma_{\text{total}}$  measurement for correlations with other measurements.

$\Gamma(K^- K^+ K^- \nu_\tau)/\Gamma(\pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0))$   $\Gamma_{108}/\Gamma_{68}$ 

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
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• • • We do not use the following data for averages, fits, limits, etc. • • •

$3.90 \pm 0.02^{+0.22}_{-0.23}$	3.2k	<sup>1</sup> LEE	10 BELL	$666 \text{ fb}^{-1}$ , $E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
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<sup>1</sup> Not independent of LEE 10  $\Gamma(\tau^- \rightarrow K^- K^+ K^- \nu_\tau)/\Gamma_{\text{total}}$  and  $\Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$  values. $\Gamma(K^- K^+ K^- \nu_\tau (\text{ex. } \phi))/\Gamma_{\text{total}}$   $\Gamma_{109}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
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$< 2.5 \times 10^{-6}$	90	AUBERT	08 BABR	$342 \text{ fb}^{-1}$ , $E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
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 $\Gamma(K^- K^+ K^- \pi^0 \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{110}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
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$< 4.8 \times 10^{-6}$	90	ARMS	05 CLE3	$7.6 \text{ fb}^{-1}$ , $E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
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 $\Gamma(\pi^- K^+ \pi^- \geq 0 \text{ neut. } \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{111}/\Gamma$ 

VALUE (%)	CL%	DOCUMENT ID	TECN	COMMENT
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$< 0.25$	95	BAUER	94 TPC	$E_{\text{cm}}^{ee} = 29 \text{ GeV}$
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 $\Gamma(e^- e^- e^+ \bar{\nu}_e \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{112}/\Gamma$ 

VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
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$2.8 \pm 1.4 \pm 0.4$	5	ALAM	96 CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
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 $\Gamma(\mu^- e^- e^+ \bar{\nu}_\mu \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{113}/\Gamma$ 

VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT
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$< 3.2$	90	ALAM	96 CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
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 $\Gamma(\pi^- e^- e^+ \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{114}/\Gamma$ 

VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
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<b>seen</b>	400	<sup>1</sup> JIN	19 BELL	$562 \text{ fb}^{-1}$ , $E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
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• • • We do not use the following data for averages, fits, limits, etc. • • •

$1.46 \pm 0.13 \pm 0.21$	400	<sup>1</sup> JIN	19 BELL	axial-vector, $562 \text{ fb}^{-1}$ , $E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
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$3.01 \pm 0.27 \pm 0.43$	400	<sup>1</sup> JIN	19 BELL	vector, $562 \text{ fb}^{-1}$ , $E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
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<sup>1</sup> JIN 19 measures  $B(\tau^- \rightarrow \pi^- e^- e^+ \nu_\tau (m_{\pi^- e^- e^+} > 1.05 \text{ GeV}/c^2)) = (5.90 \pm 0.53 \pm 0.86) \times 10^{-6}$ , which is only sensitive to the structure-dependent contribution, and assumes that the decay proceeds with either a pure axial-vector current or a pure vector current to obtain the two respective branching fraction measurements for this mode, which are 100% correlated. $\Gamma(\pi^- \mu^- \mu^+ \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{115}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
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$< 1.14 \times 10^{-5}$	90	JIN	19 BELL	$562 \text{ fb}^{-1}$ , $E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
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$$\Gamma(3h^-2h^+ \geq 0 \text{ neutrals } \nu_\tau (\text{ex. } K_S^0 \rightarrow \pi^- \pi^+) (\text{"5-prong"})) / \Gamma_{\text{total}} \quad \Gamma_{116} / \Gamma$$

$$\Gamma_{116} / \Gamma = (\Gamma_{117} + \Gamma_{123}) / \Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.099±0.004 OUR FIT</b>		(Produced by HFLAV)		
<b>0.107±0.007 OUR AVERAGE</b>		Error includes scale factor of 1.1.		
0.170±0.022±0.026		<sup>1</sup> ACHARD	01D L3	1992–1995 LEP runs
0.097±0.005±0.011	419	GIBAUT	94B CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
0.102±0.029	13	BYLSMA	87 HRS	$E_{\text{cm}}^{ee} = 29 \text{ GeV}$
• • • We use the following data for averages but not for fits. • • •				
0.093±0.009±0.012		SCHAEEL	05C ALEP	1991–1995 LEP runs
0.115±0.013±0.006	112	<sup>2</sup> ABREU	01M DLPH	1992–1995 LEP runs
0.119±0.013±0.008	119	<sup>3</sup> ACKERSTAFF	99E OPAL	1991–1995 LEP runs
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.26 ±0.06 ±0.05		ACTON	92H OPAL	$E_{\text{cm}}^{ee} = 88.2\text{--}94.2 \text{ GeV}$
0.10 $\begin{smallmatrix} +0.05 \\ -0.04 \end{smallmatrix}$ ±0.03		DECAMP	92C ALEP	1989–1990 LEP runs
0.16 ±0.13 ±0.04		BEHREND	89B CELL	$E_{\text{cm}}^{ee} = 14\text{--}47 \text{ GeV}$
0.3 ±0.1 ±0.2		BARTEL	85F JADE	$E_{\text{cm}}^{ee} = 34.6 \text{ GeV}$
0.13 ±0.04	10	BELTRAMI	85 HRS	Repl. by BYLSMA 87
0.16 ±0.08 ±0.04	4	BURCHAT	85 MRK2	$E_{\text{cm}}^{ee} = 29 \text{ GeV}$
1.0 ±0.4	10	BEHREND	82 CELL	Repl. by BEHREND 89B

<sup>1</sup> The correlation coefficients between this measurement and the ACHARD 01D measurements of  $B(\tau \rightarrow \text{"1-prong"})$  and  $B(\tau \rightarrow \text{"3-prong"})$  are  $-0.082$  and  $-0.19$  respectively.

<sup>2</sup> The correlation coefficients between this measurement and the ABREU 01M measurements of  $B(\tau \rightarrow 1\text{-prong})$  and  $B(\tau \rightarrow 3\text{-prong})$  are  $-0.08$  and  $-0.08$  respectively.

<sup>3</sup> Not independent of ACKERSTAFF 99E  $B(\tau^- \rightarrow 3h^-2h^+ \nu_\tau (\text{ex. } K^0))$  and  $B(\tau^- \rightarrow 3h^-2h^+ \pi^0 \nu_\tau (\text{ex. } K^0))$  measurements.

$$\Gamma(3h^-2h^+ \nu_\tau (\text{ex. } K^0)) / \Gamma_{\text{total}} \quad \Gamma_{117} / \Gamma = (\Gamma_{118} + \Gamma_{120} + 0.0153\Gamma_{185}) / \Gamma$$

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>8.29±0.31 OUR FIT</b>		(Produced by HFLAV)		
<b>8.32±0.35 OUR AVERAGE</b>				
9.7 ±1.5 ±0.5	96	<sup>1</sup> ABDALLAH	06A DLPH	1992–1995 LEP runs
7.2 ±0.9 ±1.2	165	<sup>2</sup> SCHAEEL	05C ALEP	1991–1995 LEP runs
9.1 ±1.4 ±0.6	97	ACKERSTAFF	99E OPAL	1991–1995 LEP runs
7.7 ±0.5 ±0.9	295	GIBAUT	94B CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
6.4 ±2.3 ±1.0	12	ALBRECHT	88B ARG	$E_{\text{cm}}^{ee} = 10 \text{ GeV}$
5.1 ±2.0	7	BYLSMA	87 HRS	$E_{\text{cm}}^{ee} = 29 \text{ GeV}$
• • • We use the following data for averages but not for fits. • • •				
8.56±0.05±0.42	34k	AUBERT,B	05W BABR	$232 \text{ fb}^{-1}$ , $E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
8.0 ±1.1 ±1.3	58	BUSKULIC	96 ALEP	Repl. by SCHAEEL 05C
6.7 ±3.0	5	<sup>3</sup> BELTRAMI	85 HRS	Repl. by BYLSMA 87

<sup>1</sup> See footnote to ABDALLAH 06A  $\Gamma(\tau^- \rightarrow h^- \nu_\tau) / \Gamma_{\text{total}}$  measurement for correlations with other measurements.

<sup>2</sup> See footnote to SCHAEEL 05C  $\Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau) / \Gamma_{\text{total}}$  measurement for correlations with other measurements.

<sup>3</sup> The error quoted is statistical only.

$$\Gamma(3\pi^- 2\pi^+ \nu_\tau(\text{ex. } K^0, \omega))/\Gamma_{\text{total}} \qquad \Gamma_{118}/\Gamma = (\Gamma_{119} + \Gamma_{173})/\Gamma$$

VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT
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**8.27±0.31 OUR FIT** (Produced by HFLAV)

• • • We use the following data for averages but not for fits. • • •

**8.33±0.04±0.43** <sup>1</sup> LEES 12X BABR 468 fb<sup>-1</sup>  $E_{\text{cm}}^{\text{ee}} = 10.6$  GeV

<sup>1</sup> Not independent of LEES 12X  $\Gamma(\tau^- \rightarrow f_1(1285)\pi^- \nu_\tau \rightarrow 3\pi^- 2\pi^+ \nu_\tau)/\Gamma$  and  $\Gamma(\tau^- \rightarrow 3\pi^- 2\pi^+ \nu_\tau(\text{ex. } K^0, \omega, f_1(1285)))/\Gamma$  values.

$$\Gamma(3\pi^- 2\pi^+ \nu_\tau(\text{ex. } K^0, \omega, f_1(1285)))/\Gamma_{\text{total}} \qquad \Gamma_{119}/\Gamma$$

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
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**7.75±0.30 OUR FIT** (Produced by HFLAV)

**7.68±0.04±0.40** 69k LEES 12X BABR 468 fb<sup>-1</sup>  $E_{\text{cm}}^{\text{ee}} = 10.6$  GeV

$$\Gamma(K^- 2\pi^- 2\pi^+ \nu_\tau(\text{ex. } K^0))/\Gamma_{\text{total}} \qquad \Gamma_{120}/\Gamma$$

VALUE (units $10^{-6}$ )	DOCUMENT ID	TECN	COMMENT
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**0.6±1.2 OUR FIT** (Produced by HFLAV)

**0.6±0.5±1.1** <sup>1</sup> LEES 12X BABR 468 fb<sup>-1</sup>  $E_{\text{cm}}^{\text{ee}} = 10.6$  GeV

<sup>1</sup> LEES 12X measurement corresponds to the lower limit of  $< 2.4 \times 10^{-6}$  at 90% CL.

$$\Gamma(K^+ 3\pi^- \pi^+ \nu_\tau)/\Gamma_{\text{total}} \qquad \Gamma_{121}/\Gamma$$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
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**<5.0 × 10<sup>-6</sup>** 90 LEES 12X BABR 468 fb<sup>-1</sup>  $E_{\text{cm}}^{\text{ee}} = 10.6$  GeV

$$\Gamma(K^+ K^- 2\pi^- \pi^+ \nu_\tau)/\Gamma_{\text{total}} \qquad \Gamma_{122}/\Gamma$$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
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**<4.5 × 10<sup>-7</sup>** 90 LEES 12X BABR 468 fb<sup>-1</sup>  $E_{\text{cm}}^{\text{ee}} = 10.6$  GeV

$$\Gamma(3h^- 2h^+ \pi^0 \nu_\tau(\text{ex. } K^0))/\Gamma_{\text{total}} \qquad \Gamma_{123}/\Gamma = (\Gamma_{124} + \Gamma_{127})/\Gamma$$

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
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**1.65±0.11 OUR FIT** (Produced by HFLAV)

**1.74±0.27 OUR AVERAGE**

1.6 ±1.2 ±0.6 13 <sup>1</sup> ABDALLAH 06A DLPH 1992–1995 LEP runs

2.1 ±0.7 ±0.9 95 <sup>2</sup> SCHAEEL 05C ALEP 1991–1995 LEP runs

1.7 ±0.2 ±0.2 231 ANASTASSOV 01 CLEO  $E_{\text{cm}}^{\text{ee}} = 10.6$  GeV

2.7 ±1.8 ±0.9 23 ACKERSTAFF 99E OPAL 1991–1995 LEP runs

• • • We do not use the following data for averages, fits, limits, etc. • • •

1.8 ±0.7 ±1.2 18 BUSKULIC 96 ALEP Repl. by SCHAEEL 05C

1.9 ±0.4 ±0.4 31 GIBAUT 94B CLEO Repl. by ANASTASSOV 01

5.1 ±2.2 6 BYLSMA 87 HRS  $E_{\text{cm}}^{\text{ee}} = 29$  GeV

6.7 ±3.0 5 <sup>3</sup> BELTRAMI 85 HRS Repl. by BYLSMA 87

<sup>1</sup> See footnote to ABDALLAH 06A  $\Gamma(\tau^- \rightarrow h^- \nu_\tau)/\Gamma_{\text{total}}$  measurement for correlations with other measurements.

<sup>2</sup> SCHAEEL 05C quote  $(1.4 \pm 0.7 \pm 0.9) \times 10^{-4}$ . We add  $0.7 \times 10^{-4}$  to remove their correction for  $\tau^- \rightarrow \eta\pi^- \pi^+ \pi^- \nu_\tau \rightarrow 3\pi^- 2\pi^+ \pi^0 \nu_\tau$  and  $\tau^- \rightarrow K^*(892)^- \eta \nu_\tau \rightarrow 3\pi^- 2\pi^+ \pi^0 \nu_\tau$  decays. See footnote to SCHAEEL 05C  $\Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)/\Gamma_{\text{total}}$  measurement for correlations with other measurements.

<sup>3</sup> The error quoted is statistical only.

$$\Gamma(3\pi^- 2\pi^+ \pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}} \quad \Gamma_{124}/\Gamma$$

$$\Gamma_{124}/\Gamma = (\Gamma_{126} + 0.2302\Gamma_{160} + 0.892\Gamma_{185})/\Gamma$$

VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT
<b>1.64±0.11 OUR FIT</b>	(Produced by HFLAV)		

• • • We use the following data for averages but not for fits. • • •

$$\mathbf{1.65\pm 0.05\pm 0.09} \quad {}^1 \text{LEES} \quad 12\text{X} \quad \text{BABR} \quad 468 \text{ fb}^{-1} \quad E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$$

<sup>1</sup>Not independent of LEES 12X measurements of  $\Gamma(\tau^- \rightarrow 2\pi^- \pi^+ \omega \nu_\tau (\text{ex. } K^0))/\Gamma$ ,  $\Gamma(\tau^- \rightarrow \eta \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0))/\Gamma$ , and  $\Gamma(\tau^- \rightarrow 3\pi^- 2\pi^+ \pi^0 \nu_\tau (\text{ex. } K^0, \eta, \omega, f_1(1285)))/\Gamma$ .

$$\Gamma(3\pi^- 2\pi^+ \pi^0 \nu_\tau (\text{ex. } K^0, \eta, f_1(1285)))/\Gamma_{\text{total}} \quad \Gamma_{125}/\Gamma$$

VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT
<b>1.11±0.04±0.09</b>	<sup>1</sup> LEES	12X	BABR 468 fb <sup>-1</sup> E <sub>cm</sub> <sup>ee</sup> = 10.6 GeV

<sup>1</sup>Not independent of LEES 12X  $\Gamma(\tau^- \rightarrow 2\pi^- \pi^+ \omega \nu_\tau (\text{ex. } K^0))/\Gamma$  and  $\Gamma(\tau^- \rightarrow 3\pi^- 2\pi^+ \pi^0 \nu_\tau (\text{ex. } K^0, \eta, \omega, f_1(1285)))/\Gamma$  values.

$$\Gamma(3\pi^- 2\pi^+ \pi^0 \nu_\tau (\text{ex. } K^0, \eta, \omega, f_1(1285)))/\Gamma_{\text{total}} \quad \Gamma_{126}/\Gamma$$

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.38±0.09 OUR FIT</b>	(Produced by HFLAV)			
<b>0.36±0.03±0.09</b>	7.3k	LEES	12X	BABR 468 fb <sup>-1</sup> E <sub>cm</sub> <sup>ee</sup> = 10.6 GeV

$$\Gamma(K^- 2\pi^- 2\pi^+ \pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}} \quad \Gamma_{127}/\Gamma$$

VALUE (units $10^{-6}$ )	DOCUMENT ID	TECN	COMMENT
<b>1.1±0.6 OUR FIT</b>	(Produced by HFLAV)		
<b>1.1±0.4±0.4</b>	<sup>1</sup> LEES	12X	BABR 468 fb <sup>-1</sup> E <sub>cm</sub> <sup>ee</sup> = 10.6 GeV

<sup>1</sup>LEES 12X measurement corresponds to the lower limit of  $< 1.9 \times 10^{-6}$  at 90% CL.

$$\Gamma(K^+ 3\pi^- \pi^+ \pi^0 \nu_\tau)/\Gamma_{\text{total}} \quad \Gamma_{128}/\Gamma$$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;8 × 10<sup>-7</sup></b>	90	LEES	12X	BABR 468 fb <sup>-1</sup> E <sub>cm</sub> <sup>ee</sup> = 10.6 GeV

$$\Gamma(3h^- 2h^+ 2\pi^0 \nu_\tau)/\Gamma_{\text{total}} \quad \Gamma_{129}/\Gamma$$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;3.4 × 10<sup>-6</sup></b>	90	AUBERT,B	06	BABR 232 fb <sup>-1</sup> E <sub>cm</sub> <sup>ee</sup> = 10.6 GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

$$<1.1 \times 10^{-4} \quad 90 \quad \text{GIBAUT} \quad 94\text{B} \quad \text{CLEO} \quad E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$$

$$\Gamma((5\pi)^- \nu_\tau)/\Gamma_{\text{total}} \quad \Gamma_{130}/\Gamma$$

$$\Gamma_{130}/\Gamma = (\Gamma_{30} + \frac{1}{2}\Gamma_{45} + \Gamma_{48} + \frac{1}{2}\Gamma_{61} + \Gamma_{85} + \Gamma_{117} + 0.5559\Gamma_{150} + 0.892\Gamma_{180})/\Gamma$$

VALUE (%)	DOCUMENT ID	TECN	COMMENT
<b>0.78±0.05 OUR FIT</b>	(Produced by HFLAV)		

• • • We use the following data for averages but not for fits. • • •

**0.61 ± 0.06 ± 0.08** <sup>1</sup>GIBAUT 94B CLEO  $E_{\text{cm}}^{ee} = 10.6$  GeV

<sup>1</sup>Not independent of GIBAUT 94B  $B(3h^- 2h^+ \nu_\tau)$ , PROCARIO 93  $B(h^- 4\pi^0 \nu_\tau)$ , and BORTOLETTO 93  $B(2h^- h^+ 2\pi^0 \nu_\tau)/B(\text{"3prong"})$  measurements. Result is corrected for  $\eta$  contributions.

**$\Gamma(4h^- 3h^+ \geq 0 \text{ neutrals } \nu_\tau (\text{"7-prong"}))/\Gamma_{\text{total}}$**   **$\Gamma_{131}/\Gamma$**

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt; 3.0 \times 10^{-7}</math></b>	90	AUBERT,B	05F BABR	$232 \text{ fb}^{-1}$ , $E_{\text{cm}}^{ee} = 10.6$ GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

$< 1.8 \times 10^{-5}$	95	ACKERSTAFF	97J OPAL	1990–1995 LEP runs
$< 2.4 \times 10^{-6}$	90	EDWARDS	97B CLEO	$E_{\text{cm}}^{ee} = 10.6$ GeV
$< 2.9 \times 10^{-4}$	90	BYLSMA	87 HRS	$E_{\text{cm}}^{ee} = 29$ GeV

**$\Gamma(4h^- 3h^+ \nu_\tau)/\Gamma_{\text{total}}$**   **$\Gamma_{132}/\Gamma$**

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt; 4.3 \times 10^{-7}</math></b>	90	AUBERT,B	05F BABR	$232 \text{ fb}^{-1}$ , $E_{\text{cm}}^{ee} = 10.6$ GeV

**$\Gamma(4h^- 3h^+ \pi^0 \nu_\tau)/\Gamma_{\text{total}}$**   **$\Gamma_{133}/\Gamma$**

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt; 2.5 \times 10^{-7}</math></b>	90	AUBERT,B	05F BABR	$232 \text{ fb}^{-1}$ , $E_{\text{cm}}^{ee} = 10.6$ GeV

**$\Gamma(X^-(S=-1)\nu_\tau)/\Gamma_{\text{total}}$**   **$\Gamma_{134}/\Gamma$**

$$\Gamma_{134}/\Gamma = (\Gamma_{10} + \Gamma_{16} + \Gamma_{23} + \Gamma_{28} + \Gamma_{36} + \Gamma_{41} + \Gamma_{45} + \Gamma_{61} + \Gamma_{97} + \Gamma_{103} + \Gamma_{120} + \Gamma_{127} + \Gamma_{152} + \Gamma_{154} + \Gamma_{156} + 0.8312\Gamma_{170} + \Gamma_{179})/\Gamma$$

VALUE (%)	DOCUMENT ID	TECN	COMMENT
<b>2.92 ± 0.04 OUR FIT</b>	(Produced by HFLAV)		

• • • We use the following data for averages but not for fits. • • •

**2.87 ± 0.12** <sup>1</sup>BARATE 99R ALEP 1991–1995 LEP runs

<sup>1</sup>BARATE 99R perform a combined analysis of all ALEPH LEP 1 data on  $\tau$  branching fraction measurements for decay modes having total strangeness equal to  $-1$ .

**$\Gamma(K^*(892)^- \geq 0 \text{ neutrals } \geq 0 K_L^0 \nu_\tau)/\Gamma_{\text{total}}$**   **$\Gamma_{135}/\Gamma$**

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.42 ± 0.18 OUR AVERAGE</b>	Error includes scale factor of 1.4. See the ideogram below.			

1.19 ± 0.15 <sup>+0.13</sup>/<sub>-0.18</sub> 104 ALBRECHT 95H ARG  $E_{\text{cm}}^{ee} = 9.4\text{--}10.6$  GeV

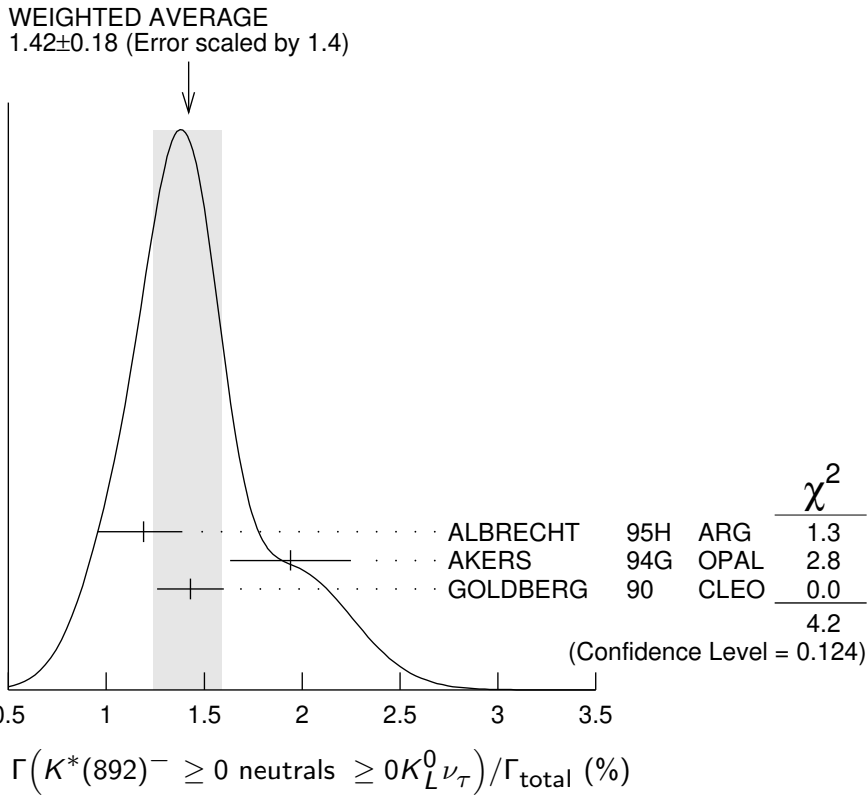
1.94 ± 0.27 ± 0.15 74 <sup>1</sup>AKERS 94G OPAL  $E_{\text{cm}}^{ee} = 88\text{--}94$  GeV

1.43 ± 0.11 ± 0.13 475 <sup>2</sup>GOLDBERG 90 CLEO  $E_{\text{cm}}^{ee} = 9.4\text{--}10.9$  GeV

<sup>1</sup>AKERS 94G reject events in which a  $K_S^0$  accompanies the  $K^*(892)^-$ . We do not correct for them.

<sup>2</sup>GOLDBERG 90 estimates that 10% of observed  $K^*(892)$  are accompanied by a  $\pi^0$ .





$\Gamma(K^*(892)^- \nu_\tau) / \Gamma_{\text{total}}$

$\Gamma_{136} / \Gamma$

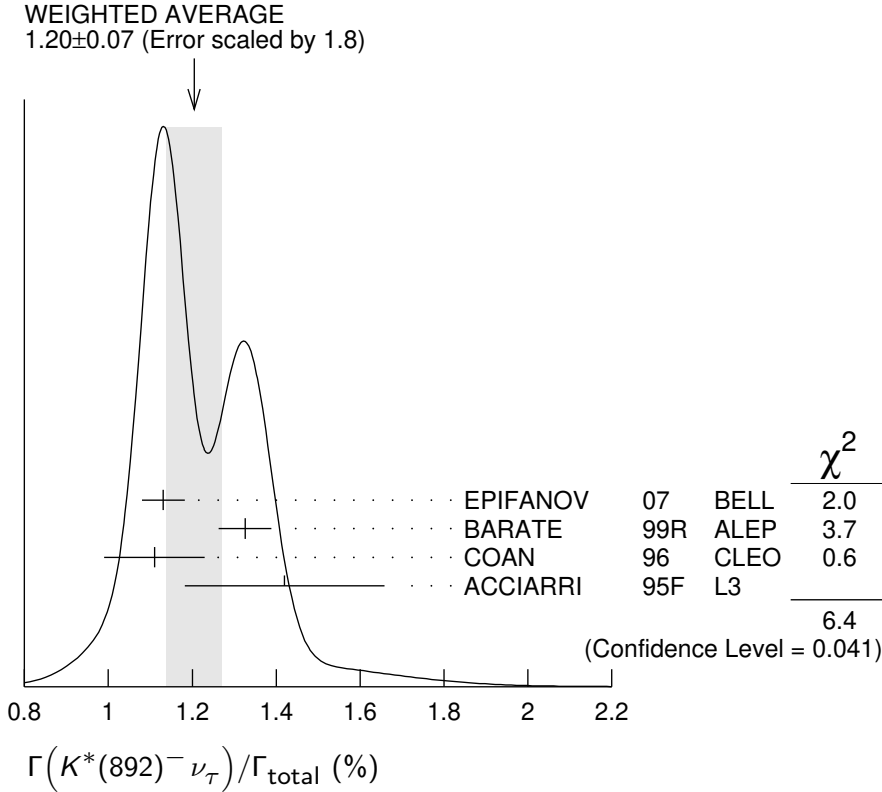
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>1.20 \pm 0.07</math></b>	<b>OUR AVERAGE</b>	Error includes scale factor of 1.8. See the ideogram below.		
$1.131 \pm 0.006 \pm 0.051$	49k	<sup>1</sup> EPIFANOV 07	BELL	$351 \text{ fb}^{-1}$ $E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$1.326 \pm 0.063$		BARATE 99R	ALEP	1991–1995 LEP runs
$1.11 \pm 0.12$		<sup>2</sup> COAN 96	CLEO	$E_{\text{cm}}^{ee} \approx 10.6 \text{ GeV}$
$1.42 \pm 0.22 \pm 0.09$		<sup>3</sup> ACCIARRI 95F	L3	1991–1993 LEP runs
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
$1.39 \pm 0.09 \pm 0.10$		<sup>4</sup> BUSKULIC 96	ALEP	Repl. by BARATE 99R
$1.45 \pm 0.13 \pm 0.11$	273	<sup>5</sup> BUSKULIC 94F	ALEP	Repl. by BUSKULIC 96
$1.23 \pm 0.21$	$+0.11$ $-0.21$	54 <sup>6</sup> ALBRECHT 88L	ARG	$E_{\text{cm}}^{ee} = 10 \text{ GeV}$
$1.9 \pm 0.3 \pm 0.4$	44	<sup>7</sup> TSCHIRHART 88	HRS	$E_{\text{cm}}^{ee} = 29 \text{ GeV}$
$1.5 \pm 0.4 \pm 0.4$	15	<sup>8</sup> AIHARA 87C	TPC	$E_{\text{cm}}^{ee} = 29 \text{ GeV}$
$1.3 \pm 0.3 \pm 0.3$	31	YELTON 86	MRK2	$E_{\text{cm}}^{ee} = 29 \text{ GeV}$
$1.7 \pm 0.7$	11	DORFAN 81	MRK2	$E_{\text{cm}}^{ee} = 4.2\text{--}6.7 \text{ GeV}$

<sup>1</sup> EPIFANOV 07 quote  $B(\tau^- \rightarrow K^*(892)^- \nu_\tau) B(K^*(892)^- \rightarrow K_S^0 \pi^-) = (3.77 \pm 0.02(\text{stat}) \pm 0.12(\text{syst}) \pm 0.12(\text{mod})) \times 10^{-3}$ . We add the systematic and model uncertainties in quadrature and divide by  $B(K^*(892)^- \rightarrow K_S^0 \pi^-) = 0.3333$ .

<sup>2</sup> Not independent of COAN 96  $B(\pi^- \bar{K}^0 \nu_\tau)$  and BATTLE 94  $B(K^- \pi^0 \nu_\tau)$  measurements.  $K\pi$  final states are consistent with and assumed to originate from  $K^*(892)^-$  production.

<sup>3</sup> This result is obtained from their  $B(\pi^- \bar{K}^0 \nu_\tau)$  assuming all those decays originate in  $K^*(892)^-$  decays.

- <sup>4</sup> Not independent of BUSKULIC 96  $B(\pi^- \bar{K}^0 \nu_\tau)$  and  $B(K^- \pi^0 \nu_\tau)$  measurements.  
<sup>5</sup> BUSKULIC 94F obtain this result from BUSKULIC 94F  $B(\bar{K}^0 \pi^- \nu_\tau)$  and BUSKULIC 94E  $B(K^- \pi^0 \nu_\tau)$  assuming all of those decays originate in  $K^*(892)^-$  decays.  
<sup>6</sup> The authors divide by  $\Gamma_2/\Gamma = 0.865$  to obtain this result.  
<sup>7</sup> Not independent of TSCHIRHART 88  $\Gamma(\tau^- \rightarrow h^- \bar{K}^0 \geq 0 \text{ neutrals} \geq 0 K_L^0 \nu_\tau) / \Gamma$ .  
<sup>8</sup> Decay  $\pi^-$  identified in this experiment, is assumed in the others.



$\Gamma(K^*(892)^- \nu_\tau) / \Gamma(\pi^- \pi^0 \nu_\tau)$   $\Gamma_{136} / \Gamma_{14}$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.075 \pm 0.027</math></b>	<sup>1</sup> ABREU	94K	DLPH LEP 1992 Z data

<sup>1</sup> ABREU 94K quote  $B(\tau^- \rightarrow K^*(892)^- \nu_\tau) B(K^*(892)^- \rightarrow K^- \pi^0) / B(\tau^- \rightarrow \rho^- \nu_\tau) = 0.025 \pm 0.009$ . We divide by  $B(K^*(892)^- \rightarrow K^- \pi^0) = 0.333$  to obtain this result.

$\Gamma(K^*(892)^- \nu_\tau \rightarrow \pi^- \bar{K}^0 \nu_\tau) / \Gamma(\pi^- \bar{K}^0 \nu_\tau)$   $\Gamma_{137} / \Gamma_{36}$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.933 \pm 0.027</math></b>	49k	EPIFANOV	07 BELL	$351 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

$\Gamma(K^*(892)^0 K^- \geq 0 \text{ neutrals} \nu_\tau) / \Gamma_{\text{total}}$   $\Gamma_{138} / \Gamma$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.32 \pm 0.08 \pm 0.12</math></b>	119	GOLDBERG	90 CLEO	$E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.9 \text{ GeV}$

$$\Gamma(K^*(892)^0 K^- \nu_\tau) / \Gamma_{\text{total}} \quad \Gamma_{139} / \Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.21 ± 0.04 OUR AVERAGE</b>				
0.213 ± 0.048		<sup>1</sup> BARATE 98	ALEP	1991–1995 LEP runs
0.20 ± 0.05 ± 0.04	47	ALBRECHT 95H	ARG	$E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.6$ GeV

<sup>1</sup> BARATE 98 measure the  $K^- (\rho^0 \rightarrow \pi^+ \pi^-)$  fraction in  $\tau^- \rightarrow K^- \pi^+ \pi^- \nu_\tau$  decays to be  $(35 \pm 11)\%$  and derive this result from their measurement of  $\Gamma(\tau^- \rightarrow K^- \pi^+ \pi^- \nu_\tau) / \Gamma_{\text{total}}$  assuming the intermediate states are all  $K^- \rho$  and  $K^- K^*(892)^0$ .

$$\Gamma(\bar{K}^*(892)^0 \pi^- \geq 0 \text{ neutrals } \nu_\tau) / \Gamma_{\text{total}} \quad \Gamma_{140} / \Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.38 ± 0.11 ± 0.13</b>	105	GOLDBERG 90	CLEO	$E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.9$ GeV

$$\Gamma(\bar{K}^*(892)^0 \pi^- \nu_\tau) / \Gamma_{\text{total}} \quad \Gamma_{141} / \Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.22 ± 0.05 OUR AVERAGE</b>				
0.209 ± 0.058		<sup>1</sup> BARATE 98	ALEP	1991–1995 LEP runs
0.25 ± 0.10 ± 0.05	27	ALBRECHT 95H	ARG	$E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.6$ GeV

<sup>1</sup> BARATE 98 measure the  $K^- K^*(892)^0$  fraction in  $\tau^- \rightarrow K^- K^+ \pi^- \nu_\tau$  decays to be  $(87 \pm 13)\%$  and derive this result from their measurement of  $\Gamma(\tau^- \rightarrow K^- K^+ \pi^- \nu_\tau) / \Gamma_{\text{total}}$ .

$$\Gamma((\bar{K}^*(892)\pi)^- \nu_\tau \rightarrow \pi^- \bar{K}^0 \pi^0 \nu_\tau) / \Gamma_{\text{total}} \quad \Gamma_{142} / \Gamma$$

VALUE (%)	DOCUMENT ID	TECN	COMMENT
<b>0.10 ± 0.04 OUR AVERAGE</b>			
0.097 ± 0.044 ± 0.036	<sup>1</sup> BARATE 99K	ALEP	1991–1995 LEP runs
0.106 ± 0.037 ± 0.032	<sup>2</sup> BARATE 98E	ALEP	1991–1995 LEP runs

<sup>1</sup> BARATE 99K measure  $K^0$ 's by detecting  $K_L^0$ 's in their hadron calorimeter. They determine the  $\bar{K}^0 \rho^-$  fraction in  $\tau^- \rightarrow \pi^- \bar{K}^0 \pi^0 \nu_\tau$  decays to be  $(0.72 \pm 0.12 \pm 0.10)$  and multiply their  $B(\pi^- \bar{K}^0 \pi^0 \nu_\tau)$  measurement by one minus this fraction to obtain the quoted result.

<sup>2</sup> BARATE 98E reconstruct  $K^0$ 's using  $K_S^0 \rightarrow \pi^+ \pi^-$  decays. They determine the  $\bar{K}^0 \rho^-$  fraction in  $\tau^- \rightarrow \pi^- \bar{K}^0 \pi^0 \nu_\tau$  decays to be  $(0.64 \pm 0.09 \pm 0.10)$  and multiply their  $B(\pi^- \bar{K}^0 \pi^0 \nu_\tau)$  measurement by one minus this fraction to obtain the quoted result.

$$\Gamma(K_1(1270)^- \nu_\tau) / \Gamma_{\text{total}} \quad \Gamma_{143} / \Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.47 ± 0.11 OUR AVERAGE</b>				
0.48 ± 0.11		BARATE 99R	ALEP	1991–1995 LEP runs
0.41 <sup>+0.41</sup> <sub>-0.35</sub> ± 0.10	5	<sup>1</sup> BAUER 94	TPC	$E_{\text{cm}}^{\text{ee}} = 29$ GeV

<sup>1</sup> We multiply 0.41% by 0.25, the relative systematic error quoted by BAUER 94, to obtain the systematic error.

$$\Gamma(K_1(1400)^- \nu_\tau) / \Gamma_{\text{total}} \qquad \Gamma_{144} / \Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.17 ± 0.26 OUR AVERAGE</b>		Error includes scale factor of 1.7.		
0.05 ± 0.17		BARATE	99R ALEP	1991–1995 LEP runs
0.76 <sup>+0.40</sup> <sub>-0.33</sub> ± 0.20	11	<sup>1</sup> BAUER	94 TPC	$E_{\text{cm}}^{ee} = 29$ GeV

<sup>1</sup> We multiply 0.76% by 0.25, the relative systematic error quoted by BAUER 94, to obtain the systematic error.

$$[\Gamma(K_1(1270)^- \nu_\tau) + \Gamma(K_1(1400)^- \nu_\tau)] / \Gamma_{\text{total}} \qquad (\Gamma_{143} + \Gamma_{144}) / \Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.17<sup>+0.41</sup><sub>-0.37</sub> ± 0.29</b>	16	<sup>1</sup> BAUER	94 TPC	$E_{\text{cm}}^{ee} = 29$ GeV

<sup>1</sup> We multiply 1.17% by 0.25, the relative systematic error quoted by BAUER 94, to obtain the systematic error. Not independent of BAUER 94  $B(K_1(1270)^- \nu_\tau)$  and BAUER 94  $B(K_1(1400)^- \nu_\tau)$  measurements.

$$\Gamma(K_1(1270)^- \nu_\tau) / [\Gamma(K_1(1270)^- \nu_\tau) + \Gamma(K_1(1400)^- \nu_\tau)] \qquad \Gamma_{143} / (\Gamma_{143} + \Gamma_{144})$$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.69 ± 0.15 OUR AVERAGE</b>			
0.71 ± 0.16 ± 0.11	<sup>1</sup> ABBIENDI 00D	OPAL	1990–1995 LEP runs
0.66 ± 0.19 ± 0.13	<sup>2</sup> ASNER 00B	CLEO	$E_{\text{cm}}^{ee} = 10.6$ GeV

<sup>1</sup> ABBIENDI 00D assume the resonance structure of  $\tau^- \rightarrow K^- \pi^+ \pi^- \nu_\tau$  decays is dominated by the  $K_1(1270)^-$  and  $K_1(1400)^-$  resonances.

<sup>2</sup> ASNER 00B assume the resonance structure of  $\tau^- \rightarrow K^- \pi^+ \pi^- \nu_\tau$  (ex.  $K^0$ ) decays is dominated by  $K_1(1270)^-$  and  $K_1(1400)^-$  resonances.

$$\Gamma(K^*(1410)^- \nu_\tau) / \Gamma_{\text{total}} \qquad \Gamma_{145} / \Gamma$$

VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT
<b>1.5<sup>+1.4</sup><sub>-1.0</sub></b>	BARATE	99R ALEP	1991–1995 LEP runs

$$\Gamma(K_0^*(1430)^- \nu_\tau) / \Gamma_{\text{total}} \qquad \Gamma_{146} / \Gamma$$

VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.5</b>	95	BARATE	99R ALEP	1991–1995 LEP runs

$$\Gamma(K_2^*(1430)^- \nu_\tau) / \Gamma_{\text{total}} \qquad \Gamma_{147} / \Gamma$$

VALUE (%)	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.3</b>	95		TSCHIRHART 88	HRS	$E_{\text{cm}}^{ee} = 29$ GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

<0.33	95	<sup>1</sup> ACCIARRI	95F L3	1991–1993 LEP runs
<0.9	95	0	DORFAN 81	MRK2 $E_{\text{cm}}^{ee} = 4.2$ –6.7 GeV

<sup>1</sup> ACCIARRI 95F quote  $B(\tau^- \rightarrow K^*(1430)^- \rightarrow \pi^- \bar{K}^0 \nu_\tau) < 0.11\%$ . We divide by  $B(K^*(1430)^- \rightarrow \pi^- \bar{K}^0) = 0.33$  to obtain the limit shown.

$$\Gamma(a_0(980)^- \geq 0 \text{ neutrals } \nu_\tau) / \Gamma_{\text{total}} \times B(a_0(980) \rightarrow K^0 K^-) \qquad \Gamma_{148} / \Gamma \times B$$

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;2.8</b>	90	GOLDBERG 90	CLEO	$E_{\text{cm}}^{ee} = 9.4$ –10.9 GeV

$\Gamma(\eta\pi^-\nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{149}/\Gamma$ 

VALUE (units $10^{-4}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>&lt; 0.99</b>	95		<sup>1</sup> DEL-AMO-SA...11E	BABR	$470 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
< 6.2	95		BUSKULIC	97C ALEP	1991–1994 LEP runs
< 1.4	95	0	BARTELT	96 CLEO	$E_{\text{cm}}^{ee} \approx 10.6 \text{ GeV}$
< 3.4	95		ARTUSO	92 CLEO	$E_{\text{cm}}^{ee} \approx 10.6 \text{ GeV}$
< 90	95		ALBRECHT	88M ARG	$E_{\text{cm}}^{ee} \approx 10 \text{ GeV}$
<140	90		BEHREND	88 CELL	$E_{\text{cm}}^{ee} = 14\text{--}46.8 \text{ GeV}$
<180	95		BARINGER	87 CLEO	$E_{\text{cm}}^{ee} = 10.5 \text{ GeV}$
<250	90	0	COFFMAN	87 MRK3	$E_{\text{cm}}^{ee} = 3.77 \text{ GeV}$
510 $\pm 100 \pm 120$		65	DERRICK	87 HRS	$E_{\text{cm}}^{ee} = 29 \text{ GeV}$
<100	95		GAN	87B MRK2	$E_{\text{cm}}^{ee} = 29 \text{ GeV}$

<sup>1</sup> DEL-AMO-SANCHEZ 11E also quote  $B(\tau^- \rightarrow \eta\pi^-\nu_\tau) = (3.4 \pm 3.4 \pm 2.1) \times 10^{-5}$ .

 $\Gamma(\eta\pi^-\pi^0\nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{150}/\Gamma$ 

VALUE (units $10^{-3}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.39 <math>\pm</math> 0.07 OUR FIT</b>	(Produced by HFLAV)				
<b>1.38 <math>\pm</math> 0.09 OUR AVERAGE</b>	Error includes scale factor of 1.2.				
1.35 $\pm$ 0.03 $\pm$ 0.07		6.0k	INAMI	09 BELL	$490 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
1.8 $\pm$ 0.4 $\pm$ 0.2			BUSKULIC	97C ALEP	1991–1994 LEP runs
1.7 $\pm$ 0.2 $\pm$ 0.2		125	ARTUSO	92 CLEO	$E_{\text{cm}}^{ee} \approx 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

< 11.0	95		ALBRECHT	88M ARG	$E_{\text{cm}}^{ee} \approx 10 \text{ GeV}$
< 21.0	95		BARINGER	87 CLEO	$E_{\text{cm}}^{ee} = 10.5 \text{ GeV}$
42.0 $\pm$ 7.0 $\pm$ 16.0			<sup>1</sup> GAN	87 MRK2	$E_{\text{cm}}^{ee} = 29 \text{ GeV}$

<sup>1</sup> Highly correlated with GAN 87  $\Gamma(\pi^- 3\pi^0\nu_\tau)/\Gamma(\text{total})$  value.

 $\Gamma(\eta\pi^-\pi^0\pi^0\nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{151}/\Gamma$ 

VALUE (units $10^{-4}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.9 <math>\pm</math> 0.4 OUR FIT</b>	(Produced by HFLAV)				
<b>1.81 <math>\pm</math> 0.31 OUR AVERAGE</b>					
2.01 $\pm$ 0.34 $\pm$ 0.22		381	LEES	12X BABR	$468 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

• • • We use the following data for averages but not for fits. • • •

1.5  $\pm$  0.5      30      <sup>1</sup> ANASTASSOV 01      CLEO       $E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

1.4 $\pm$ 0.6 $\pm$ 0.3		15	<sup>2</sup> BERGFELD	97 CLEO	Repl. by ANASTASSOV 01
< 4.3	95		ARTUSO	92 CLEO	$E_{\text{cm}}^{ee} \approx 10.6 \text{ GeV}$
<120	95		ALBRECHT	88M ARG	$E_{\text{cm}}^{ee} \approx 10 \text{ GeV}$

<sup>1</sup> Weighted average of BERGFELD 97 and ANASTASSOV 01 value of  $(1.5 \pm 0.6 \pm 0.3) \times 10^{-4}$  obtained using  $\eta$ 's reconstructed from  $\eta \rightarrow \pi^+\pi^-\pi^0$  decays.

<sup>2</sup> BERGFELD 97 reconstruct  $\eta$ 's using  $\eta \rightarrow \gamma\gamma$  decays.

$\Gamma(\eta K^- \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{152}/\Gamma$ 

VALUE (units $10^{-4}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.55±0.08 OUR FIT</b>					(Produced by HFLAV)
<b>1.54±0.08 OUR AVERAGE</b>					
1.42±0.11±0.07		690	DEL-AMO-SA..11E	BABR	470 fb <sup>-1</sup> $E_{\text{cm}}^{ee} = 10.6$ GeV
1.58±0.05±0.09		1.6k	INAMI 09	BELL	490 fb <sup>-1</sup> $E_{\text{cm}}^{ee} = 10.6$ GeV
2.9 <sup>+1.3</sup> / <sub>-1.2</sub> ±0.7			BUSKULIC 97C	ALEP	1991–1994 LEP runs
2.6 ±0.5 ±0.5		85	BARTELT 96	CLEO	$E_{\text{cm}}^{ee} \approx 10.6$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •					
< 4.7		95	ARTUSO 92	CLEO	$E_{\text{cm}}^{ee} \approx 10.6$ GeV

 $\Gamma(\eta K^*(892)^- \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{153}/\Gamma$ 

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.38±0.15 OUR AVERAGE</b>				
1.34±0.12±0.09	245	<sup>1</sup> INAMI 09	BELL	490 fb <sup>-1</sup> $E_{\text{cm}}^{ee} = 10.6$ GeV
2.90±0.80±0.42	25	BISHAI 99	CLEO	$E_{\text{cm}}^{ee} = 10.6$ GeV

<sup>1</sup> Not independent of INAMI 09  $B(\tau^- \rightarrow \eta K^- \pi^0 \nu_\tau)$  and  $B(\tau^- \rightarrow \eta \bar{K}^0 \pi^- \nu_\tau)$  values.

 $\Gamma(\eta K^- \pi^0 \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{154}/\Gamma$ 

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.48±0.12 OUR FIT</b>				(Produced by HFLAV)
<b>0.48±0.12 OUR AVERAGE</b>				
0.46±0.11±0.04	270	INAMI 09	BELL	490 fb <sup>-1</sup> $E_{\text{cm}}^{ee} = 10.6$ GeV
1.77±0.56±0.71	36	BISHAI 99	CLEO	$E_{\text{cm}}^{ee} = 10.6$ GeV

 $\Gamma(\eta K^- \pi^0 (\text{non-}K^*(892)) \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{155}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;3.5 × 10<sup>-5</sup></b>	90	INAMI 09	BELL	490 fb <sup>-1</sup> $E_{\text{cm}}^{ee} = 10.6$ GeV

 $\Gamma(\eta \bar{K}^0 \pi^- \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{156}/\Gamma$ 

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.94±0.15 OUR FIT</b>				(Produced by HFLAV)
<b>0.93±0.15 OUR AVERAGE</b>				
0.88±0.14±0.06	161	<sup>1</sup> INAMI 09	BELL	490 fb <sup>-1</sup> $E_{\text{cm}}^{ee} = 10.6$ GeV
2.20±0.70±0.22	15	<sup>2</sup> BISHAI 99	CLEO	$E_{\text{cm}}^{ee} = 10.6$ GeV

<sup>1</sup> We multiply the INAMI 09 measurement  $B(\tau^- \rightarrow \eta K_S^0 \pi^- \nu_\tau) = (0.44 \pm 0.07 \pm 0.03) \times 10^{-4}$  by 2 to obtain the listed value.

<sup>2</sup> We multiply the BISHAI 99 measurement  $B(\tau^- \rightarrow \eta K_S^0 \pi^- \nu_\tau) = (1.10 \pm 0.35 \pm 0.11) \times 10^{-4}$  by 2 to obtain the listed value.

 $\Gamma(\eta \bar{K}^0 \pi^- \pi^0 \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{157}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;5.0 × 10<sup>-5</sup></b>	90	<sup>1</sup> INAMI 09	BELL	490 fb <sup>-1</sup> $E_{\text{cm}}^{ee} = 10.6$ GeV

<sup>1</sup> We multiply the INAMI 09 measurement  $B(\tau^- \rightarrow \eta K_S^0 \pi^- \pi^0 \nu_\tau) < 2.5 \times 10^{-5}$  by 2 to obtain the listed value.

$\Gamma(\eta K^- K^0 \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{158}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<9.0 \times 10^{-6}$	90	<sup>1</sup> INAMI 09	BELL	490 fb <sup>-1</sup> $E_{\text{cm}}^{ee} = 10.6$ GeV

<sup>1</sup>We multiply the INAMI 09 measurement  $B(\tau^- \rightarrow \eta K^- K_S^0 \nu_\tau) < 4.5 \times 10^{-6}$  by 2 to obtain the listed value.

 $\Gamma(\eta \pi^+ \pi^- \pi^- \geq 0 \text{ neutrals } \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{159}/\Gamma$ 

VALUE (%)	CL%	DOCUMENT ID	TECN	COMMENT
$<0.3$	90	ABACHI 87B	HRS	$E_{\text{cm}}^{ee} = 29$ GeV

 $\Gamma(\eta \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$   $\Gamma_{160}/\Gamma$ 

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2.20 ± 0.13 OUR FIT</b>		(Produced by HFLAV)		
<b>2.23 ± 0.12 OUR AVERAGE</b>				

2.10 ± 0.09 ± 0.13	2.9k	<sup>1</sup> LEES	12X	BABR $\eta \rightarrow \gamma\gamma$
2.37 ± 0.12 ± 0.18	1.4k	<sup>1</sup> LEES	12X	BABR $\eta \rightarrow \pi^+ \pi^- \pi^0$
2.54 ± 0.27 ± 0.25	315	<sup>1</sup> LEES	12X	BABR $\eta \rightarrow 3\pi^0$

• • • We use the following data for averages but not for fits. • • •

2.3 ± 0.5	170	<sup>2</sup> ANASTASSOV 01	CLEO	$E_{\text{cm}}^{ee} = 10.6$ GeV
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• • • We do not use the following data for averages, fits, limits, etc. • • •

1.60 ± 0.05 ± 0.11	1.8 k	AUBERT	08AE	BABR Repl. by LEES 12X
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3.4 $^{+0.6}_{-0.5}$ ± 0.6	89	<sup>3</sup> BERGFELD	97	CLEO Repl. by ANASTASSOV 01
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<sup>1</sup>LEES 12X uses 468 fb<sup>-1</sup> of data taken at  $E_{\text{cm}}^{ee} = 10.6$  GeV. It gives the average of the three measurements listed here as  $(2.25 \pm 0.07 \pm 0.12) \times 10^{-4}$ .

<sup>2</sup>Weighted average of BERGFELD 97 and ANASTASSOV 01 measurements using  $\eta$ 's reconstructed from  $\eta \rightarrow \pi^+ \pi^- \pi^0$  and  $\eta \rightarrow 3\pi^0$  decays.

<sup>3</sup>BERGFELD 97 reconstruct  $\eta$ 's using  $\eta \rightarrow \gamma\gamma$  and  $\eta \rightarrow 3\pi^0$  decays.

 $\Gamma(\eta \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0, f_1(1285)))/\Gamma_{\text{total}}$   $\Gamma_{161}/\Gamma$ 

VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT
<b>0.99 ± 0.09 ± 0.13</b>	<sup>1</sup> LEES	12X	BABR 468 fb <sup>-1</sup> $E_{\text{cm}}^{ee} = 10.6$ GeV

<sup>1</sup>LEES 12X obtain this result by subtracting their  $B(\tau^- \rightarrow f_1(1285) \pi^- \nu_\tau \rightarrow \eta \pi^- \pi^+ \pi^- \nu_\tau)$  measurement from their  $B(\tau^- \rightarrow \eta \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0))$  measurement.

 $\Gamma(\eta a_1(1260)^- \nu_\tau \rightarrow \eta \pi^- \rho^0 \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{162}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<3.9 \times 10^{-4}$	90	BERGFELD 97	CLEO	$E_{\text{cm}}^{ee} = 10.6$ GeV

 $\Gamma(\eta \eta \pi^- \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{163}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<7.4 \times 10^{-6}$	90	INAMI 09	BELL	490 fb <sup>-1</sup> $E_{\text{cm}}^{ee} = 10.6$ GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

$<1.1 \times 10^{-4}$	95	ARTUSO 92	CLEO	$E_{\text{cm}}^{ee} \approx 10.6$ GeV
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$<8.3 \times 10^{-3}$	95	ALBRECHT 88M	ARG	$E_{\text{cm}}^{ee} \approx 10$ GeV
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$\Gamma(\eta\eta\pi^-\pi^0\nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{164}/\Gamma$ 

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt; 2.0</b>	95	ARTUSO	92	CLEO $E_{\text{cm}}^{ee} \approx 10.6$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<90	95	ALBRECHT	88M	ARG $E_{\text{cm}}^{ee} \approx 10$ GeV

 $\Gamma(\eta\eta K^-\nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{165}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt; <math>3.0 \times 10^{-6}</math></b>	90	INAMI	09	BELL $490 \text{ fb}^{-1}$ $E_{\text{cm}}^{ee} = 10.6$ GeV

 $\Gamma(\eta'(958)\pi^-\nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{166}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt; <math>4.0 \times 10^{-6}</math></b>	90	LEES	12X	BABR $468 \text{ fb}^{-1}$ $E_{\text{cm}}^{ee} = 10.6$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
< $7.2 \times 10^{-6}$	90	AUBERT	08AE	BABR $384 \text{ fb}^{-1}$ , $E_{\text{cm}}^{ee} = 10.6$ GeV
< $7.4 \times 10^{-5}$	90	BERGFELD	97	CLEO $E_{\text{cm}}^{ee} = 10.6$ GeV

 $\Gamma(\eta'(958)\pi^-\pi^0\nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{167}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt; <math>1.2 \times 10^{-5}</math></b>	90	LEES	12X	BABR $468 \text{ fb}^{-1}$ $E_{\text{cm}}^{ee} = 10.6$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
< $8.0 \times 10^{-5}$	90	BERGFELD	97	CLEO $E_{\text{cm}}^{ee} = 10.6$ GeV

 $\Gamma(\eta'(958)K^-\nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{168}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt; <math>2.4 \times 10^{-6}</math></b>	90	LEES	12X	BABR $468 \text{ fb}^{-1}$ $E_{\text{cm}}^{ee} = 10.6$ GeV

 $\Gamma(\phi\pi^-\nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{169}/\Gamma$ 

VALUE (units $10^{-5}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>3.42 \pm 0.55 \pm 0.25</math></b>		344	AUBERT	08	BABR $342 \text{ fb}^{-1}$ $E_{\text{cm}}^{ee} = 10.6$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •					
< 20	90		<sup>1</sup> AVERY	97	CLEO $E_{\text{cm}}^{ee} = 10.6$ GeV
< 35	90		ALBRECHT	95H	ARG $E_{\text{cm}}^{ee} = 9.4\text{--}10.6$ GeV

<sup>1</sup>AVERY 97 limit varies from  $(1.2\text{--}2.0) \times 10^{-4}$  depending on decay model assumptions.

 $\Gamma(\phi K^-\nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{170}/\Gamma$ 

VALUE (units $10^{-5}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>4.4 \pm 1.6</math></b>					<b>OUR FIT</b> (Produced by HFLAV)

**$3.70 \pm 0.33$  OUR AVERAGE** Error includes scale factor of 1.3.

• • • We use the following data for averages but not for fits. • • •

$3.39 \pm 0.20 \pm 0.28$	274	AUBERT	08	BABR	$342 \text{ fb}^{-1}$ $E_{\text{cm}}^{ee} = 10.6$ GeV
$4.05 \pm 0.25 \pm 0.26$	551	INAMI	06	BELL	$401 \text{ fb}^{-1}$ $E_{\text{cm}}^{ee} = 10.6$ GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

< 6.7	90	<sup>1</sup> AVERY	97	CLEO	$E_{\text{cm}}^{ee} = 10.6$ GeV
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<sup>1</sup>AVERY 97 limit varies from  $(5.4\text{--}6.7) \times 10^{-5}$  depending on decay model assumptions.



$$\Gamma(f_1(1285)\pi^-\nu_\tau)/\Gamma_{\text{total}} \qquad \Gamma_{171}/\Gamma$$

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>3.9 ± 0.5 OUR AVERAGE</b>		Error includes scale factor of 1.9.		
4.73 ± 0.28 ± 0.45	3.7k	<sup>1</sup> LEES	12X BABR	468 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
3.60 ± 0.18 ± 0.23	2.5k	<sup>2</sup> LEES	12X BABR	468 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
3.19 ± 0.18 ± 1.00	1.3 k	<sup>3</sup> AUBERT	08AE BABR	Repl. by LEES 12X
3.9 ± 0.7 ± 0.5	1.4 k	<sup>4</sup> AUBERT,B	05W BABR	Repl. by LEES 12X
5.8 $\begin{smallmatrix} +1.4 \\ -1.3 \end{smallmatrix}$ ± 1.8	54	<sup>5</sup> BERGFELD	97 CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6$ GeV

<sup>1</sup> LEES 12X obtain this value by dividing their  $B(\tau^- \rightarrow f_1(1285)\pi^-\nu_\tau \rightarrow 3\pi^- 2\pi^+\nu_\tau)$  measurement by the PDG 12 value of  $B(f_1(1285) \rightarrow 2\pi^+ 2\pi^-) = 0.111^{+0.007}_{-0.006}$ .

<sup>2</sup> LEES 12X obtain this value by dividing their  $B(\tau^- \rightarrow f_1(1285)\pi^-\nu_\tau \rightarrow \eta\pi^-\pi^+\pi^-\nu_\tau)$  measurement by 2/3 of the PDG 12 value of  $B(f_1(1285) \rightarrow \eta\pi\pi) = 0.524^{+0.019}_{-0.021}$ .

<sup>3</sup> AUBERT 08AE obtain this value by dividing their  $B(\tau^- \rightarrow f_1(1285)\pi^-\nu_\tau \rightarrow \eta\pi^-\pi^+\pi^-\nu_\tau)$  measurement by the PDG 06 value of  $B(f_1(1285) \rightarrow \eta\pi^-\pi^+) = 0.35 \pm 0.11$ . The quote  $(3.19 \pm 0.18 \pm 0.16 \pm 0.99) \times 10^{-4}$  where the final error is due to the uncertainty on  $B(f_1(1285) \rightarrow \eta\pi^-\pi^+)$ . We combine the two systematic errors in quadrature.

<sup>4</sup> AUBERT,B 05W use the  $f_1(1285) \rightarrow 2\pi^+ 2\pi^-$  decay mode and the PDG 04 value of  $B(f_1(1285) \rightarrow 2\pi^+ 2\pi^-) = 0.110^{+0.007}_{-0.006}$ .

<sup>5</sup> BERGFELD 97 use the  $f_1(1285) \rightarrow \eta\pi^+\pi^-$  decay mode.

$$\Gamma(f_1(1285)\pi^-\nu_\tau \rightarrow \eta\pi^-\pi^+\pi^-\nu_\tau)/\Gamma_{\text{total}} \qquad \Gamma_{172}/\Gamma$$

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.18 ± 0.07 OUR AVERAGE</b>		Error includes scale factor of 1.3.		
1.26 ± 0.06 ± 0.06	2.5k	LEES	12X BABR	468 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
1.11 ± 0.06 ± 0.05	1.3 k	AUBERT	08AE BABR	384 fb <sup>-1</sup> , $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV

$$\Gamma(f_1(1285)\pi^-\nu_\tau \rightarrow \eta\pi^-\pi^+\pi^-\nu_\tau)/\Gamma(\eta\pi^-\pi^+\pi^-\nu_\tau(\text{ex.}K^0)) \qquad \Gamma_{172}/\Gamma_{160}$$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.69 ± 0.01 ± 0.05</b>	<sup>1</sup> AUBERT	08AE BABR	384 fb <sup>-1</sup> , $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
0.55 ± 0.14	BERGFELD	97 CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6$ GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

<sup>1</sup> Not independent of AUBERT 08AE  $B(\tau^- \rightarrow f_1(1285)\pi^-\nu_\tau \rightarrow \eta\pi^-\pi^+\pi^-\nu_\tau)$  and  $B(\tau^- \rightarrow \eta\pi^-\pi^+\pi^-\nu_\tau(\text{ex.}K^0))$  values.

$$\Gamma(f_1(1285)\pi^-\nu_\tau \rightarrow 3\pi^- 2\pi^+\nu_\tau)/\Gamma_{\text{total}} \qquad \Gamma_{173}/\Gamma$$

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.52 ± 0.04 OUR FIT</b>		(Produced by HFLAV)		
<b>0.520 ± 0.031 ± 0.037</b>	3.7k	LEES	12X BABR	468 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV

$$\Gamma(\pi(1300)^-\nu_\tau \rightarrow (\rho\pi)^-\nu_\tau \rightarrow (3\pi)^-\nu_\tau)/\Gamma_{\text{total}} \qquad \Gamma_{174}/\Gamma$$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt; 1.0 × 10<sup>-4</sup></b>	90	ASNER	00 CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6$ GeV

$\Gamma(\pi(1300)^- \nu_\tau \rightarrow ((\pi\pi)_{S\text{-wave}} \pi)^- \nu_\tau \rightarrow (3\pi)^- \nu_\tau) / \Gamma_{\text{total}}$					$\Gamma_{175} / \Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$<1.9 \times 10^{-4}$	90	ASNER	00	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

$\Gamma(h^- \omega \geq 0 \text{ neutrals } \nu_\tau) / \Gamma_{\text{total}}$					$\Gamma_{176} / \Gamma$
$\Gamma_{176} / \Gamma = (\Gamma_{178} + \Gamma_{179} + \Gamma_{180}) / \Gamma$					
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>2.40 ± 0.08 OUR FIT</b>	(Produced by HFLAV)				

• • • We use the following data for averages but not for fits. • • •

<b>1.65 ± 0.3 ± 0.2</b>	1513	ALBRECHT	88M	ARG	$E_{\text{cm}}^{\text{ee}} \approx 10 \text{ GeV}$
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$\Gamma(h^- \omega \nu_\tau) / \Gamma_{\text{total}}$					$\Gamma_{177} / \Gamma = (\Gamma_{178} + \Gamma_{179}) / \Gamma$
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>1.99 ± 0.06 OUR FIT</b>	(Produced by HFLAV)				

**1.92 ± 0.07 OUR AVERAGE**

1.91 ± 0.07 ± 0.06	5803	BUSKULIC	97C	ALEP	1991–1994 LEP runs
1.60 ± 0.27 ± 0.41	139	BARINGER	87	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.5 \text{ GeV}$

• • • We use the following data for averages but not for fits. • • •

1.95 ± 0.07 ± 0.11	2223	<sup>1</sup> BALEST	95C	CLEO	$E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$
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<sup>1</sup> Not independent of BALEST 95C  $B(\tau^- \rightarrow h^- \omega \nu_\tau) / B(\tau^- \rightarrow h^- h^- h^+ \pi^0 \nu_\tau)$  value.

$[\Gamma(\pi^- \omega \nu_\tau) + \Gamma(K^- \omega \nu_\tau)] / \Gamma(h^- h^- h^+ \pi^0 \nu_\tau \text{ (ex. } K^0))$					$(\Gamma_{178} + \Gamma_{179}) / \Gamma_{74}$
$(\Gamma_{178} + \Gamma_{179}) / \Gamma_{74} = (\Gamma_{178} + \Gamma_{179}) / (\Gamma_{78} + \Gamma_{103} + \Gamma_{107} + 0.2302\Gamma_{152} + 0.892\Gamma_{178} + 0.892\Gamma_{179} + 0.0153\Gamma_{180})$					
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>43.5 ± 1.4 OUR FIT</b>	(Produced by HFLAV)				
<b>45.3 ± 1.9 OUR AVERAGE</b>					

43.1 ± 3.3	2350	<sup>1</sup> BUSKULIC	96	ALEP	LEP 1991–1993 data
46.4 ± 1.6 ± 1.7	2223	<sup>2</sup> BALEST	95C	CLEO	$E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

37 ± 5 ± 2	458	<sup>3</sup> ALBRECHT	91D	ARG	$E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.6 \text{ GeV}$
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<sup>1</sup> BUSKULIC 96 quote the fraction of  $\tau \rightarrow h^- h^- h^+ \pi^0 \nu_\tau$  (ex.  $K^0$ ) decays which originate in a  $h^- \omega$  final state =  $0.383 \pm 0.029$ . We divide this by the  $\omega(782) \rightarrow \pi^+ \pi^- \pi^0$  branching fraction (0.888).

<sup>2</sup> BALEST 95C quote the fraction of  $\tau^- \rightarrow h^- h^- h^+ \pi^0 \nu_\tau$  (ex.  $K^0$ ) decays which originate in a  $h^- \omega$  final state equals  $0.412 \pm 0.014 \pm 0.015$ . We divide this by the  $\omega(782) \rightarrow \pi^+ \pi^- \pi^0$  branching fraction (0.888).

<sup>3</sup> ALBRECHT 91D quote the fraction of  $\tau^- \rightarrow h^- h^- h^+ \pi^0 \nu_\tau$  decays which originate in a  $\pi^- \omega$  final state equals  $0.33 \pm 0.04 \pm 0.02$ . We divide this by the  $\omega(782) \rightarrow \pi^+ \pi^- \pi^0$  branching fraction (0.888).

$\Gamma(\pi^- \omega \nu_\tau) / \Gamma_{\text{total}}$					$\Gamma_{178} / \Gamma$
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>1.95 ± 0.06 OUR FIT</b>	(Produced by HFLAV)				

$\Gamma(K^- \omega \nu_\tau) / \Gamma_{\text{total}}$   $\Gamma_{179} / \Gamma$ 

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>4.1 ± 0.9 OUR FIT</b>				(Produced by HFLAV)
<b>4.1 ± 0.6 ± 0.7</b>	500	ARMS	05 CLE3	7.6 fb <sup>-1</sup> , $E_{\text{cm}}^{ee} = 10.6$ GeV

 $\Gamma(h^- \omega \pi^0 \nu_\tau) / \Gamma_{\text{total}}$   $\Gamma_{180} / \Gamma$ 

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.41 ± 0.04 OUR FIT</b>				(Produced by HFLAV)
<b>0.43 ± 0.06 ± 0.05</b>	7283	BUSKULIC	97c ALEP	1991–1994 LEP runs

 $\Gamma(h^- \omega \pi^0 \nu_\tau) / \Gamma(h^- h^- h^+ \geq 0 \text{ neutrals} \geq 0 K_L^0 \nu_\tau)$   $\Gamma_{180} / \Gamma_{62}$ 

$$\Gamma_{180} / \Gamma_{62} = \Gamma_{180} / (0.34598\Gamma_{36} + 0.34598\Gamma_{38} + 0.34598\Gamma_{41} + 0.34598\Gamma_{43} + 0.4247\Gamma_{48} + 0.6920\Gamma_{49} + 0.8494\Gamma_{52} + 0.6920\Gamma_{56} + 0.6534\Gamma_{61} + \Gamma_{70} + \Gamma_{78} + \Gamma_{85} + \Gamma_{89} + \Gamma_{97} + \Gamma_{103} + \Gamma_{106} + \Gamma_{107} + 0.2804\Gamma_{150} + 0.2302\Gamma_{151} + 0.2804\Gamma_{152} + 0.2804\Gamma_{154} + 0.3759\Gamma_{156} + 0.3257\Gamma_{160} + 0.7259\Gamma_{170} + 0.9078\Gamma_{178} + 0.9078\Gamma_{179} + 0.9078\Gamma_{180} + 0.892\Gamma_{182})$$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>(2.69 ± 0.28) × 10<sup>-2</sup> OUR FIT</b>				(Produced by HFLAV)

• • • We use the following data for averages but not for fits. • • •

**0.028 ± 0.003 ± 0.003** 430 <sup>1</sup>BORTOLETTO 93 CLEO  $E_{\text{cm}}^{ee} \approx 10.6$  GeV

<sup>1</sup>Not independent of BORTOLETTO 93  $\Gamma(\tau^- \rightarrow h^- \omega \pi^0 \nu_\tau) / \Gamma(\tau^- \rightarrow h^- h^- h^+ 2\pi^0 \nu_\tau \text{ (ex. } K^0))$  value.

 $\Gamma(h^- \omega \pi^0 \nu_\tau) / \Gamma(h^- h^- h^+ 2\pi^0 \nu_\tau \text{ (ex. } K^0))$   $\Gamma_{180} / \Gamma_{84}$ 

$$\Gamma_{180} / \Gamma_{84} = \Gamma_{180} / (\Gamma_{85} + 0.2302\Gamma_{150} + 0.2302\Gamma_{154} + 0.892\Gamma_{180})$$

VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT
<b>82 ± 8 OUR FIT</b>			(Produced by HFLAV)
<b>81 ± 6 ± 6</b>	BORTOLETTO93	CLEO	$E_{\text{cm}}^{ee} \approx 10.6$ GeV

 $\Gamma(h^- \omega 2\pi^0 \nu_\tau) / \Gamma_{\text{total}}$   $\Gamma_{181} / \Gamma$ 

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.4 ± 0.4 ± 0.3</b>	53	ANASTASSOV 01	CLEO	$E_{\text{cm}}^{ee} = 10.6$ GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

1.89<sup>+0.74</sup><sub>-0.67</sub> ± 0.40 19 ANDERSON 97 CLEO Repl. by ANASTASSOV 01

 $\Gamma(\pi^- \omega 2\pi^0 \nu_\tau) / \Gamma_{\text{total}}$   $\Gamma_{182} / \Gamma$ 

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.72 ± 0.16 OUR FIT</b>				(Produced by HFLAV)
<b>0.73 ± 0.12 ± 0.12</b>	1.1k	LEES	12X BABR	468 fb <sup>-1</sup> $E_{\text{cm}}^{ee} = 10.6$ GeV

 $\Gamma(h^- 2\omega \nu_\tau) / \Gamma_{\text{total}}$   $\Gamma_{183} / \Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt; 5.4 × 10<sup>-7</sup></b>	90	AUBERT,B	06 BABR	232 fb <sup>-1</sup> $E_{\text{cm}}^{ee} = 10.6$ GeV

$\Gamma(2h^- h^+ \omega \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{184}/\Gamma$ 

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>1.2 \pm 0.2 \pm 0.1</math></b>	110	ANASTASSOV 01	CLEO	$E_{\text{cm}}^{ee} = 10.6$ GeV

 $\Gamma(2\pi^- \pi^+ \omega \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$   $\Gamma_{185}/\Gamma$ 

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.84 \pm 0.06</math> OUR FIT</b>				(Produced by HFLAV)
<b><math>0.84 \pm 0.04 \pm 0.06</math></b>	2.4k	LEES	12X BABR	$468 \text{ fb}^{-1}$ $E_{\text{cm}}^{ee} = 10.6$ GeV

 $\Gamma(e^- \gamma)/\Gamma_{\text{total}}$   $\Gamma_{186}/\Gamma$ 

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt; 3.3 \times 10^{-8}</math></b>	90	AUBERT	10B BABR	$516 \text{ fb}^{-1}$ , $E_{\text{cm}}^{ee} = 10.6$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$< 5.6 \times 10^{-8}$	90	UNO	21 BELL	$988 \text{ fb}^{-1}$ , $E_{\text{cm}}^{ee} = 10.6$ GeV
$< 1.2 \times 10^{-7}$	90	HAYASAKA	08 BELL	$535 \text{ fb}^{-1}$ , $E_{\text{cm}}^{ee} = 10.6$ GeV
$< 1.1 \times 10^{-7}$	90	AUBERT	06C BABR	$232 \text{ fb}^{-1}$ , $E_{\text{cm}}^{ee} = 10.6$ GeV
$< 3.9 \times 10^{-7}$	90	HAYASAKA	05 BELL	$86.7 \text{ fb}^{-1}$ , $E_{\text{cm}}^{ee} = 10.6$ GeV
$< 2.7 \times 10^{-6}$	90	EDWARDS	97 CLEO	
$< 1.1 \times 10^{-4}$	90	ABREU	95U DLPH	1990–1993 LEP runs
$< 1.2 \times 10^{-4}$	90	ALBRECHT	92K ARG	$E_{\text{cm}}^{ee} = 10$ GeV
$< 2.0 \times 10^{-4}$	90	KEH	88 CBAL	$E_{\text{cm}}^{ee} = 10$ GeV
$< 6.4 \times 10^{-4}$	90	HAYES	82 MRK2	$E_{\text{cm}}^{ee} = 3.8\text{--}6.8$ GeV

 $\Gamma(e^- \gamma \gamma)/\Gamma_{\text{total}}$   $\Gamma_{187}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt; 2.5 \times 10^{-4}</math></b>	90	<sup>1</sup> BRYMAN	21 RVUE	$516 \text{ fb}^{-1}$ , $E_{\text{cm}}^{ee} = 10.6$ GeV

<sup>1</sup> BRYMAN 21 reinterprets the upper limit result on  $B(\tau^- \rightarrow e^- \gamma)$  and  $B(\tau^- \rightarrow \mu^- \gamma)$  by AUBERT 10B, estimating with a simulation the efficiency for this decay mode to be detected as the corresponding AUBERT 10B decay mode.

 $\Gamma(\mu^- \gamma)/\Gamma_{\text{total}}$   $\Gamma_{188}/\Gamma$ 

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt; 4.2 \times 10^{-8}</math></b>	90	UNO	21 BELL	$988 \text{ fb}^{-1}$ , $E_{\text{cm}}^{ee} = 10.6$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$< 4.4 \times 10^{-8}$	90	AUBERT	10B BABR	$516 \text{ fb}^{-1}$ , $E_{\text{cm}}^{ee} = 10.6$ GeV
$< 4.5 \times 10^{-8}$	90	HAYASAKA	08 BELL	$535 \text{ fb}^{-1}$ , $E_{\text{cm}}^{ee} = 10.6$ GeV
$< 6.8 \times 10^{-8}$	90	AUBERT,B	05A BABR	$232 \text{ fb}^{-1}$ , $E_{\text{cm}}^{ee} = 10.6$ GeV
$< 3.1 \times 10^{-7}$	90	ABE	04B BELL	$86.3 \text{ fb}^{-1}$ , $E_{\text{cm}}^{ee} = 10.6$ GeV
$< 1.1 \times 10^{-6}$	90	AHMED	00 CLEO	$E_{\text{cm}}^{ee} = 10.6$ GeV

$< 3.0 \times 10^{-6}$	90	EDWARDS	97	CLEO	
$< 6.2 \times 10^{-5}$	90	ABREU	95U	DLPH	1990–1993 LEP runs
$< 0.42 \times 10^{-5}$	90	BEAN	93	CLEO	$E_{cm}^{ee} = 10.6$ GeV
$< 3.4 \times 10^{-5}$	90	ALBRECHT	92K	ARG	$E_{cm}^{ee} = 10$ GeV
$< 55 \times 10^{-5}$	90	HAYES	82	MRK2	$E_{cm}^{ee} = 3.8\text{--}6.8$ GeV

### $\Gamma(\mu^- \gamma \gamma)/\Gamma_{\text{total}}$ $\Gamma_{189}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt; 5.8 \times 10^{-4}</math></b>	90	<sup>1</sup> BRYMAN	21	RVUE $516 \text{ fb}^{-1}$ , $E_{cm}^{ee} = 10.6$ GeV

<sup>1</sup> BRYMAN 21 reinterprets the upper limit result on  $B(\tau^- \rightarrow e^- \gamma)$  and  $B(\tau^- \rightarrow \mu^- \gamma)$  by AUBERT 10B, estimating with a simulation the efficiency for this decay mode to be detected as the corresponding AUBERT 10B decay mode.

### $\Gamma(e^- \pi^0)/\Gamma_{\text{total}}$ $\Gamma_{190}/\Gamma$

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt; 8.0 \times 10^{-8}</math></b>	90	MIYAZAKI	07	BELL $401 \text{ fb}^{-1}$ , $E_{cm}^{ee} = 10.6$ GeV
• • •		We do not use the following data for averages, fits, limits, etc. • • •		
$< 1.3 \times 10^{-7}$	90	AUBERT	07i	BABR $339 \text{ fb}^{-1}$ , $E_{cm}^{ee} = 10.6$ GeV
$< 1.9 \times 10^{-7}$	90	ENARI	05	BELL $154 \text{ fb}^{-1}$ , $E_{cm}^{ee} = 10.6$ GeV
$< 3.7 \times 10^{-6}$	90	BONVICINI	97	CLEO $E_{cm}^{ee} = 10.6$ GeV
$< 17 \times 10^{-5}$	90	ALBRECHT	92K	ARG $E_{cm}^{ee} = 10$ GeV
$< 14 \times 10^{-5}$	90	KEH	88	CBAL $E_{cm}^{ee} = 10$ GeV
$< 210 \times 10^{-5}$	90	HAYES	82	MRK2 $E_{cm}^{ee} = 3.8\text{--}6.8$ GeV

### $\Gamma(\mu^- \pi^0)/\Gamma_{\text{total}}$ $\Gamma_{191}/\Gamma$

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt; 1.1 \times 10^{-7}</math></b>	90	AUBERT	07i	BABR $339 \text{ fb}^{-1}$ , $E_{cm}^{ee} = 10.6$ GeV
• • •		We do not use the following data for averages, fits, limits, etc. • • •		
$< 1.2 \times 10^{-7}$	90	MIYAZAKI	07	BELL $401 \text{ fb}^{-1}$ , $E_{cm}^{ee} = 10.6$ GeV
$< 4.1 \times 10^{-7}$	90	ENARI	05	BELL $154 \text{ fb}^{-1}$ , $E_{cm}^{ee} = 10.6$ GeV
$< 4.0 \times 10^{-6}$	90	BONVICINI	97	CLEO $E_{cm}^{ee} = 10.6$ GeV
$< 4.4 \times 10^{-5}$	90	ALBRECHT	92K	ARG $E_{cm}^{ee} = 10$ GeV
$< 82 \times 10^{-5}$	90	HAYES	82	MRK2 $E_{cm}^{ee} = 3.8\text{--}6.8$ GeV

### $\Gamma(e^- K_S^0)/\Gamma_{\text{total}}$ $\Gamma_{192}/\Gamma$

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt; 2.6 \times 10^{-8}</math></b>	90	MIYAZAKI	10A	BELL $671 \text{ fb}^{-1}$ $E_{cm}^{ee} = 10.6$ GeV
• • •		We do not use the following data for averages, fits, limits, etc. • • •		
$< 3.3 \times 10^{-8}$	90	AUBERT	09D	BABR $469 \text{ fb}^{-1}$ $E_{cm}^{ee} = 10.6$ GeV
$< 5.6 \times 10^{-8}$	90	MIYAZAKI	06A	BELL $281 \text{ fb}^{-1}$ $E_{cm}^{ee} = 10.6$ GeV
$< 9.1 \times 10^{-7}$	90	CHEN	02C	CLEO $E_{cm}^{ee} = 10.6$ GeV
$< 1.3 \times 10^{-3}$	90	HAYES	82	MRK2 $E_{cm}^{ee} = 3.8\text{--}6.8$ GeV

$\Gamma(\mu^- K_S^0)/\Gamma_{\text{total}}$   $\Gamma_{193}/\Gamma$ 

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 2.3 \times 10^{-8}$	90	MIYAZAKI 10A	BELL	$671 \text{ fb}^{-1}$ $E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
$< 4.0 \times 10^{-8}$	90	AUBERT 09D	BABR	$469 \text{ fb}^{-1}$ $E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 4.9 \times 10^{-8}$	90	MIYAZAKI 06A	BELL	$281 \text{ fb}^{-1}$ $E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 9.5 \times 10^{-7}$	90	CHEN 02C	CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 1.0 \times 10^{-3}$	90	HAYES 82	MRK2	$E_{\text{cm}}^{ee} = 3.8\text{--}6.8 \text{ GeV}$

 $\Gamma(e^- \eta)/\Gamma_{\text{total}}$   $\Gamma_{194}/\Gamma$ 

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 9.2 \times 10^{-8}$	90	MIYAZAKI 07	BELL	$401 \text{ fb}^{-1}$ , $E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
$< 1.6 \times 10^{-7}$	90	AUBERT 07I	BABR	$339 \text{ fb}^{-1}$ , $E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 2.4 \times 10^{-7}$	90	ENARI 05	BELL	$154 \text{ fb}^{-1}$ , $E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 8.2 \times 10^{-6}$	90	BONVICINI 97	CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 6.3 \times 10^{-5}$	90	ALBRECHT 92K	ARG	$E_{\text{cm}}^{ee} = 10 \text{ GeV}$
$< 24 \times 10^{-5}$	90	KEH 88	CBAL	$E_{\text{cm}}^{ee} = 10 \text{ GeV}$

 $\Gamma(\mu^- \eta)/\Gamma_{\text{total}}$   $\Gamma_{195}/\Gamma$ 

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 6.5 \times 10^{-8}$	90	MIYAZAKI 07	BELL	$401 \text{ fb}^{-1}$ , $E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
$< 1.5 \times 10^{-7}$	90	AUBERT 07I	BABR	$339 \text{ fb}^{-1}$ , $E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 1.5 \times 10^{-7}$	90	ENARI 05	BELL	$154 \text{ fb}^{-1}$ , $E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 3.4 \times 10^{-7}$	90	ENARI 04	BELL	$84.3 \text{ fb}^{-1}$ , $E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 9.6 \times 10^{-6}$	90	BONVICINI 97	CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 7.3 \times 10^{-5}$	90	ALBRECHT 92K	ARG	$E_{\text{cm}}^{ee} = 10 \text{ GeV}$

 $\Gamma(e^- \rho^0)/\Gamma_{\text{total}}$   $\Gamma_{196}/\Gamma$ 

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 2.2 \times 10^{-8}$	90	<sup>1</sup> TSUZUKI 23	BELL	$980 \text{ fb}^{-1}$ $E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
$< 1.8 \times 10^{-8}$	90	MIYAZAKI 11	BELL	$854 \text{ fb}^{-1}$ $E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 4.6 \times 10^{-8}$	90	AUBERT 09W	BABR	$451 \text{ fb}^{-1}$ $E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 6.3 \times 10^{-8}$	90	NISHIO 08	BELL	$543 \text{ fb}^{-1}$ $E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 6.5 \times 10^{-7}$	90	YUSA 06	BELL	$158 \text{ fb}^{-1}$ $E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 2.0 \times 10^{-6}$	90	BLISS 98	CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 4.2 \times 10^{-6}$	90	<sup>2</sup> BARTELT 94	CLEO	Repl. by BLISS 98
$< 1.9 \times 10^{-5}$	90	ALBRECHT 92K	ARG	$E_{\text{cm}}^{ee} = 10 \text{ GeV}$
$< 37 \times 10^{-5}$	90	HAYES 82	MRK2	$E_{\text{cm}}^{ee} = 3.8\text{--}6.8 \text{ GeV}$

<sup>1</sup> Supersedes MIYAZAKI 11.<sup>2</sup> BARTELT 94 assume phase space decays.

### $\Gamma(\mu^- \rho^0)/\Gamma_{\text{total}}$ $\Gamma_{197}/\Gamma$

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 1.7 \times 10^{-8}$	90	<sup>1</sup> TSUZUKI	23 BELL	$980 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
••• We do not use the following data for averages, fits, limits, etc. •••				
$< 1.2 \times 10^{-8}$	90	MIYAZAKI	11 BELL	$854 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 2.6 \times 10^{-8}$	90	AUBERT	09W BABR	$451 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 6.8 \times 10^{-8}$	90	NISHIO	08 BELL	$543 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 2.0 \times 10^{-7}$	90	YUSA	06 BELL	$158 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 6.3 \times 10^{-6}$	90	BLISS	98 CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 5.7 \times 10^{-6}$	90	<sup>2</sup> BARTELT	94 CLEO	Repl. by BLISS 98
$< 2.9 \times 10^{-5}$	90	ALBRECHT	92K ARG	$E_{\text{cm}}^{ee} = 10 \text{ GeV}$
$< 44 \times 10^{-5}$	90	HAYES	82 MRK2	$E_{\text{cm}}^{ee} = 3.8\text{--}6.8 \text{ GeV}$

<sup>1</sup>Supersedes MIYAZAKI 11.<sup>2</sup>BARTELT 94 assume phase space decays.

### $\Gamma(e^- \omega)/\Gamma_{\text{total}}$ $\Gamma_{198}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 2.4 \times 10^{-8}$	90	<sup>1</sup> TSUZUKI	23 BELL	$980 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
••• We do not use the following data for averages, fits, limits, etc. •••				
$< 4.8 \times 10^{-8}$	90	MIYAZAKI	11 BELL	$854 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 1.1 \times 10^{-7}$	90	AUBERT	08K BABR	$384 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 1.8 \times 10^{-7}$	90	NISHIO	08 BELL	$543 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

<sup>1</sup>Supersedes MIYAZAKI 11.

### $\Gamma(\mu^- \omega)/\Gamma_{\text{total}}$ $\Gamma_{199}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 3.9 \times 10^{-8}$	90	<sup>1</sup> TSUZUKI	23 BELL	$980 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
••• We do not use the following data for averages, fits, limits, etc. •••				
$< 4.7 \times 10^{-8}$	90	MIYAZAKI	11 BELL	$854 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 1.0 \times 10^{-7}$	90	AUBERT	08K BABR	$384 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 8.9 \times 10^{-8}$	90	NISHIO	08 BELL	$543 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

<sup>1</sup>Supersedes MIYAZAKI 11.

### $\Gamma(e^- K^*(892)^0)/\Gamma_{\text{total}}$ $\Gamma_{200}/\Gamma$

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 1.9 \times 10^{-8}$	90	<sup>1</sup> TSUZUKI	23 BELL	$980 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
••• We do not use the following data for averages, fits, limits, etc. •••				
$< 3.2 \times 10^{-8}$	90	MIYAZAKI	11 BELL	$854 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 5.9 \times 10^{-8}$	90	AUBERT	09W BABR	$451 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 7.8 \times 10^{-8}$	90	NISHIO	08 BELL	$543 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 3.0 \times 10^{-7}$	90	YUSA	06 BELL	$158 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 5.1 \times 10^{-6}$	90	BLISS	98 CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 6.3 \times 10^{-6}$	90	<sup>2</sup> BARTELT	94 CLEO	Repl. by BLISS 98
$< 3.8 \times 10^{-5}$	90	ALBRECHT	92K ARG	$E_{\text{cm}}^{ee} = 10 \text{ GeV}$

<sup>1</sup>Supersedes MIYAZAKI 11.<sup>2</sup>BARTELT 94 assume phase space decays.

$\Gamma(\mu^- K^*(892)^0)/\Gamma_{\text{total}}$   $\Gamma_{201}/\Gamma$ 

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;2.9 × 10<sup>-8</sup></b>	90	<sup>1</sup> TSUZUKI	23	BELL 980 fb <sup>-1</sup> E <sub>cm</sub> <sup>ee</sup> = 10.6 GeV
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
<7.2 × 10 <sup>-8</sup>	90	MIYAZAKI	11	BELL 854 fb <sup>-1</sup> E <sub>cm</sub> <sup>ee</sup> = 10.6 GeV
<1.7 × 10 <sup>-7</sup>	90	AUBERT	09W	BABR 451 fb <sup>-1</sup> E <sub>cm</sub> <sup>ee</sup> = 10.6 GeV
<5.9 × 10 <sup>-8</sup>	90	NISHIO	08	BELL 543 fb <sup>-1</sup> E <sub>cm</sub> <sup>ee</sup> = 10.6 GeV
<3.9 × 10 <sup>-7</sup>	90	YUSA	06	BELL 158 fb <sup>-1</sup> E <sub>cm</sub> <sup>ee</sup> = 10.6 GeV
<7.5 × 10 <sup>-6</sup>	90	BLISS	98	CLEO E <sub>cm</sub> <sup>ee</sup> = 10.6 GeV
<9.4 × 10 <sup>-6</sup>	90	<sup>2</sup> BARTELT	94	CLEO Repl. by BLISS 98
<4.5 × 10 <sup>-5</sup>	90	ALBRECHT	92K	ARG E <sub>cm</sub> <sup>ee</sup> = 10 GeV

<sup>1</sup> Supersedes MIYAZAKI 11.<sup>2</sup> BARTELT 94 assume phase space decays. $\Gamma(e^- \bar{K}^*(892)^0)/\Gamma_{\text{total}}$   $\Gamma_{202}/\Gamma$ 

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;1.7 × 10<sup>-8</sup></b>	90	<sup>1</sup> TSUZUKI	23	BELL 980 fb <sup>-1</sup> E <sub>cm</sub> <sup>ee</sup> = 10.6 GeV
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
<3.4 × 10 <sup>-8</sup>	90	MIYAZAKI	11	BELL 854 fb <sup>-1</sup> E <sub>cm</sub> <sup>ee</sup> = 10.6 GeV
<4.6 × 10 <sup>-8</sup>	90	AUBERT	09W	BABR 451 fb <sup>-1</sup> E <sub>cm</sub> <sup>ee</sup> = 10.6 GeV
<7.7 × 10 <sup>-8</sup>	90	NISHIO	08	BELL 543 fb <sup>-1</sup> E <sub>cm</sub> <sup>ee</sup> = 10.6 GeV
<4.0 × 10 <sup>-7</sup>	90	YUSA	06	BELL 158 fb <sup>-1</sup> E <sub>cm</sub> <sup>ee</sup> = 10.6 GeV
<7.4 × 10 <sup>-6</sup>	90	BLISS	98	CLEO E <sub>cm</sub> <sup>ee</sup> = 10.6 GeV
<1.1 × 10 <sup>-5</sup>	90	<sup>2</sup> BARTELT	94	CLEO Repl. by BLISS 98

<sup>1</sup> Supersedes MIYAZAKI 11.<sup>2</sup> BARTELT 94 assume phase space decays. $\Gamma(\mu^- \bar{K}^*(892)^0)/\Gamma_{\text{total}}$   $\Gamma_{203}/\Gamma$ 

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;4.3 × 10<sup>-8</sup></b>	90	<sup>1</sup> TSUZUKI	23	BELL 980 fb <sup>-1</sup> E <sub>cm</sub> <sup>ee</sup> = 10.6 GeV
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
<7.0 × 10 <sup>-8</sup>	90	MIYAZAKI	11	BELL 854 fb <sup>-1</sup> E <sub>cm</sub> <sup>ee</sup> = 10.6 GeV
<7.3 × 10 <sup>-8</sup>	90	AUBERT	09W	BABR 451 fb <sup>-1</sup> E <sub>cm</sub> <sup>ee</sup> = 10.6 GeV
<1.0 × 10 <sup>-7</sup>	90	NISHIO	08	BELL 543 fb <sup>-1</sup> E <sub>cm</sub> <sup>ee</sup> = 10.6 GeV
<4.0 × 10 <sup>-7</sup>	90	YUSA	06	BELL 158 fb <sup>-1</sup> E <sub>cm</sub> <sup>ee</sup> = 10.6 GeV
<7.5 × 10 <sup>-6</sup>	90	BLISS	98	CLEO E <sub>cm</sub> <sup>ee</sup> = 10.6 GeV
<8.7 × 10 <sup>-6</sup>	90	<sup>2</sup> BARTELT	94	CLEO Repl. by BLISS 98

<sup>1</sup> Supersedes MIYAZAKI 11.<sup>2</sup> BARTELT 94 assume phase space decays. $\Gamma(e^- \eta'(958))/\Gamma_{\text{total}}$   $\Gamma_{204}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt; 1.6 × 10<sup>-7</sup></b>	90	MIYAZAKI	07	BELL 401 fb <sup>-1</sup> , E <sub>cm</sub> <sup>ee</sup> = 10.6 GeV



• • • We do not use the following data for averages, fits, limits, etc. • • •

$< 2.4 \times 10^{-7}$	90	AUBERT	07I	BABR	$339 \text{ fb}^{-1}$	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 10. \times 10^{-7}$	90	ENARI	05	BELL	$154 \text{ fb}^{-1}$	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

### $\Gamma(\mu^- \eta'(958))/\Gamma_{\text{total}}$ $\Gamma_{205}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 1.3 \times 10^{-7}$	90	MIYAZAKI 07	BELL	$401 \text{ fb}^{-1}$ , $E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$< 1.4 \times 10^{-7}$	90	AUBERT	07I	BABR	$339 \text{ fb}^{-1}$	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 4.7 \times 10^{-7}$	90	ENARI	05	BELL	$154 \text{ fb}^{-1}$	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

### $\Gamma(e^- f_0(980) \rightarrow e^- \pi^+ \pi^-)/\Gamma_{\text{total}}$ $\Gamma_{206}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 3.2 \times 10^{-8}$	90	MIYAZAKI 09	BELL	$671 \text{ fb}^{-1}$ , $E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

### $\Gamma(\mu^- f_0(980) \rightarrow \mu^- \pi^+ \pi^-)/\Gamma_{\text{total}}$ $\Gamma_{207}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 3.4 \times 10^{-8}$	90	MIYAZAKI 09	BELL	$671 \text{ fb}^{-1}$ , $E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

### $\Gamma(e^- \phi)/\Gamma_{\text{total}}$ $\Gamma_{208}/\Gamma$

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 2.0 \times 10^{-8}$	90	<sup>1</sup> TSUZUKI 23	BELL	$980 \text{ fb}^{-1}$ , $E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$< 3.1 \times 10^{-8}$	90	MIYAZAKI	11	BELL	$854 \text{ fb}^{-1}$	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 3.1 \times 10^{-8}$	90	AUBERT	09W	BABR	$451 \text{ fb}^{-1}$	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 7.3 \times 10^{-8}$	90	NISHIO	08	BELL	$543 \text{ fb}^{-1}$	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 7.3 \times 10^{-7}$	90	YUSA	06	BELL	$158 \text{ fb}^{-1}$	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 6.9 \times 10^{-6}$	90	BLISS	98	CLEO		$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

<sup>1</sup>Supersedes MIYAZAKI 11.

### $\Gamma(\mu^- \phi)/\Gamma_{\text{total}}$ $\Gamma_{209}/\Gamma$

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 2.3 \times 10^{-8}$	90	<sup>1</sup> TSUZUKI 23	BELL	$980 \text{ fb}^{-1}$ , $E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$< 8.4 \times 10^{-8}$	90	MIYAZAKI	11	BELL	$854 \text{ fb}^{-1}$	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 1.9 \times 10^{-7}$	90	AUBERT	09W	BABR	$451 \text{ fb}^{-1}$	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 1.3 \times 10^{-7}$	90	NISHIO	08	BELL	$543 \text{ fb}^{-1}$	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 7.7 \times 10^{-7}$	90	YUSA	06	BELL	$158 \text{ fb}^{-1}$	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 7.0 \times 10^{-6}$	90	BLISS	98	CLEO		$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

<sup>1</sup>Supersedes MIYAZAKI 11.

$\Gamma(e^- e^+ e^-)/\Gamma_{\text{total}}$  $\Gamma_{210}/\Gamma$ 

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 2.7 \times 10^{-8}$	90	HAYASAKA	10 BELL	$782 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
$< 2.9 \times 10^{-8}$	90	LEES	10A BABR	$468 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 3.6 \times 10^{-8}$	90	MIYAZAKI	08 BELL	$535 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 4.3 \times 10^{-8}$	90	AUBERT	07BK BABR	$376 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 2.0 \times 10^{-7}$	90	AUBERT	04J BABR	$91.5 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 3.5 \times 10^{-7}$	90	YUSA	04 BELL	$87.1 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 2.9 \times 10^{-6}$	90	BLISS	98 CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 0.33 \times 10^{-5}$	90	<sup>1</sup> BARTELT	94 CLEO	Repl. by BLISS 98
$< 1.3 \times 10^{-5}$	90	ALBRECHT	92K ARG	$E_{\text{cm}}^{ee} = 10 \text{ GeV}$
$< 2.7 \times 10^{-5}$	90	BOWCOCK	90 CLEO	$E_{\text{cm}}^{ee} = 10.4\text{--}10.9$
$< 40 \times 10^{-5}$	90	HAYES	82 MRK2	$E_{\text{cm}}^{ee} = 3.8\text{--}6.8 \text{ GeV}$

<sup>1</sup> BARTELT 94 assume phase space decays. $\Gamma(e^- \mu^+ \mu^-)/\Gamma_{\text{total}}$  $\Gamma_{211}/\Gamma$ 

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 2.7 \times 10^{-8}$	90	HAYASAKA	10 BELL	$782 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
$< 3.2 \times 10^{-8}$	90	LEES	10A BABR	$468 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 4.1 \times 10^{-8}$	90	MIYAZAKI	08 BELL	$535 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 3.7 \times 10^{-8}$	90	AUBERT	07BK BABR	$376 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 3.3 \times 10^{-7}$	90	AUBERT	04J BABR	$91.5 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 2.0 \times 10^{-7}$	90	YUSA	04 BELL	$87.1 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 1.8 \times 10^{-6}$	90	BLISS	98 CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 0.36 \times 10^{-5}$	90	<sup>1</sup> BARTELT	94 CLEO	Repl. by BLISS 98
$< 1.9 \times 10^{-5}$	90	ALBRECHT	92K ARG	$E_{\text{cm}}^{ee} = 10 \text{ GeV}$
$< 2.7 \times 10^{-5}$	90	BOWCOCK	90 CLEO	$E_{\text{cm}}^{ee} = 10.4\text{--}10.9$
$< 33 \times 10^{-5}$	90	HAYES	82 MRK2	$E_{\text{cm}}^{ee} = 3.8\text{--}6.8 \text{ GeV}$

<sup>1</sup> BARTELT 94 assume phase space decays. $\Gamma(e^+ \mu^- \mu^-)/\Gamma_{\text{total}}$  $\Gamma_{212}/\Gamma$ 

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 1.7 \times 10^{-8}$	90	HAYASAKA	10 BELL	$782 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
$< 2.6 \times 10^{-8}$	90	LEES	10A BABR	$468 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 2.3 \times 10^{-8}$	90	MIYAZAKI	08 BELL	$535 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 5.6 \times 10^{-8}$	90	AUBERT	07BK BABR	$376 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 1.3 \times 10^{-7}$	90	AUBERT	04J BABR	$91.5 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 2.0 \times 10^{-7}$	90	YUSA	04 BELL	$87.1 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

$<1.5 \times 10^{-6}$	90	BLISS	98	CLEO	$E_{cm}^{ee} = 10.6$ GeV
$<0.35 \times 10^{-5}$	90	<sup>1</sup> BARTELT	94	CLEO	Repl. by BLISS 98
$<1.8 \times 10^{-5}$	90	ALBRECHT	92k	ARG	$E_{cm}^{ee} = 10$ GeV
$<1.6 \times 10^{-5}$	90	BOWCOCK	90	CLEO	$E_{cm}^{ee} = 10.4-10.9$

<sup>1</sup> BARTELT 94 assume phase space decays.

$\Gamma(\mu^- e^+ e^-)/\Gamma_{total}$   $\Gamma_{213}/\Gamma$

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt; 1.8 \times 10^{-8}</math></b>	90	HAYASAKA	10 BELL	$782 \text{ fb}^{-1} E_{cm}^{ee} = 10.6$ GeV
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
$< 2.2 \times 10^{-8}$	90	LEES	10A BABR	$468 \text{ fb}^{-1} E_{cm}^{ee} = 10.6$ GeV
$< 2.7 \times 10^{-8}$	90	MIYAZAKI	08 BELL	$535 \text{ fb}^{-1} E_{cm}^{ee} = 10.6$ GeV
$< 8.0 \times 10^{-8}$	90	AUBERT	07BK BABR	$376 \text{ fb}^{-1} E_{cm}^{ee} = 10.6$ GeV
$< 2.7 \times 10^{-7}$	90	AUBERT	04J BABR	$91.5 \text{ fb}^{-1} E_{cm}^{ee} = 10.6$ GeV
$< 1.9 \times 10^{-7}$	90	YUSA	04 BELL	$87.1 \text{ fb}^{-1} E_{cm}^{ee} = 10.6$ GeV
$< 1.7 \times 10^{-6}$	90	BLISS	98 CLEO	$E_{cm}^{ee} = 10.6$ GeV
$< 0.34 \times 10^{-5}$	90	<sup>1</sup> BARTELT	94 CLEO	Repl. by BLISS 98
$< 1.4 \times 10^{-5}$	90	ALBRECHT	92k ARG	$E_{cm}^{ee} = 10$ GeV
$< 2.7 \times 10^{-5}$	90	BOWCOCK	90 CLEO	$E_{cm}^{ee} = 10.4-10.9$
$< 44 \times 10^{-5}$	90	HAYES	82 MRK2	$E_{cm}^{ee} = 3.8-6.8$ GeV

<sup>1</sup> BARTELT 94 assume phase space decays.

$\Gamma(\mu^+ e^- e^-)/\Gamma_{total}$   $\Gamma_{214}/\Gamma$

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt; 1.5 \times 10^{-8}</math></b>	90	HAYASAKA	10 BELL	$782 \text{ fb}^{-1} E_{cm}^{ee} = 10.6$ GeV
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
$< 1.8 \times 10^{-8}$	90	LEES	10A BABR	$468 \text{ fb}^{-1} E_{cm}^{ee} = 10.6$ GeV
$< 2.0 \times 10^{-8}$	90	MIYAZAKI	08 BELL	$535 \text{ fb}^{-1} E_{cm}^{ee} = 10.6$ GeV
$< 5.8 \times 10^{-8}$	90	AUBERT	07BK BABR	$376 \text{ fb}^{-1} E_{cm}^{ee} = 10.6$ GeV
$< 1.1 \times 10^{-7}$	90	AUBERT	04J BABR	$91.5 \text{ fb}^{-1} E_{cm}^{ee} = 10.6$ GeV
$< 2.0 \times 10^{-7}$	90	YUSA	04 BELL	$87.1 \text{ fb}^{-1} E_{cm}^{ee} = 10.6$ GeV
$< 1.5 \times 10^{-6}$	90	BLISS	98 CLEO	$E_{cm}^{ee} = 10.6$ GeV
$< 0.34 \times 10^{-5}$	90	<sup>1</sup> BARTELT	94 CLEO	Repl. by BLISS 98
$< 1.4 \times 10^{-5}$	90	ALBRECHT	92k ARG	$E_{cm}^{ee} = 10$ GeV
$< 1.6 \times 10^{-5}$	90	BOWCOCK	90 CLEO	$E_{cm}^{ee} = 10.4-10.9$

<sup>1</sup> BARTELT 94 assume phase space decays.

$\Gamma(\mu^- \mu^+ \mu^-)/\Gamma_{total}$   $\Gamma_{215}/\Gamma$

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt; 2.1 \times 10^{-8}</math></b>	90	HAYASAKA	10 BELL	$782 \text{ fb}^{-1} E_{cm}^{ee} = 10.6$ GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

$< 8.0 \times 10^{-8}$	90	SIRUNYAN	21D	CMS	$33.2 \text{ fb}^{-1}$ , $pp$ at 13 TeV
$< 3.8 \times 10^{-7}$	90	AAD	16BA	ATLS	$20.3 \text{ fb}^{-1}$ , $pp$ at 8 TeV
$< 4.6 \times 10^{-8}$	90	AAIJ	15AI	LHCB	$3.0 \text{ fb}^{-1}$ , $pp$ at 7, 8 TeV
$< 8.0 \times 10^{-8}$	90	<sup>1</sup> AAIJ	13AH	LHCB	$1.0 \text{ fb}^{-1}$ , $pp$ at 7 TeV
$< 3.3 \times 10^{-8}$	90	LEES	10A	BABR	$468 \text{ fb}^{-1}$ $E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 3.2 \times 10^{-8}$	90	MIYAZAKI	08	BELL	$535 \text{ fb}^{-1}$ $E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 5.3 \times 10^{-8}$	90	AUBERT	07BK	BABR	$376 \text{ fb}^{-1}$ $E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 1.9 \times 10^{-7}$	90	AUBERT	04J	BABR	$91.5 \text{ fb}^{-1}$ $E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 2.0 \times 10^{-7}$	90	YUSA	04	BELL	$87.1 \text{ fb}^{-1}$ $E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 1.9 \times 10^{-6}$	90	BLISS	98	CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 0.43 \times 10^{-5}$	90	<sup>2</sup> BARTELT	94	CLEO	Repl. by BLISS 98
$< 1.9 \times 10^{-5}$	90	ALBRECHT	92K	ARG	$E_{\text{cm}}^{ee} = 10 \text{ GeV}$
$< 1.7 \times 10^{-5}$	90	BOWCOCK	90	CLEO	$E_{\text{cm}}^{ee} = 10.4\text{--}10.9$
$< 49 \times 10^{-5}$	90	HAYES	82	MRK2	$E_{\text{cm}}^{ee} = 3.8\text{--}6.8 \text{ GeV}$

<sup>1</sup> Repl. by AAIJ 15AI.

<sup>2</sup> BARTELT 94 assume phase space decays.

### $\Gamma(e^- \pi^+ \pi^-) / \Gamma_{\text{total}}$

Test of lepton family number conservation.

$\Gamma_{216} / \Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt; 2.3 \times 10^{-8}</math></b>	90	MIYAZAKI	13	BELL $854 \text{ fb}^{-1}$ $E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$< 4.4 \times 10^{-8}$	90	MIYAZAKI	10	BELL	Repl. by MIYAZAKI 13
$< 7.3 \times 10^{-7}$	90	YUSA	06	BELL	$158 \text{ fb}^{-1}$ $E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 1.2 \times 10^{-7}$	90	AUBERT, BE	05D	BABR	$221 \text{ fb}^{-1}$ , $E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 2.2 \times 10^{-6}$	90	BLISS	98	CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 4.4 \times 10^{-6}$	90	<sup>1</sup> BARTELT	94	CLEO	Repl. by BLISS 98
$< 2.7 \times 10^{-5}$	90	ALBRECHT	92K	ARG	$E_{\text{cm}}^{ee} = 10 \text{ GeV}$
$< 6.0 \times 10^{-5}$	90	BOWCOCK	90	CLEO	$E_{\text{cm}}^{ee} = 10.4\text{--}10.9$

<sup>1</sup> BARTELT 94 assume phase space decays.

### $\Gamma(e^+ \pi^- \pi^-) / \Gamma_{\text{total}}$

Test of lepton number conservation.

$\Gamma_{217} / \Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt; 2.0 \times 10^{-8}</math></b>	90	MIYAZAKI	13	BELL $854 \text{ fb}^{-1}$ $E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$< 8.8 \times 10^{-8}$	90	MIYAZAKI	10	BELL	Repl. by MIYAZAKI 13
$< 2.0 \times 10^{-7}$	90	YUSA	06	BELL	$158 \text{ fb}^{-1}$ $E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 2.7 \times 10^{-7}$	90	AUBERT, BE	05D	BABR	$221 \text{ fb}^{-1}$ , $E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 1.9 \times 10^{-6}$	90	BLISS	98	CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 4.4 \times 10^{-6}$	90	<sup>1</sup> BARTELT	94	CLEO	Repl. by BLISS 98
$< 1.8 \times 10^{-5}$	90	ALBRECHT	92K	ARG	$E_{\text{cm}}^{ee} = 10 \text{ GeV}$
$< 1.7 \times 10^{-5}$	90	BOWCOCK	90	CLEO	$E_{\text{cm}}^{ee} = 10.4\text{--}10.9$

<sup>1</sup> BARTELT 94 assume phase space decays.

$\Gamma(\mu^- \pi^+ \pi^-)/\Gamma_{\text{total}}$   $\Gamma_{218}/\Gamma$ 

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<2.1 \times 10^{-8}$	90	MIYAZAKI 13	BELL	$854 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<3.3 \times 10^{-8}$	90	MIYAZAKI 10	BELL	Repl. by MIYAZAKI 13
$<4.8 \times 10^{-7}$	90	YUSA 06	BELL	$158 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$<2.9 \times 10^{-7}$	90	AUBERT, BE 05D	BABR	$221 \text{ fb}^{-1}, E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$<8.2 \times 10^{-6}$	90	BLISS 98	CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$<7.4 \times 10^{-6}$	90	<sup>1</sup> BARTELT 94	CLEO	Repl. by BLISS 98
$<3.6 \times 10^{-5}$	90	ALBRECHT 92K	ARG	$E_{\text{cm}}^{ee} = 10 \text{ GeV}$
$<3.9 \times 10^{-5}$	90	BOWCOCK 90	CLEO	$E_{\text{cm}}^{ee} = 10.4\text{--}10.9$

<sup>1</sup> BARTELT 94 assume phase space decays. $\Gamma(\mu^+ \pi^- \pi^-)/\Gamma_{\text{total}}$   $\Gamma_{219}/\Gamma$ 

Test of lepton number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<3.9 \times 10^{-8}$	90	MIYAZAKI 13	BELL	$854 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<3.7 \times 10^{-8}$	90	MIYAZAKI 10	BELL	Repl. by MIYAZAKI 13
$<3.4 \times 10^{-7}$	90	YUSA 06	BELL	$158 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$<7 \times 10^{-8}$	90	AUBERT, BE 05D	BABR	$221 \text{ fb}^{-1}, E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$<3.4 \times 10^{-6}$	90	BLISS 98	CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$<6.9 \times 10^{-6}$	90	<sup>1</sup> BARTELT 94	CLEO	Repl. by BLISS 98
$<6.3 \times 10^{-5}$	90	ALBRECHT 92K	ARG	$E_{\text{cm}}^{ee} = 10 \text{ GeV}$
$<3.9 \times 10^{-5}$	90	BOWCOCK 90	CLEO	$E_{\text{cm}}^{ee} = 10.4\text{--}10.9$

<sup>1</sup> BARTELT 94 assume phase space decays. $\Gamma(e^- \pi^+ K^-)/\Gamma_{\text{total}}$   $\Gamma_{220}/\Gamma$ 

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<3.7 \times 10^{-8}$	90	MIYAZAKI 13	BELL	$854 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<5.8 \times 10^{-8}$	90	MIYAZAKI 10	BELL	Repl. by MIYAZAKI 13
$<7.2 \times 10^{-7}$	90	YUSA 06	BELL	$158 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$<3.2 \times 10^{-7}$	90	AUBERT, BE 05D	BABR	$221 \text{ fb}^{-1}, E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$<6.4 \times 10^{-6}$	90	BLISS 98	CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$<7.7 \times 10^{-6}$	90	<sup>1</sup> BARTELT 94	CLEO	Repl. by BLISS 98
$<2.9 \times 10^{-5}$	90	ALBRECHT 92K	ARG	$E_{\text{cm}}^{ee} = 10 \text{ GeV}$
$<5.8 \times 10^{-5}$	90	BOWCOCK 90	CLEO	$E_{\text{cm}}^{ee} = 10.4\text{--}10.9$

<sup>1</sup> BARTELT 94 assume phase space decays. $\Gamma(e^- \pi^- K^+)/\Gamma_{\text{total}}$   $\Gamma_{221}/\Gamma$ 

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<3.1 \times 10^{-8}$	90	MIYAZAKI 13	BELL	$854 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$<5.2 \times 10^{-8}$	90	MIYAZAKI	10	BELL	Repl. by MIYAZAKI 13
$<1.6 \times 10^{-7}$	90	YUSA	06	BELL	$158 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$<1.7 \times 10^{-7}$	90	AUBERT,BE	05D	BABR	$221 \text{ fb}^{-1}, E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$<3.8 \times 10^{-6}$	90	BLISS	98	CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$<4.6 \times 10^{-6}$	90	<sup>1</sup> BARTELT	94	CLEO	Repl. by BLISS 98
$<5.8 \times 10^{-5}$	90	BOWCOCK	90	CLEO	$E_{\text{cm}}^{ee} = 10.4\text{--}10.9$

<sup>1</sup> BARTELT 94 assume phase space decays.

### $\Gamma(e^+ \pi^- K^-)/\Gamma_{\text{total}}$ $\Gamma_{222}/\Gamma$

Test of lepton number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;3.2 \times 10^{-8}</math></b>	90	MIYAZAKI	13	BELL $854 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$<6.7 \times 10^{-8}$	90	MIYAZAKI	10	BELL	Repl. by MIYAZAKI 13
$<1.9 \times 10^{-7}$	90	YUSA	06	BELL	$158 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$<1.8 \times 10^{-7}$	90	AUBERT,BE	05D	BABR	$221 \text{ fb}^{-1}, E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$<2.1 \times 10^{-6}$	90	BLISS	98	CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$<4.5 \times 10^{-6}$	90	<sup>1</sup> BARTELT	94	CLEO	Repl. by BLISS 98
$<2.0 \times 10^{-5}$	90	ALBRECHT	92K	ARG	$E_{\text{cm}}^{ee} = 10 \text{ GeV}$
$<4.9 \times 10^{-5}$	90	BOWCOCK	90	CLEO	$E_{\text{cm}}^{ee} = 10.4\text{--}10.9$

<sup>1</sup> BARTELT 94 assume phase space decays.

### $\Gamma(e^- K_S^0 K_S^0)/\Gamma_{\text{total}}$ $\Gamma_{223}/\Gamma$

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;7.1 \times 10^{-8}</math></b>	90	MIYAZAKI	10A	BELL $671 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$<2.2 \times 10^{-6}$	90	CHEN	02c	CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
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### $\Gamma(e^- K^+ K^-)/\Gamma_{\text{total}}$ $\Gamma_{224}/\Gamma$

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;3.4 \times 10^{-8}</math></b>	90	MIYAZAKI	13	BELL $854 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$<5.4 \times 10^{-8}$	90	MIYAZAKI	10	BELL	Repl. by MIYAZAKI 13
$<3.0 \times 10^{-7}$	90	YUSA	06	BELL	$158 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$<1.4 \times 10^{-7}$	90	AUBERT,BE	05D	BABR	$221 \text{ fb}^{-1}, E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$<6.0 \times 10^{-6}$	90	BLISS	98	CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

### $\Gamma(e^+ K^- K^-)/\Gamma_{\text{total}}$ $\Gamma_{225}/\Gamma$

Test of lepton number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;3.3 \times 10^{-8}</math></b>	90	MIYAZAKI	13	BELL $854 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$<6.0 \times 10^{-8}$	90	MIYAZAKI	10	BELL	Repl. by MIYAZAKI 13
$<3.1 \times 10^{-7}$	90	YUSA	06	BELL	$158 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$<1.5 \times 10^{-7}$	90	AUBERT,BE	05D	BABR	$221 \text{ fb}^{-1}, E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$<3.8 \times 10^{-6}$	90	BLISS	98	CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

$\Gamma(\mu^- \pi^+ K^-)/\Gamma_{\text{total}}$   $\Gamma_{226}/\Gamma$ 

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 8.6 \times 10^{-8}$	90	MIYAZAKI 13	BELL	$854 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
$< 1.6 \times 10^{-7}$	90	MIYAZAKI 10	BELL	Repl. by MIYAZAKI 13
$< 2.7 \times 10^{-7}$	90	YUSA 06	BELL	$158 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 2.6 \times 10^{-7}$	90	AUBERT, BE 05D	BABR	$221 \text{ fb}^{-1}, E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 7.5 \times 10^{-6}$	90	BLISS 98	CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 8.7 \times 10^{-6}$	90	<sup>1</sup> BARTELT 94	CLEO	Repl. by BLISS 98
$< 11 \times 10^{-5}$	90	ALBRECHT 92K	ARG	$E_{\text{cm}}^{ee} = 10 \text{ GeV}$
$< 7.7 \times 10^{-5}$	90	BOWCOCK 90	CLEO	$E_{\text{cm}}^{ee} = 10.4\text{--}10.9$

<sup>1</sup> BARTELT 94 assume phase space decays. $\Gamma(\mu^- \pi^- K^+)/\Gamma_{\text{total}}$   $\Gamma_{227}/\Gamma$ 

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 4.5 \times 10^{-8}$	90	MIYAZAKI 13	BELL	$854 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
$< 1.0 \times 10^{-7}$	90	MIYAZAKI 10	BELL	Repl. by MIYAZAKI 13
$< 7.3 \times 10^{-7}$	90	YUSA 06	BELL	$158 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 3.2 \times 10^{-7}$	90	AUBERT, BE 05D	BABR	$221 \text{ fb}^{-1}, E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 7.4 \times 10^{-6}$	90	BLISS 98	CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 1.5 \times 10^{-5}$	90	<sup>1</sup> BARTELT 94	CLEO	Repl. by BLISS 98
$< 7.7 \times 10^{-5}$	90	BOWCOCK 90	CLEO	$E_{\text{cm}}^{ee} = 10.4\text{--}10.9$

<sup>1</sup> BARTELT 94 assume phase space decays. $\Gamma(\mu^+ \pi^- K^-)/\Gamma_{\text{total}}$   $\Gamma_{228}/\Gamma$ 

Test of lepton number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 4.8 \times 10^{-8}$	90	MIYAZAKI 13	BELL	$854 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
$< 9.4 \times 10^{-8}$	90	MIYAZAKI 10	BELL	Repl. by MIYAZAKI 13
$< 2.9 \times 10^{-7}$	90	YUSA 06	BELL	$158 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 2.2 \times 10^{-7}$	90	AUBERT, BE 05D	BABR	$221 \text{ fb}^{-1}, E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 7.0 \times 10^{-6}$	90	BLISS 98	CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 2.0 \times 10^{-5}$	90	<sup>1</sup> BARTELT 94	CLEO	Repl. by BLISS 98
$< 5.8 \times 10^{-5}$	90	ALBRECHT 92K	ARG	$E_{\text{cm}}^{ee} = 10 \text{ GeV}$
$< 4.0 \times 10^{-5}$	90	BOWCOCK 90	CLEO	$E_{\text{cm}}^{ee} = 10.4\text{--}10.9$

<sup>1</sup> BARTELT 94 assume phase space decays. $\Gamma(\mu^- K_S^0 K_S^0)/\Gamma_{\text{total}}$   $\Gamma_{229}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 8.0 \times 10^{-8}$	90	MIYAZAKI 10A	BELL	$671 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
$< 3.4 \times 10^{-6}$	90	CHEN 02C	CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

$\Gamma(\mu^- K^+ K^-)/\Gamma_{\text{total}}$   $\Gamma_{230}/\Gamma$ 

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 4.4 \times 10^{-8}$	90	MIYAZAKI 13	BELL	$854 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$< 6.8 \times 10^{-8}$	90	MIYAZAKI 10	BELL	Repl. by MIYAZAKI 13
$< 8.0 \times 10^{-7}$	90	YUSA 06	BELL	$158 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 2.5 \times 10^{-7}$	90	AUBERT, BE 05D	BABR	$221 \text{ fb}^{-1}, E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 15 \times 10^{-6}$	90	BLISS 98	CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

 $\Gamma(\mu^+ K^- K^-)/\Gamma_{\text{total}}$   $\Gamma_{231}/\Gamma$ 

Test of lepton number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 4.7 \times 10^{-8}$	90	MIYAZAKI 13	BELL	$854 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$< 9.6 \times 10^{-8}$	90	MIYAZAKI 10	BELL	Repl. by MIYAZAKI 13
$< 4.4 \times 10^{-7}$	90	YUSA 06	BELL	$158 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 4.8 \times 10^{-7}$	90	AUBERT, BE 05D	BABR	$221 \text{ fb}^{-1}, E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$< 6.0 \times 10^{-6}$	90	BLISS 98	CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

 $\Gamma(e^- \pi^0 \pi^0)/\Gamma_{\text{total}}$   $\Gamma_{232}/\Gamma$ 

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 6.5 \times 10^{-6}$	90	BONVICINI 97	CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

 $\Gamma(\mu^- \pi^0 \pi^0)/\Gamma_{\text{total}}$   $\Gamma_{233}/\Gamma$ 

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 14 \times 10^{-6}$	90	BONVICINI 97	CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

 $\Gamma(e^- \eta \eta)/\Gamma_{\text{total}}$   $\Gamma_{234}/\Gamma$ 

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 35 \times 10^{-6}$	90	BONVICINI 97	CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

 $\Gamma(\mu^- \eta \eta)/\Gamma_{\text{total}}$   $\Gamma_{235}/\Gamma$ 

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 60 \times 10^{-6}$	90	BONVICINI 97	CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

 $\Gamma(e^- \pi^0 \eta)/\Gamma_{\text{total}}$   $\Gamma_{236}/\Gamma$ 

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 24 \times 10^{-6}$	90	BONVICINI 97	CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

 $\Gamma(\mu^- \pi^0 \eta)/\Gamma_{\text{total}}$   $\Gamma_{237}/\Gamma$ 

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 22 \times 10^{-6}$	90	BONVICINI 97	CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$



$\Gamma(pe^-e^-)/\Gamma_{\text{total}}$   $\Gamma_{238}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<3.0 \times 10^{-8}$	90	SAHOO 20	BELL	$921 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

 $\Gamma(\bar{p}e^+e^-)/\Gamma_{\text{total}}$   $\Gamma_{239}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<3.0 \times 10^{-8}$	90	SAHOO 20	BELL	$921 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

 $\Gamma(\bar{p}e^+\mu^-)/\Gamma_{\text{total}}$   $\Gamma_{240}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<2.0 \times 10^{-8}$	90	SAHOO 20	BELL	$921 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

 $\Gamma(\bar{p}e^-\mu^+)/\Gamma_{\text{total}}$   $\Gamma_{241}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<1.8 \times 10^{-8}$	90	SAHOO 20	BELL	$921 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

 $\Gamma(p\mu^-\mu^-)/\Gamma_{\text{total}}$   $\Gamma_{242}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<4.0 \times 10^{-8}$	90	SAHOO 20	BELL	$921 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$<4.4 \times 10^{-7}$	90	AAIJ 13AH	LHCB	$1.0 \text{ fb}^{-1}, pp \text{ at } 7 \text{ TeV}$
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 $\Gamma(\bar{p}\mu^+\mu^-)/\Gamma_{\text{total}}$   $\Gamma_{243}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<1.8 \times 10^{-8}$	90	SAHOO 20	BELL	$921 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$<3.3 \times 10^{-7}$	90	AAIJ 13AH	LHCB	$1.0 \text{ fb}^{-1}, pp \text{ at } 7 \text{ TeV}$
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 $\Gamma(\bar{p}\gamma)/\Gamma_{\text{total}}$   $\Gamma_{244}/\Gamma$ 

Test of lepton number and baryon number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<3.5 \times 10^{-6}$	90	GODANG 99	CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$<29 \times 10^{-5}$	90	ALBRECHT 92K	ARG	$E_{\text{cm}}^{ee} = 10 \text{ GeV}$
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 $\Gamma(\bar{p}\pi^0)/\Gamma_{\text{total}}$   $\Gamma_{245}/\Gamma$ 

Test of lepton number and baryon number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<15 \times 10^{-6}$	90	GODANG 99	CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$<66 \times 10^{-5}$	90	ALBRECHT 92K	ARG	$E_{\text{cm}}^{ee} = 10 \text{ GeV}$
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 $\Gamma(\bar{p}2\pi^0)/\Gamma_{\text{total}}$   $\Gamma_{246}/\Gamma$ 

Test of lepton number and baryon number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<33 \times 10^{-6}$	90	GODANG 99	CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

$\Gamma(\bar{p}\eta)/\Gamma_{\text{total}}$   $\Gamma_{247}/\Gamma$ 

Test of lepton number and baryon number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 8.9 \times 10^{-6}$	90	GODANG	99	CLEO $E_{\text{cm}}^{ee} = 10.6$ GeV
$< 130 \times 10^{-5}$	90	ALBRECHT	92K	ARG $E_{\text{cm}}^{ee} = 10$ GeV

 $\Gamma(\bar{p}\pi^0\eta)/\Gamma_{\text{total}}$   $\Gamma_{248}/\Gamma$ 

Test of lepton number and baryon number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 27 \times 10^{-6}$	90	GODANG	99	CLEO $E_{\text{cm}}^{ee} = 10.6$ GeV

 $\Gamma(\Lambda\pi^-)/\Gamma_{\text{total}}$   $\Gamma_{249}/\Gamma$ 

Test of lepton number and baryon number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 0.72 \times 10^{-7}$	90	MIYAZAKI	06	BELL $154 \text{ fb}^{-1}$ , $E_{\text{cm}}^{ee} = 10.6$ GeV

 $\Gamma(\bar{\Lambda}\pi^-)/\Gamma_{\text{total}}$   $\Gamma_{250}/\Gamma$ 

Test of lepton number and baryon number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 1.4 \times 10^{-7}$	90	MIYAZAKI	06	BELL $154 \text{ fb}^{-1}$ , $E_{\text{cm}}^{ee} = 10.6$ GeV

 $\Gamma(e^- \text{ light boson})/\Gamma(e^- \bar{\nu}_e \nu_\tau)$   $\Gamma_{251}/\Gamma_5$ 

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 5.3 \times 10^{-3}$	95	<sup>1</sup> ADACHI	23A	BELL $E_{\text{cm}}^{ee} = 10.58$ GeV
$< 0.008$	95	<sup>2</sup> BRYMAN	21	RVUE
$< 0.015$	95	<sup>3</sup> ALBRECHT	95G	ARG $E_{\text{cm}}^{ee} = 9.4\text{--}10.6$ GeV
$< 0.018$	95	<sup>4</sup> ALBRECHT	90E	ARG $E_{\text{cm}}^{ee} = 9.4\text{--}10.6$ GeV
$< 0.040$	95	<sup>5</sup> BALTRUSAITIS	85	MRK3 $E_{\text{cm}}^{ee} = 3.77$ GeV

<sup>1</sup> ADACHI 23A limit holds for bosons with mass = 0.0 GeV. The limit rises to  $9.7 \times 10^{-3}$  for a mass of 1.0 GeV, then falls to  $1.1 \times 10^{-3}$  at the upper mass limit of 1.6 GeV.

<sup>2</sup> BRYMAN 21 reports indirect limits obtained from the consistency of the world averages of  $B(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)$  and  $B(\tau^- \rightarrow \mu^- \bar{\nu}_\mu \nu_\tau)$  with their Standard Model predictions, without a simulation of the efficiency as a function of the X mass for the searched decay modes to be detected as the corresponding Standard Model decay modes.

<sup>3</sup> ALBRECHT 95G limit holds for bosons with mass < 0.4 GeV. The limit rises to 0.036 for a mass of 1.0 GeV, then falls to 0.006 at the upper mass limit of 1.6 GeV.

<sup>4</sup> ALBRECHT 90E limit applies for spinless boson with mass < 100 MeV, and rises to 0.050 for mass = 500 MeV.

<sup>5</sup> BALTRUSAITIS 85 limit applies for spinless boson with mass < 100 MeV.

 $\Gamma(\mu^- \text{ light boson})/\Gamma(e^- \bar{\nu}_e \nu_\tau)$   $\Gamma_{252}/\Gamma_5$ 

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 3.4 \times 10^{-3}$	95	<sup>1</sup> ADACHI	23A	BELL $E_{\text{cm}}^{ee} = 10.58$ GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

<0.011	95	<sup>2</sup> BRYMAN	21	RVUE	
<0.026	95	<sup>3</sup> ALBRECHT	95G	ARG	$E_{\text{cm}}^{ee} = 9.4\text{--}10.6$ GeV
<0.033	95	<sup>4</sup> ALBRECHT	90E	ARG	$E_{\text{cm}}^{ee} = 9.4\text{--}10.6$ GeV
<0.125	95	<sup>5</sup> BALTRUSAITIS	85	MRK3	$E_{\text{cm}}^{ee} = 3.77$ GeV

<sup>1</sup> ADACHI 23A limit holds for bosons with mass = 0.0 GeV. The limit rises to  $12.2 \times 10^{-3}$  for a mass of 1.0 GeV, then falls to  $0.7 \times 10^{-3}$  at the upper mass limit of 1.6 GeV.

<sup>2</sup> BRYMAN 21 reports indirect limits obtained from the consistency of the world averages of  $B(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)$  and  $B(\tau^- \rightarrow \mu^- \bar{\nu}_\mu \nu_\tau)$  with their Standard Model predictions, without a simulation of the efficiency as a function of the  $X$  mass for the searched decay modes to be detected as the corresponding Standard Model decay modes.

<sup>3</sup> ALBRECHT 95G limit holds for bosons with mass < 1.3 GeV. The limit rises to 0.034 for a mass of 1.4 GeV, then falls to 0.003 at the upper mass limit of 1.6 GeV.

<sup>4</sup> ALBRECHT 90E limit applies for spinless boson with mass < 100 MeV, and rises to 0.071 for mass = 500 MeV.

<sup>5</sup> BALTRUSAITIS 85 limit applies for spinless boson with mass < 100 MeV.

### $\tau$ -DECAY PARAMETERS

See the related review(s):

[\$\tau\$ -Lepton Decay Parameters](#)

#### $\rho(e \text{ or } \mu)$ PARAMETER

( $V-A$ ) theory predicts  $\rho = 0.75$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.745 ± 0.008 OUR FIT</b>				
<b>0.749 ± 0.008 OUR AVERAGE</b>				
0.742 ± 0.014 ± 0.006	81k	HEISTER	01E ALEP	1991–1995 LEP runs
0.775 ± 0.023 ± 0.020	36k	ABREU	00L DLPH	1992–1995 runs
0.781 ± 0.028 ± 0.018	46k	ACKERSTAFF	99D OPAL	1990–1995 LEP runs
0.762 ± 0.035	54k	ACCIARRI	98R L3	1991–1995 LEP runs
0.731 ± 0.031		<sup>1</sup> ALBRECHT	98 ARG	$E_{\text{cm}}^{ee} = 9.5\text{--}10.6$ GeV
0.72 ± 0.09 ± 0.03		<sup>2</sup> ABE	97O SLD	1993–1995 SLC runs
0.747 ± 0.010 ± 0.006	55k	ALEXANDER	97F CLEO	$E_{\text{cm}}^{ee} = 10.6$ GeV
0.79 ± 0.10 ± 0.10	3732	FORD	87B MAC	$E_{\text{cm}}^{ee} = 29$ GeV
0.71 ± 0.09 ± 0.03	1426	BEHRENDIS	85 CLEO	$e^+ e^-$ near $\Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.735 ± 0.013 ± 0.008	31k	AMMAR	97B CLEO	Repl. by ALEXANDER 97F
0.794 ± 0.039 ± 0.031	18k	ACCIARRI	96H L3	Repl. by ACCIARRI 98R
0.732 ± 0.034 ± 0.020	8.2k	<sup>3</sup> ALBRECHT	95 ARG	$E_{\text{cm}}^{ee} = 9.5\text{--}10.6$ GeV
0.738 ± 0.038		<sup>4</sup> ALBRECHT	95C ARG	Repl. by ALBRECHT 98
0.751 ± 0.039 ± 0.022		BUSKULIC	95D ALEP	Repl. by HEISTER 01E
0.742 ± 0.035 ± 0.020	8000	ALBRECHT	90E ARG	$E_{\text{cm}}^{ee} = 9.4\text{--}10.6$ GeV

<sup>1</sup> Combined fit to ARGUS tau decay parameter measurements in ALBRECHT 98, ALBRECHT 95C, ALBRECHT 93G, and ALBRECHT 94E. ALBRECHT 98 use tau pair events of the type  $\tau^- \tau^+ \rightarrow (\ell^- \bar{\nu}_\ell \nu_\tau)(\pi^+ \pi^0 \bar{\nu}_\tau)$ , and their charged conjugates.

<sup>2</sup> ABE 97O assume  $\eta = 0$  in their fit. Letting  $\eta$  vary in the fit gives a  $\rho$  value of  $0.69 \pm 0.13 \pm 0.05$ .

<sup>3</sup> Value is from a simultaneous fit for the  $\rho$  and  $\eta$  decay parameters to the lepton energy spectrum. Not independent of ALBRECHT 90E  $\rho(e \text{ or } \mu)$  value which assumes  $\eta = 0$ . Result is strongly correlated with ALBRECHT 95C.

<sup>4</sup> Combined fit to ARGUS tau decay parameter measurements in ALBRECHT 95C, ALBRECHT 93G, and ALBRECHT 94E.

## $\rho(e)$ PARAMETER

( $V-A$ ) theory predicts  $\rho = 0.75$ .

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.747±0.010 OUR FIT</b>				
<b>0.744±0.010 OUR AVERAGE</b>				
0.747±0.019±0.014	44k	HEISTER	01E ALEP	1991–1995 LEP runs
0.744±0.036±0.037	17k	ABREU	00L DLPH	1992–1995 runs
0.779±0.047±0.029	25k	ACKERSTAFF	99D OPAL	1990–1995 LEP runs
0.68 ±0.04 ±0.07		<sup>1</sup> ALBRECHT	98 ARG	$E_{\text{cm}}^{ee} = 9.5\text{--}10.6$ GeV
0.71 ±0.14 ±0.05		ABE	97O SLD	1993–1995 SLC runs
0.747±0.012±0.004	34k	ALEXANDER	97F CLEO	$E_{\text{cm}}^{ee} = 10.6$ GeV
0.735±0.036±0.020	4.7k	<sup>2</sup> ALBRECHT	95 ARG	$E_{\text{cm}}^{ee} = 9.5\text{--}10.6$ GeV
0.79 ±0.08 ±0.06	3230	<sup>3</sup> ALBRECHT	93G ARG	$E_{\text{cm}}^{ee} = 9.4\text{--}10.6$ GeV
0.64 ±0.06 ±0.07	2753	JANSSEN	89 CBAL	$E_{\text{cm}}^{ee} = 9.4\text{--}10.6$ GeV
0.62 ±0.17 ±0.14	1823	FORD	87B MAC	$E_{\text{cm}}^{ee} = 29$ GeV
0.60 ±0.13	699	BEHRENDIS	85 CLEO	$e^+e^-$ near $\Upsilon(4S)$
0.72 ±0.10 ±0.11	594	BACINO	79B DLCO	$E_{\text{cm}}^{ee} = 3.5\text{--}7.4$ GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.732±0.014±0.009	19k	AMMAR	97B CLEO	Repl. by ALEXANDER 97F
0.793±0.050±0.025		BUSKULIC	95D ALEP	Repl. by HEISTER 01E
0.747±0.045±0.028	5106	ALBRECHT	90E ARG	Repl. by ALBRECHT 95

<sup>1</sup> ALBRECHT 98 use tau pair events of the type  $\tau^- \tau^+ \rightarrow (\ell^- \bar{\nu}_\ell \nu_\tau)(\pi^+ \pi^0 \bar{\nu}_\tau)$ , and their charged conjugates.

<sup>2</sup> ALBRECHT 95 use tau pair events of the type  $\tau^- \tau^+ \rightarrow (\ell^- \bar{\nu}_\ell \nu_\tau)(h^+ h^- h^+ (\pi^0) \bar{\nu}_\tau)$  and their charged conjugates.

<sup>3</sup> ALBRECHT 93G use tau pair events of the type  $\tau^- \tau^+ \rightarrow (\mu^- \bar{\nu}_\mu \nu_\tau)(e^+ \nu_e \bar{\nu}_\tau)$  and their charged conjugates.

## $\rho(\mu)$ PARAMETER

( $V-A$ ) theory predicts  $\rho = 0.75$ .

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.763±0.020 OUR FIT</b>				
<b>0.770±0.022 OUR AVERAGE</b>				
0.776±0.045±0.019	46k	HEISTER	01E ALEP	1991–1995 LEP runs
0.999±0.098±0.045	22k	ABREU	00L DLPH	1992–1995 runs
0.777±0.044±0.016	27k	ACKERSTAFF	99D OPAL	1990–1995 LEP runs
0.69 ±0.06 ±0.06		<sup>1</sup> ALBRECHT	98 ARG	$E_{\text{cm}}^{ee} = 9.5\text{--}10.6$ GeV
0.54 ±0.28 ±0.14		ABE	97O SLD	1993–1995 SLC runs
0.750±0.017±0.045	22k	ALEXANDER	97F CLEO	$E_{\text{cm}}^{ee} = 10.6$ GeV
0.76 ±0.07 ±0.08	3230	ALBRECHT	93G ARG	$E_{\text{cm}}^{ee} = 9.4\text{--}10.6$ GeV
0.734±0.055±0.027	3041	ALBRECHT	90E ARG	$E_{\text{cm}}^{ee} = 9.4\text{--}10.6$ GeV
0.89 ±0.14 ±0.08	1909	FORD	87B MAC	$E_{\text{cm}}^{ee} = 29$ GeV
0.81 ±0.13	727	BEHRENDIS	85 CLEO	$e^+e^-$ near $\Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$0.747 \pm 0.048 \pm 0.044$	13k	AMMAR	97B	CLEO	Repl. by ALEXANDER 97F
$0.693 \pm 0.057 \pm 0.028$		BUSKULIC	95D	ALEP	Repl. by HEISTER 01E

<sup>1</sup> ALBRECHT 98 use tau pair events of the type  $\tau^- \tau^+ \rightarrow (\ell^- \bar{\nu}_\ell \nu_\tau)(\pi^+ \pi^0 \bar{\nu}_\tau)$ , and their charged conjugates.

### $\xi(e \text{ or } \mu)$ PARAMETER

( $V-A$ ) theory predicts  $\xi = 1$ .

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
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**$0.985 \pm 0.030$  OUR FIT**

**$0.981 \pm 0.031$  OUR AVERAGE**

$0.986 \pm 0.068 \pm 0.031$	81k	HEISTER	01E	ALEP	1991–1995 LEP runs
$0.929 \pm 0.070 \pm 0.030$	36k	ABREU	00L	DLPH	1992–1995 runs
$0.98 \pm 0.22 \pm 0.10$	46k	ACKERSTAFF	99D	OPAL	1990–1995 LEP runs
$0.70 \pm 0.16$	54k	ACCIARRI	98R	L3	1991–1995 LEP runs
$1.03 \pm 0.11$		<sup>1</sup> ALBRECHT	98	ARG	$E_{\text{cm}}^{ee} = 9.5\text{--}10.6$ GeV
$1.05 \pm 0.35 \pm 0.04$		<sup>2</sup> ABE	97O	SLD	1993–1995 SLC runs
$1.007 \pm 0.040 \pm 0.015$	55k	ALEXANDER	97F	CLEO	$E_{\text{cm}}^{ee} = 10.6$ GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

$0.94 \pm 0.21 \pm 0.07$	18k	ACCIARRI	96H	L3	Repl. by ACCIARRI 98R
$0.97 \pm 0.14$		<sup>3</sup> ALBRECHT	95C	ARG	Repl. by ALBRECHT 98
$1.18 \pm 0.15 \pm 0.16$		BUSKULIC	95D	ALEP	Repl. by HEISTER 01E
$0.90 \pm 0.15 \pm 0.10$	3230	<sup>4</sup> ALBRECHT	93G	ARG	$E_{\text{cm}}^{ee} = 9.4\text{--}10.6$ GeV

<sup>1</sup> Combined fit to ARGUS tau decay parameter measurements in ALBRECHT 98, ALBRECHT 95C, ALBRECHT 93G, and ALBRECHT 94E. ALBRECHT 98 use tau pair events of the type  $\tau^- \tau^+ \rightarrow (\ell^- \bar{\nu}_\ell \nu_\tau)(\pi^+ \pi^0 \bar{\nu}_\tau)$ , and their charged conjugates.

<sup>2</sup> ABE 97O assume  $\eta = 0$  in their fit. Letting  $\eta$  vary in the fit gives a  $\xi$  value of  $1.02 \pm 0.36 \pm 0.05$ .

<sup>3</sup> Combined fit to ARGUS tau decay parameter measurements in ALBRECHT 95C, ALBRECHT 93G, and ALBRECHT 94E. ALBRECHT 95C uses events of the type  $\tau^- \tau^+ \rightarrow (\ell^- \bar{\nu}_\ell \nu_\tau)(h^+ h^- h^+ \bar{\nu}_\tau)$  and their charged conjugates.

<sup>4</sup> ALBRECHT 93G measurement determines  $|\xi|$  for the case  $\xi(e) = \xi(\mu)$ , but the authors point out that other LEP experiments determine the sign to be positive.

### $\xi(e)$ PARAMETER

( $V-A$ ) theory predicts  $\xi = 1$ .

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
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**$0.994 \pm 0.040$  OUR FIT**

**$1.00 \pm 0.04$  OUR AVERAGE**

$1.011 \pm 0.094 \pm 0.038$	44k	HEISTER	01E	ALEP	1991–1995 LEP runs
$1.01 \pm 0.12 \pm 0.05$	17k	ABREU	00L	DLPH	1992–1995 runs
$1.13 \pm 0.39 \pm 0.14$	25k	ACKERSTAFF	99D	OPAL	1990–1995 LEP runs
$1.11 \pm 0.20 \pm 0.08$		<sup>1</sup> ALBRECHT	98	ARG	$E_{\text{cm}}^{ee} = 9.5\text{--}10.6$ GeV
$1.16 \pm 0.52 \pm 0.06$		ABE	97O	SLD	1993–1995 SLC runs
$0.979 \pm 0.048 \pm 0.016$	34k	ALEXANDER	97F	CLEO	$E_{\text{cm}}^{ee} = 10.6$ GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

$1.03 \pm 0.23 \pm 0.09$		BUSKULIC	95D	ALEP	Repl. by HEISTER 01E
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<sup>1</sup> ALBRECHT 98 use tau pair events of the type  $\tau^- \tau^+ \rightarrow (\ell^- \bar{\nu}_\ell \nu_\tau)(\pi^+ \pi^0 \bar{\nu}_\tau)$ , and their charged conjugates.

**$\xi(\mu)$  PARAMETER** $(V-A)$  theory predicts  $\xi = 1$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
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**1.030 ± 0.059 OUR FIT****1.06 ± 0.06 OUR AVERAGE**

1.030 ± 0.120 ± 0.050	46k	HEISTER	01E	ALEP	1991–1995 LEP runs
1.16 ± 0.19 ± 0.06	22k	ABREU	00L	DLPH	1992–1995 runs
0.79 ± 0.41 ± 0.09	27k	ACKERSTAFF	99D	OPAL	1990–1995 LEP runs
1.26 ± 0.27 ± 0.14		<sup>1</sup> ALBRECHT	98	ARG	$E_{\text{cm}}^{ee} = 9.5\text{--}10.6$ GeV
0.75 ± 0.50 ± 0.14		ABE	97O	SLD	1993–1995 SLC runs
1.054 ± 0.069 ± 0.047	22k	ALEXANDER	97F	CLEO	$E_{\text{cm}}^{ee} = 10.6$ GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

1.23 ± 0.22 ± 0.10		BUSKULIC	95D	ALEP	Repl. by HEISTER 01E
		<sup>1</sup> ALBRECHT 98			use tau pair events of the type $\tau^- \tau^+ \rightarrow (\ell^- \bar{\nu}_\ell \nu_\tau)(\pi^+ \pi^0 \bar{\nu}_\tau)$ , and their charged conjugates.

 **$\eta(e \text{ or } \mu)$  PARAMETER** $(V-A)$  theory predicts  $\eta = 0$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
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**0.013 ± 0.020 OUR FIT****0.015 ± 0.021 OUR AVERAGE**

0.012 ± 0.026 ± 0.004	81k	HEISTER	01E	ALEP	1991–1995 LEP runs
−0.005 ± 0.036 ± 0.037		ABREU	00L	DLPH	1992–1995 runs
0.027 ± 0.055 ± 0.005	46k	ACKERSTAFF	99D	OPAL	1990–1995 LEP runs
0.27 ± 0.14	54k	ACCIARRI	98R	L3	1991–1995 LEP runs
−0.13 ± 0.47 ± 0.15		ABE	97O	SLD	1993–1995 SLC runs
−0.015 ± 0.061 ± 0.062	31k	AMMAR	97B	CLEO	$E_{\text{cm}}^{ee} = 10.6$ GeV
0.03 ± 0.18 ± 0.12	8.2k	ALBRECHT	95	ARG	$E_{\text{cm}}^{ee} = 9.5\text{--}10.6$ GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.25 ± 0.17 ± 0.11	18k	ACCIARRI	96H	L3	Repl. by ACCIARRI 98R
−0.04 ± 0.15 ± 0.11		BUSKULIC	95D	ALEP	Repl. by HEISTER 01E

 **$\eta(\mu)$  PARAMETER** $(V-A)$  theory predicts  $\eta = 0$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
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**0.094 ± 0.073 OUR FIT****0.17 ± 0.15 OUR AVERAGE** Error includes scale factor of 1.2.

0.160 ± 0.150 ± 0.060	46k	HEISTER	01E	ALEP	1991–1995 LEP runs
0.72 ± 0.32 ± 0.15		ABREU	00L	DLPH	1992–1995 runs
−0.59 ± 0.82 ± 0.45		<sup>1</sup> ABE	97O	SLD	1993–1995 SLC runs
0.010 ± 0.149 ± 0.171	13k	<sup>2</sup> AMMAR	97B	CLEO	$E_{\text{cm}}^{ee} = 10.6$ GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.010 ± 0.065 ± 0.001	27k	<sup>3</sup> ACKERSTAFF	99D	OPAL	1990–1995 LEP runs
−0.24 ± 0.23 ± 0.18		BUSKULIC	95D	ALEP	Repl. by HEISTER 01E

<sup>1</sup> Highly correlated (corr. = 0.92) with ABE 97O  $\rho(\mu)$  measurement.<sup>2</sup> Highly correlated (corr. = 0.949) with AMMAR 97B  $\rho(\mu)$  value.<sup>3</sup> ACKERSTAFF 99D result is dominated by a constraint on  $\eta$  from the OPAL measurements of the  $\tau$  lifetime and  $B(\tau^- \rightarrow \mu^- \bar{\nu}_\mu \nu_\tau)$  assuming lepton universality for the total coupling strength.

**( $\delta\xi$ )(e or  $\mu$ ) PARAMETER** $(V-A)$  theory predicts  $(\delta\xi) = 0.75$ .

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.746<math>\pm</math>0.021 OUR FIT</b>				
<b>0.744<math>\pm</math>0.022 OUR AVERAGE</b>				
0.776 $\pm$ 0.045 $\pm$ 0.024	81k	HEISTER	01E ALEP	1991–1995 LEP runs
0.779 $\pm$ 0.070 $\pm$ 0.028	36k	ABREU	00L DLPH	1992–1995 runs
0.65 $\pm$ 0.14 $\pm$ 0.07	46k	ACKERSTAFF	99D OPAL	1990–1995 LEP runs
0.70 $\pm$ 0.11	54k	ACCIARRI	98R L3	1991–1995 LEP runs
0.63 $\pm$ 0.09		<sup>1</sup> ALBRECHT	98 ARG	$E_{\text{cm}}^{ee} = 9.5\text{--}10.6$ GeV
0.88 $\pm$ 0.27 $\pm$ 0.04		<sup>2</sup> ABE	97O SLD	1993–1995 SLC runs
0.745 $\pm$ 0.026 $\pm$ 0.009	55k	ALEXANDER	97F CLEO	$E_{\text{cm}}^{ee} = 10.6$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.81 $\pm$ 0.14 $\pm$ 0.06	18k	ACCIARRI	96H L3	Repl. by ACCIARRI 98R
0.65 $\pm$ 0.12		<sup>3</sup> ALBRECHT	95C ARG	Repl. by ALBRECHT 98
0.88 $\pm$ 0.11 $\pm$ 0.07		BUSKULIC	95D ALEP	Repl. by HEISTER 01E

<sup>1</sup> Combined fit to ARGUS tau decay parameter measurements in ALBRECHT 98, ALBRECHT 95C, ALBRECHT 93G, and ALBRECHT 94E. ALBRECHT 98 use tau pair events of the type  $\tau^- \tau^+ \rightarrow (\ell^- \bar{\nu}_\ell \nu_\tau)(\pi^+ \pi^0 \bar{\nu}_\tau)$ , and their charged conjugates.

<sup>2</sup> ABE 97O assume  $\eta = 0$  in their fit. Letting  $\eta$  vary in the fit gives a  $(\delta\xi)$  value of  $0.87 \pm 0.27 \pm 0.04$ .

<sup>3</sup> Combined fit to ARGUS tau decay parameter measurements in ALBRECHT 95C, ALBRECHT 93G, and ALBRECHT 94E. ALBRECHT 95C uses events of the type  $\tau^- \tau^+ \rightarrow (\ell^- \bar{\nu}_\ell \nu_\tau)(h^+ h^- h^+ \bar{\nu}_\tau)$  and their charged conjugates.

**( $\delta\xi$ )(e) PARAMETER** $(V-A)$  theory predicts  $(\delta\xi) = 0.75$ .

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.734<math>\pm</math>0.028 OUR FIT</b>				
<b>0.731<math>\pm</math>0.029 OUR AVERAGE</b>				
0.778 $\pm$ 0.066 $\pm$ 0.024	44k	HEISTER	01E ALEP	1991–1995 LEP runs
0.85 $\pm$ 0.12 $\pm$ 0.04	17k	ABREU	00L DLPH	1992–1995 runs
0.72 $\pm$ 0.31 $\pm$ 0.14	25k	ACKERSTAFF	99D OPAL	1990–1995 LEP runs
0.56 $\pm$ 0.14 $\pm$ 0.06		<sup>1</sup> ALBRECHT	98 ARG	$E_{\text{cm}}^{ee} = 9.5\text{--}10.6$ GeV
0.85 $\pm$ 0.43 $\pm$ 0.08		ABE	97O SLD	1993–1995 SLC runs
0.720 $\pm$ 0.032 $\pm$ 0.010	34k	ALEXANDER	97F CLEO	$E_{\text{cm}}^{ee} = 10.6$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1.11 $\pm$ 0.17 $\pm$ 0.07		BUSKULIC	95D ALEP	Repl. by HEISTER 01E

<sup>1</sup> ALBRECHT 98 use tau pair events of the type  $\tau^- \tau^+ \rightarrow (\ell^- \bar{\nu}_\ell \nu_\tau)(\pi^+ \pi^0 \bar{\nu}_\tau)$ , and their charged conjugates.

**( $\delta\xi$ )( $\mu$ ) PARAMETER** $(V-A)$  theory predicts  $(\delta\xi) = 0.75$ .

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.778<math>\pm</math>0.037 OUR FIT</b>				
<b>0.79 <math>\pm</math>0.04 OUR AVERAGE</b>				
0.786 $\pm$ 0.066 $\pm$ 0.028	46k	HEISTER	01E ALEP	1991–1995 LEP runs
0.86 $\pm$ 0.13 $\pm$ 0.04	22k	ABREU	00L DLPH	1992–1995 runs
0.63 $\pm$ 0.23 $\pm$ 0.05	27k	ACKERSTAFF	99D OPAL	1990–1995 LEP runs
0.73 $\pm$ 0.18 $\pm$ 0.10		<sup>1</sup> ALBRECHT	98 ARG	$E_{\text{cm}}^{ee} = 9.5\text{--}10.6$ GeV
0.82 $\pm$ 0.32 $\pm$ 0.07		ABE	97O SLD	1993–1995 SLC runs
0.786 $\pm$ 0.041 $\pm$ 0.032	22k	ALEXANDER	97F CLEO	$E_{\text{cm}}^{ee} = 10.6$ GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.71  $\pm$ 0.14  $\pm$ 0.06 BUSKULIC 95D ALEP Repl. by HEISTER 01E  
<sup>1</sup>ALBRECHT 98 use tau pair events of the type  $\tau^- \tau^+ \rightarrow (\ell^- \bar{\nu}_\ell \nu_\tau)(\pi^+ \pi^0 \bar{\nu}_\tau)$ , and their charged conjugates.

### $\xi(\pi)$ PARAMETER

(V-A) theory predicts  $\xi(\pi) = 1$ .

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.993<math>\pm</math>0.022 OUR FIT</b>				
<b>0.994<math>\pm</math>0.023 OUR AVERAGE</b>				
0.994 $\pm$ 0.020 $\pm$ 0.014	27k	HEISTER	01E ALEP	1991–1995 LEP runs
0.81 $\pm$ 0.17 $\pm$ 0.02		ABE	97O SLD	1993–1995 SLC runs
1.03 $\pm$ 0.06 $\pm$ 0.04	2.0k	COAN	97 CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6$ GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.987 $\pm$ 0.057 $\pm$ 0.027 BUSKULIC 95D ALEP Repl. by HEISTER 01E  
0.95  $\pm$ 0.11  $\pm$ 0.05 <sup>1</sup>BUSKULIC 94D ALEP 1990+1991 LEP run  
<sup>1</sup>Superseded by BUSKULIC 95D.

### $\xi(\rho)$ PARAMETER

(V-A) theory predicts  $\xi(\rho) = 1$ .

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.994<math>\pm</math>0.008 OUR FIT</b>				
<b>0.994<math>\pm</math>0.009 OUR AVERAGE</b>				
0.987 $\pm$ 0.012 $\pm$ 0.011	59k	HEISTER	01E ALEP	1991–1995 LEP runs
0.99 $\pm$ 0.12 $\pm$ 0.04		ABE	97O SLD	1993–1995 SLC runs
0.995 $\pm$ 0.010 $\pm$ 0.003	66k	ALEXANDER	97F CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
1.022 $\pm$ 0.028 $\pm$ 0.030	1.7k	<sup>1</sup> ALBRECHT	94E ARG	$E_{\text{cm}}^{\text{ee}} = 9.4$ –10.6 GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

1.045 $\pm$ 0.058 $\pm$ 0.032 BUSKULIC 95D ALEP Repl. by HEISTER 01E  
1.03  $\pm$ 0.11  $\pm$ 0.05 <sup>2</sup>BUSKULIC 94D ALEP 1990+1991 LEP run  
<sup>1</sup>ALBRECHT 94E measure the square of this quantity and use the sign determined by ALBRECHT 90I to obtain the quoted result.  
<sup>2</sup>Superseded by BUSKULIC 95D.

### $\xi(a_1)$ PARAMETER

(V-A) theory predicts  $\xi(a_1) = 1$ .

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>1.001<math>\pm</math>0.027 OUR FIT</b>				
<b>1.002<math>\pm</math>0.028 OUR AVERAGE</b>				
1.000 $\pm$ 0.016 $\pm$ 0.024	35k	<sup>1</sup> HEISTER	01E ALEP	1991–1995 LEP runs
1.02 $\pm$ 0.13 $\pm$ 0.03	17.2k	ASNER	00 CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
1.29 $\pm$ 0.26 $\pm$ 0.11	7.4k	<sup>2</sup> ACKERSTAFF	97R OPAL	1992–1994 LEP runs
0.85 $\begin{smallmatrix} +0.15 \\ -0.17 \end{smallmatrix}$ $\pm$ 0.05		ALBRECHT	95C ARG	$E_{\text{cm}}^{\text{ee}} = 9.5$ –10.6 GeV
1.25 $\pm$ 0.23 $\begin{smallmatrix} +0.15 \\ -0.08 \end{smallmatrix}$	7.5k	ALBRECHT	93C ARG	$E_{\text{cm}}^{\text{ee}} = 9.4$ –10.6 GeV



• • • We do not use the following data for averages, fits, limits, etc. • • •

1.08	$\begin{matrix} +0.46 & +0.14 \\ -0.41 & -0.25 \end{matrix}$	2.6k	<sup>3</sup> AKERS	95P	OPAL	Repl. by ACKER-STAFF 97R
0.937 ± 0.116 ± 0.064			BUSKULIC	95D	ALEP	Repl. by HEISTER 01E

<sup>1</sup> HEISTER 01E quote  $1.000 \pm 0.016 \pm 0.013 \pm 0.020$  where the errors are statistical, systematic, and an uncertainty due to the final state model. We combine the systematic error and model uncertainty.

<sup>2</sup> ACKERSTAFF 97R obtain this result with a model independent fit to the hadronic structure functions. Fitting with the model of Kuhn and Santamaria (ZPHY **C48**, 445 (1990)) gives  $0.87 \pm 0.16 \pm 0.04$ , and with the model of of Isgur *et al.* (PR **D39**,1357 (1989)) they obtain  $1.20 \pm 0.21 \pm 0.14$ .

<sup>3</sup> AKERS 95P obtain this result with a model independent fit to the hadronic structure functions. Fitting with the model of Kuhn and Santamaria (ZPHY **C48**, 445 (1990)) gives  $0.87 \pm 0.27 \begin{matrix} +0.05 \\ -0.06 \end{matrix}$ , and with the model of of Isgur *et al.* (PR **D39**,1357 (1989)) they obtain  $1.10 \pm 0.31 \begin{matrix} +0.13 \\ -0.14 \end{matrix}$ .

### ξ(all hadronic modes) PARAMETER

(V-A) theory predicts  $\xi = 1$ .

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.995 ± 0.007 OUR FIT</b>				
<b>0.997 ± 0.007 OUR AVERAGE</b>				
0.992 ± 0.007 ± 0.008	102k	<sup>1</sup> HEISTER	01E ALEP	1991–1995 LEP runs
0.997 ± 0.027 ± 0.011	39k	<sup>2</sup> ABREU	00L DLPH	1992–1995 runs
1.02 ± 0.13 ± 0.03	17.2k	<sup>3</sup> ASNER	00 CLEO	$E_{cm}^{ee} = 10.6$ GeV
1.032 ± 0.031	37k	<sup>4</sup> ACCIARRI	98R L3	1991–1995 LEP runs
0.93 ± 0.10 ± 0.04		ABE	97O SLD	1993–1995 SLC runs
1.29 ± 0.26 ± 0.11	7.4k	<sup>5</sup> ACKERSTAFF	97R OPAL	1992–1994 LEP runs
0.995 ± 0.010 ± 0.003	66k	<sup>6</sup> ALEXANDER	97F CLEO	$E_{cm}^{ee} = 10.6$ GeV
1.03 ± 0.06 ± 0.04	2.0k	<sup>7</sup> COAN	97 CLEO	$E_{cm}^{ee} = 10.6$ GeV
1.017 ± 0.039		<sup>8</sup> ALBRECHT	95C ARG	$E_{cm}^{ee} = 9.5–10.6$ GeV
1.25 ± 0.23 $\begin{matrix} +0.15 \\ -0.08 \end{matrix}$	7.5k	<sup>9</sup> ALBRECHT	93C ARG	$E_{cm}^{ee} = 9.4–10.6$ GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.970 ± 0.053 ± 0.011	14k	<sup>10</sup> ACCIARRI	96H L3	Repl. by ACCIARRI 98R
1.08 $\begin{matrix} +0.46 & +0.14 \\ -0.41 & -0.25 \end{matrix}$	2.6k	<sup>11</sup> AKERS	95P OPAL	Repl. by ACKER-STAFF 97R
1.006 ± 0.032 ± 0.019		<sup>12</sup> BUSKULIC	95D ALEP	Repl. by HEISTER 01E
1.022 ± 0.028 ± 0.030	1.7k	<sup>13</sup> ALBRECHT	94E ARG	$E_{cm}^{ee} = 9.4–10.6$ GeV
0.99 ± 0.07 ± 0.04		<sup>14</sup> BUSKULIC	94D ALEP	1990+1991 LEP run
1.14 ± 0.34 $\begin{matrix} +0.34 \\ -0.17 \end{matrix}$	3.9k	<sup>9</sup> ALBRECHT	90I ARG	Repl. by AL-BRECHT 93C

<sup>1</sup> HEISTER 01E quote  $0.992 \pm 0.007 \pm 0.006 \pm 0.005$  where the errors are statistical, systematic, and an uncertainty due to the final state model. We combine the systematic error and model uncertainty. They use  $\tau \rightarrow \pi \nu_\tau$ ,  $\tau \rightarrow K \nu_\tau$ ,  $\tau \rightarrow \rho \nu_\tau$ , and  $\tau \rightarrow a_1 \nu_\tau$  decays.

<sup>2</sup> ABREU 00L use  $\tau^- \rightarrow h^- \geq 0\pi^0 \nu_\tau$  decays.

<sup>3</sup> ASNER 00 use  $\tau^- \rightarrow \pi^- 2\pi^0 \nu_\tau$  decays.

<sup>4</sup> ACCIARRI 98R use  $\tau \rightarrow \pi \nu_\tau$ ,  $\tau \rightarrow K \nu_\tau$ , and  $\tau \rightarrow \rho \nu_\tau$  decays.

<sup>5</sup> ACKERSTAFF 97R use  $\tau \rightarrow a_1 \nu_\tau$  decays.

<sup>6</sup> ALEXANDER 97F use  $\tau \rightarrow \rho \nu_\tau$  decays.

<sup>7</sup> COAN 97 use  $h^+ h^-$  energy correlations.

<sup>8</sup> Combined fit to ARGUS tau decay parameter measurements in ALBRECHT 95C, ALBRECHT 93G, and ALBRECHT 94E.

<sup>9</sup> Uses  $\tau \rightarrow a_1 \nu_\tau$  decays. Replaced by ALBRECHT 95C.

<sup>10</sup> ACCIARRI 96H use  $\tau \rightarrow \pi \nu_\tau$ ,  $\tau \rightarrow K \nu_\tau$ , and  $\tau \rightarrow \rho \nu_\tau$  decays.

<sup>11</sup> AKERS 95P use  $\tau \rightarrow a_1 \nu_\tau$  decays.

<sup>12</sup> BUSKULIC 95D use  $\tau \rightarrow \pi \nu_\tau$ ,  $\tau \rightarrow \rho \nu_\tau$ , and  $\tau \rightarrow a_1 \nu_\tau$  decays.

<sup>13</sup> ALBRECHT 94E measure the square of this quantity and use the sign determined by ALBRECHT 90I to obtain the quoted result. Uses  $\tau \rightarrow a_1 \nu_\tau$  decays. Replaced by ALBRECHT 95C.

<sup>14</sup> BUSKULIC 94D use  $\tau \rightarrow \pi \nu_\tau$  and  $\tau \rightarrow \rho \nu_\tau$  decays. Superseded by BUSKULIC 95D.

## $\xi'(\mu)$ PARAMETER

(V-A) theory predicts  $\xi' = 1$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.22±0.94±0.42</b>	165	BODROV	23	BELL $E_{\text{cm}}^e = 10.6$ GeV

## $\bar{\eta}(\mu)$ PARAMETER

(V-A) theory predicts  $\bar{\eta}(\mu) = 0$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>-1.3±1.5±0.8</b>	71k	<sup>1</sup> SHIMIZU	18A	BELL $\tau^- \rightarrow \nu_\tau \mu^- \bar{\nu}_\mu \gamma$

<sup>1</sup> SHIMIZU 18A measurement procedure fits a distribution affected by  $\bar{\eta}(\mu)$ ,  $\xi\kappa(\mu)$  and  $\eta''(\mu)$ , floating  $\bar{\eta}(\mu)$  and  $\xi\kappa(\mu)$  and fixing  $\eta''(\mu) = 0$ . The contribution of  $\eta''(\mu)$  is suppressed by  $m_\mu/m_\tau$ .

## $(\xi\kappa)(e \text{ or } \mu)$ PARAMETER

(V-A) theory predicts  $(\xi\kappa)(e \text{ or } \mu) = 0$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.5±0.4±0.2</b>	149k	<sup>1,2</sup> SHIMIZU	18A	BELL $\tau^- \rightarrow \nu_\tau e^- \bar{\nu}_e \gamma$ and $\tau^- \rightarrow \nu_\tau \mu^- \bar{\nu}_\mu \gamma$

<sup>1</sup> SHIMIZU 18A measurement procedure fits a distribution of radiative tau decays into both electrons and muons affected by  $\bar{\eta}(e \text{ or } \mu)$ ,  $\xi\kappa(e \text{ or } \mu)$  and  $\eta''(e \text{ or } \mu)$ , floating  $\bar{\eta}(e \text{ or } \mu)$  and  $\xi\kappa(e \text{ or } \mu)$  and fixing  $\eta''(e \text{ or } \mu) = 0$ . The contribution of  $\eta''(e \text{ or } \mu)$  is suppressed by  $m_e/m_\tau$  for tau decaying to electrons and by  $m_\mu/m_\tau$  for tau decaying to muons.

<sup>2</sup> Error correlated with SHIMIZU 18A  $(\xi\kappa)(e)$  and  $(\xi\kappa)(\mu)$  values.

## $(\xi\kappa)(e)$ PARAMETER

(V-A) theory predicts  $(\xi\kappa)(e) = 0$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>-0.4±0.8±0.9</b>	78k	<sup>1,2</sup> SHIMIZU	18A	BELL $\tau^- \rightarrow \nu_\tau e^- \bar{\nu}_e \gamma$

<sup>1</sup> SHIMIZU 18A measurement procedure fits a distribution affected by  $\bar{\eta}(e)$ ,  $(\xi\kappa)(e)$  and  $\eta''(e)$ , floating  $(\xi\kappa)(e)$  and fixing  $\bar{\eta}(e) = 0$  and  $\eta''(e) = 0$ . The contribution of  $\eta''(e)$  is suppressed by  $m_e/m_\tau$ .

<sup>2</sup> Error correlated with SHIMIZU 18A  $(\xi\kappa)(e \text{ or } \mu)$  value.

## $(\xi\kappa)(\mu)$ PARAMETER

(V-A) theory predicts  $(\xi\kappa)(\mu) = 0$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.8±0.5±0.3</b>	71k	<sup>1,2</sup> SHIMIZU	18A	BELL $\tau^- \rightarrow \nu_\tau \mu^- \bar{\nu}_\mu \gamma$

<sup>1</sup> SHIMIZU 18A measurement procedure fits a distribution affected by  $\bar{\eta}(\mu)$ ,  $\xi\kappa(\mu)$  and  $\eta''(\mu)$ , floating  $\bar{\eta}(\mu)$  and  $\xi\kappa(\mu)$  and fixing  $\eta''(\mu) = 0$ . The contribution of  $\eta''(\mu)$  is suppressed by  $m_\mu/m_\tau$ .

<sup>2</sup> Error correlated with SHIMIZU 18A ( $\xi\kappa$ )( $e$  or  $\mu$ ) value. **$\tau$  REFERENCES**

AAD	23BM	PRL 131 151802	G. Aad <i>et al.</i>	(ATLAS Collab.)
ADACHI	23A	PRL 130 181803	I. Adachi <i>et al.</i>	(BELLE II Collab.)
ADACHI	23C	PR D108 032006	I. Adachi <i>et al.</i>	(BELLE II Collab.)
ANASHIN	23A	PPN 54 185	V.V. Anashin <i>et al.</i>	(VEPP-4M KEDR Collab.)
BODROV	23	PRL 131 021801	D. Bodrov <i>et al.</i>	(BELLE Collab.)
	Also	PR D108 012003	D. Bodrov <i>et al.</i>	(BELLE Collab.)
TSUZUKI	23	JHEP 2306 118	N. Tsuzuki <i>et al.</i>	(BELLE Collab.)
TUMASYAN	23AS	PRL 131 151803	A. Tumasyan <i>et al.</i>	(CMS Collab.)
INAMI	22	JHEP 2204 110	K. Inami <i>et al.</i>	(BELLE Collab.)
BRYMAN	21	PR D104 075032	D.A. Bryman, S. Ito, R. Shrock	(BRCO, TRIU+)
SIRUNYAN	21D	JHEP 2101 163	A.M. Sirunyan <i>et al.</i>	(CMS Collab.)
UNO	21	JHEP 2110 019	K. Uno <i>et al.</i>	(BELLE Collab.)
SAHOO	20	PR D102 111101	D. Sahoo <i>et al.</i>	(BELLE Collab.)
JIN	19	PR D100 071101	Y. Jin <i>et al.</i>	(BELLE Collab.)
LEES	18B	PR D98 032010	J.P. Lees <i>et al.</i>	(BABAR Collab.)
SHIMIZU	18A	PTEP 2018 023C01	N. Shimizu <i>et al.</i>	(BELLE Collab.)
AAD	16BA	EPJ C76 232	G.Aad <i>et al.</i>	(ATLAS Collab.)
AAIJ	15AI	JHEP 1502 121	R. Aaij <i>et al.</i>	(LHCb Collab.)
LEES	15G	PR D91 051103	J.P. Lees <i>et al.</i>	(BABAR Collab.)
ABLIKIM	14D	PR D90 012001	M. Ablikim <i>et al.</i>	(BESIII Collab.)
BELOUS	14	PRL 112 031801	K. Belous <i>et al.</i>	(BELLE Collab.)
LEVICHEV	14	SPU 57 66	E.B. Levichev <i>et al.</i>	
RYU	14	PR D89 072009	S. Ryu <i>et al.</i>	(BELLE Collab.)
AAIJ	13AH	PL B724 36	R. Aaij <i>et al.</i>	(LHCb Collab.)
MIYAZAKI	13	PL B719 346	Y. Miyazaki <i>et al.</i>	(BELLE Collab.)
LEES	12M	PR D85 031102	J.P. Lees <i>et al.</i>	(BABAR Collab.)
	Also	PR D85 099904 (errat.)	J.P. Lees <i>et al.</i>	(BABAR Collab.)
LEES	12X	PR D86 092010	J.P. Lees <i>et al.</i>	(BABAR Collab.)
LEES	12Y	PR D86 092013	J.P. Lees <i>et al.</i>	(BABAR Collab.)
PDG	12	PR D86 010001	J. Beringer <i>et al.</i>	(PDG Collab.)
DEL-AMO-SA...	11E	PR D83 032002	P. del Amo Sanchez <i>et al.</i>	(BABAR Collab.)
MIYAZAKI	11	PL B699 251	Y. Miyazaki <i>et al.</i>	(BELLE Collab.)
AUBERT	10B	PRL 104 021802	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	10F	PRL 105 051602	B. Aubert <i>et al.</i>	(BABAR Collab.)
HAYASAKA	10	PL B687 139	K. Hayasaka <i>et al.</i>	(BELLE Collab.)
LEE	10	PR D81 113007	M.J. Lee <i>et al.</i>	(BELLE Collab.)
LEES	10A	PR D81 111101	J.P. Lees <i>et al.</i>	(BABAR Collab.)
MIYAZAKI	10	PL B682 355	Y. Miyazaki <i>et al.</i>	(BELLE Collab.)
MIYAZAKI	10A	PL B692 4	Y. Miyazaki <i>et al.</i>	(BELLE Collab.)
AUBERT	09AK	PR D80 092005	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	09D	PR D79 012004	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	09W	PRL 103 021801	B. Aubert <i>et al.</i>	(BABAR Collab.)
GROZIN	09A	PAN 72 1203	A.G. Grozin, I.B. Khriplovich, A.S. Rudenko	(NOVO)
INAMI	09	PL B672 209	K. Inami <i>et al.</i>	(BELLE Collab.)
MIYAZAKI	09	PL B672 317	Y. Miyazaki <i>et al.</i>	(BELLE Collab.)
AUBERT	08	PRL 100 011801	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	08AE	PR D77 112002	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	08K	PRL 100 071802	B. Aubert <i>et al.</i>	(BABAR Collab.)
FUJIKAWA	08	PR D78 072006	M. Fujikawa <i>et al.</i>	(BELLE Collab.)
HAYASAKA	08	PL B666 16	K. Hayasaka <i>et al.</i>	(BELLE Collab.)
MIYAZAKI	08	PL B660 154	Y. Miyazaki <i>et al.</i>	(BELLE Collab.)
NISHIO	08	PL B664 35	Y. Nishio <i>et al.</i>	(BELLE Collab.)
ANASHIN	07	JETPL 85 347	V.V. Anashin <i>et al.</i>	(KEDR Collab.)
		Translated from ZETFP 85 429.		
AUBERT	07AP	PR D76 051104	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	07BK	PRL 99 251803	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	07I	PRL 98 061803	B. Aubert <i>et al.</i>	(BABAR Collab.)
BELOUS	07	PRL 99 011801	K. Belous <i>et al.</i>	(BELLE Collab.)
EIDELMAN	07	MPL A22 159	S. Eidelman, M. Passera	(NOVO, PADO)
EPIFANOV	07	PL B654 65	D. Epifanov <i>et al.</i>	(BELLE Collab.)
MIYAZAKI	07	PL B648 341	Y. Miyazaki <i>et al.</i>	(BELLE Collab.)
ABDALLAH	06A	EPJ C46 1	J. Abdallah <i>et al.</i>	(DELPHI Collab.)
AUBERT	06C	PRL 96 041801	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT,B	06	PR D73 112003	B. Aubert <i>et al.</i>	(BABAR Collab.)
INAMI	06	PL B643 5	K. Inami <i>et al.</i>	(BELLE Collab.)
MIYAZAKI	06	PL B632 51	Y. Miyazaki <i>et al.</i>	(BELLE Collab.)

MIYAZAKI	06A	PL B639 159	Y. Miyazaki <i>et al.</i>	(BELLE Collab.)
PDG	06	JP G33 1	W.-M. Yao <i>et al.</i>	(PDG Collab.)
YUSA	06	PL B640 138	Y. Yusa <i>et al.</i>	(BELLE Collab.)
ARMS	05	PRL 94 241802	K. Arms <i>et al.</i>	(CLEO Collab.)
AUBERT,B	05A	PRL 95 041802	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT,B	05F	PR D72 012003	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT,B	05W	PR D72 072001	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT,BE	05D	PRL 95 191801	B. Aubert <i>et al.</i>	(BABAR Collab.)
ENARI	05	PL B622 218	Y. Enari <i>et al.</i>	(BELLE Collab.)
HAYASAKA	05	PL B613 20	K. Hayasaka <i>et al.</i>	(BELLE Collab.)
SCHAEEL	05C	PRPL 421 191	S. Schael <i>et al.</i>	(ALEPH Collab.)
ABBIENDI	04J	EPJ C35 437	G. Abbiendi <i>et al.</i>	(OPAL Collab.)
ABDALLAH	04K	EPJ C35 159	J. Abdallah <i>et al.</i>	(DELPHI Collab.)
ABDALLAH	04T	EPJ C36 283	J. Abdallah <i>et al.</i>	(DELPHI Collab.)
ABE	04B	PRL 92 171802	K. Abe <i>et al.</i>	(BELLE Collab.)
ACHARD	04G	PL B585 53	P. Achard <i>et al.</i>	(L3 Collab.)
AUBERT	04J	PRL 92 121801	B. Aubert <i>et al.</i>	(BABAR Collab.)
ENARI	04	PRL 93 081803	Y. Enari <i>et al.</i>	(BELLE Collab.)
PDG	04	PL B592 1	S. Eidelman <i>et al.</i>	(PDG Collab.)
YUSA	04	PL B589 103	Y. Yusa <i>et al.</i>	(BELLE Collab.)
ABBIENDI	03	PL B551 35	G. Abbiendi <i>et al.</i>	(OPAL Collab.)
BRIERE	03	PRL 90 181802	R. A. Briere <i>et al.</i>	(CLEO Collab.)
HEISTER	03F	EPJ C30 291	A. Heister <i>et al.</i>	(ALEPH Collab.)
INAMI	03	PL B551 16	K. Inami <i>et al.</i>	(BELLE Collab.)
CHEN	02C	PR D66 071101	S. Chen <i>et al.</i>	(CLEO Collab.)
REGAN	02	PRL 88 071805	B.C. Regan <i>et al.</i>	
ABBIENDI	01J	EPJ C19 653	G. Abbiendi <i>et al.</i>	(OPAL Collab.)
ABREU	01M	EPJ C20 617	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ACCIARRI	01F	PL B507 47	M. Acciarri <i>et al.</i>	(L3 Collab.)
ACHARD	01D	PL B519 189	P. Achard <i>et al.</i>	(L3 Collab.)
ANASTASSOV	01	PRL 86 4467	A. Anastassov <i>et al.</i>	(CLEO Collab.)
HEISTER	01E	EPJ C22 217	A. Heister <i>et al.</i>	(ALEPH Collab.)
ABBIENDI	00A	PL B492 23	G. Abbiendi <i>et al.</i>	(OPAL Collab.)
ABBIENDI	00C	EPJ C13 213	G. Abbiendi <i>et al.</i>	(OPAL Collab.)
ABBIENDI	00D	EPJ C13 197	G. Abbiendi <i>et al.</i>	(OPAL Collab.)
ABREU	00L	EPJ C16 229	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ACCIARRI	00B	PL B479 67	M. Acciarri <i>et al.</i>	(L3 Collab.)
AHMED	00	PR D61 071101	S. Ahmed <i>et al.</i>	(CLEO Collab.)
ALBRECHT	00	PL B485 37	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ASNER	00	PR D61 012002	D.M. Asner <i>et al.</i>	(CLEO Collab.)
ASNER	00B	PR D62 072006	D.M. Asner <i>et al.</i>	(CLEO Collab.)
BERGFELD	00	PRL 84 830	T. Bergfeld <i>et al.</i>	(CLEO Collab.)
BROWDER	00	PR D61 052004	T.E. Browder <i>et al.</i>	(CLEO Collab.)
EDWARDS	00A	PR D61 072003	K.W. Edwards <i>et al.</i>	(CLEO Collab.)
GONZALEZ-S...	00	NP B582 3	G.A. Gonzalez-Sprinberg <i>et al.</i>	
ABBIENDI	99H	PL B447 134	G. Abbiendi <i>et al.</i>	(OPAL Collab.)
ABREU	99X	EPJ C10 201	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ACKERSTAFF	99D	EPJ C8 3	K. Akerstaff <i>et al.</i>	(OPAL Collab.)
ACKERSTAFF	99E	EPJ C8 183	K. Akerstaff <i>et al.</i>	(OPAL Collab.)
BARATE	99K	EPJ C10 1	R. Barate <i>et al.</i>	(ALEPH Collab.)
BARATE	99R	EPJ C11 599	R. Barate <i>et al.</i>	(ALEPH Collab.)
BISHAI	99	PRL 82 281	M. Bishai <i>et al.</i>	(CLEO Collab.)
GODANG	99	PR D59 091303	R. Godang <i>et al.</i>	(CLEO Collab.)
RICHICHI	99	PR D60 112002	S.J. Richichi <i>et al.</i>	(CLEO Collab.)
ACCIARRI	98C	PL B426 207	M. Acciarri <i>et al.</i>	(L3 Collab.)
ACCIARRI	98E	PL B434 169	M. Acciarri <i>et al.</i>	(L3 Collab.)
ACCIARRI	98R	PL B438 405	M. Acciarri <i>et al.</i>	(L3 Collab.)
ACKERSTAFF	98M	EPJ C4 193	K. Akerstaff <i>et al.</i>	(OPAL Collab.)
ACKERSTAFF	98N	PL B431 188	K. Akerstaff <i>et al.</i>	(OPAL Collab.)
ALBRECHT	98	PL B431 179	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
BARATE	98	EPJ C1 65	R. Barate <i>et al.</i>	(ALEPH Collab.)
BARATE	98E	EPJ C4 29	R. Barate <i>et al.</i>	(ALEPH Collab.)
BLISS	98	PR D57 5903	D.W. Bliss <i>et al.</i>	(CLEO Collab.)
ABE	97O	PRL 78 4691	K. Abe <i>et al.</i>	(SLD Collab.)
ACKERSTAFF	97J	PL B404 213	K. Akerstaff <i>et al.</i>	(OPAL Collab.)
ACKERSTAFF	97L	ZPHY C74 403	K. Akerstaff <i>et al.</i>	(OPAL Collab.)
ACKERSTAFF	97R	ZPHY C75 593	K. Akerstaff <i>et al.</i>	(OPAL Collab.)
ALEXANDER	97F	PR D56 5320	J.P. Alexander <i>et al.</i>	(CLEO Collab.)
AMMAR	97B	PRL 78 4686	R. Ammar <i>et al.</i>	(CLEO Collab.)
ANASTASSOV	97	PR D55 2559	A. Anastassov <i>et al.</i>	(CLEO Collab.)
Also		PR D58 119903 (errat.)	A. Anastassov <i>et al.</i>	(CLEO Collab.)

ANDERSON	97	PRL 79 3814	S. Anderson <i>et al.</i>	(CLEO Collab.)
AVERY	97	PR D55 1119	P. Avery <i>et al.</i>	(CLEO Collab.)
BARATE	97I	ZPHY C74 387	R. Barate <i>et al.</i>	(ALEPH Collab.)
BARATE	97R	PL B414 362	R. Barate <i>et al.</i>	(ALEPH Collab.)
BERGFELD	97	PRL 79 2406	T. Bergfeld <i>et al.</i>	(CLEO Collab.)
BONVICINI	97	PRL 79 1221	G. Bonvicini <i>et al.</i>	(CLEO Collab.)
BUSKULIC	97C	ZPHY C74 263	D. Buskulic <i>et al.</i>	(ALEPH Collab.)
COAN	97	PR D55 7291	T.E. Coan <i>et al.</i>	(CLEO Collab.)
EDWARDS	97	PR D55 3919	K.W. Edwards <i>et al.</i>	(CLEO Collab.)
EDWARDS	97B	PR D56 5297	K.W. Edwards <i>et al.</i>	(CLEO Collab.)
ESCRIBANO	97	PL B395 369	R. Escrivano, E. Masso	(BARC, PARIT)
ABREU	96B	PL B365 448	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ACCIARRI	96H	PL B377 313	M. Acciarri <i>et al.</i>	(L3 Collab.)
ACCIARRI	96K	PL B389 187	M. Acciarri <i>et al.</i>	(L3 Collab.)
ALAM	96	PRL 76 2637	M.S. Alam <i>et al.</i>	(CLEO Collab.)
ALBRECHT	96E	PRPL 276 223	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALEXANDER	96D	PL B369 163	G. Alexander <i>et al.</i>	(OPAL Collab.)
ALEXANDER	96E	PL B374 341	G. Alexander <i>et al.</i>	(OPAL Collab.)
ALEXANDER	96S	PL B388 437	G. Alexander <i>et al.</i>	(OPAL Collab.)
BAI	96	PR D53 20	J.Z. Bai <i>et al.</i>	(BES Collab.)
BALEST	96	PL B388 402	R. Balest <i>et al.</i>	(CLEO Collab.)
BARTELT	96	PRL 76 4119	J.E. Bartelt <i>et al.</i>	(CLEO Collab.)
BUSKULIC	96	ZPHY C70 579	D. Buskulic <i>et al.</i>	(ALEPH Collab.)
BUSKULIC	96C	ZPHY C70 561	D. Buskulic <i>et al.</i>	(ALEPH Collab.)
COAN	96	PR D53 6037	T.E. Coan <i>et al.</i>	(CLEO Collab.)
ABE	95Y	PR D52 4828	K. Abe <i>et al.</i>	(SLD Collab.)
ABREU	95T	PL B357 715	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ABREU	95U	PL B359 411	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ACCIARRI	95	PL B345 93	M. Acciarri <i>et al.</i>	(L3 Collab.)
ACCIARRI	95F	PL B352 487	M. Acciarri <i>et al.</i>	(L3 Collab.)
AKERS	95F	ZPHY C66 31	R. Akers <i>et al.</i>	(OPAL Collab.)
AKERS	95I	ZPHY C66 543	R. Akers <i>et al.</i>	(OPAL Collab.)
AKERS	95P	ZPHY C67 45	R. Akers <i>et al.</i>	(OPAL Collab.)
AKERS	95Y	ZPHY C68 555	R. Akers <i>et al.</i>	(OPAL Collab.)
ALBRECHT	95	PL B341 441	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALBRECHT	95C	PL B349 576	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALBRECHT	95G	ZPHY C68 25	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALBRECHT	95H	ZPHY C68 215	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
BALEST	95C	PRL 75 3809	R. Balest <i>et al.</i>	(CLEO Collab.)
BERNABEU	95	NP B436 474	J. Bernabeu <i>et al.</i>	
BUSKULIC	95C	PL B346 371	D. Buskulic <i>et al.</i>	(ALEPH Collab.)
BUSKULIC	95D	PL B346 379	D. Buskulic <i>et al.</i>	(ALEPH Collab.)
Also		PL B363 265 (errat.)	D. Buskulic <i>et al.</i>	(ALEPH Collab.)
ABREU	94K	PL B334 435	P. Abreu <i>et al.</i>	(DELPHI Collab.)
AKERS	94E	PL B328 207	R. Akers <i>et al.</i>	(OPAL Collab.)
AKERS	94G	PL B339 278	R. Akers <i>et al.</i>	(OPAL Collab.)
ALBRECHT	94E	PL B337 383	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ARTUSO	94	PRL 72 3762	M. Artuso <i>et al.</i>	(CLEO Collab.)
BARTELT	94	PRL 73 1890	J.E. Bartelt <i>et al.</i>	(CLEO Collab.)
BATTLE	94	PRL 73 1079	M. Battle <i>et al.</i>	(CLEO Collab.)
BAUER	94	PR D50 13	D.A. Bauer <i>et al.</i>	(TPC/2gamma Collab.)
BUSKULIC	94D	PL B321 168	D. Buskulic <i>et al.</i>	(ALEPH Collab.)
BUSKULIC	94E	PL B332 209	D. Buskulic <i>et al.</i>	(ALEPH Collab.)
BUSKULIC	94F	PL B332 219	D. Buskulic <i>et al.</i>	(ALEPH Collab.)
GIBAUT	94B	PRL 73 934	D. Gibaut <i>et al.</i>	(CLEO Collab.)
ADRIANI	93M	PRPL 236 1	O. Adriani <i>et al.</i>	(L3 Collab.)
ALBRECHT	93C	ZPHY C58 61	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALBRECHT	93G	PL B316 608	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
BALEST	93	PR D47 3671	R. Balest <i>et al.</i>	(CLEO Collab.)
BEAN	93	PRL 70 138	A. Bean <i>et al.</i>	(CLEO Collab.)
BORTOLETTO	93	PRL 71 1791	D. Bortoletto <i>et al.</i>	(CLEO Collab.)
ESCRIBANO	93	PL B301 419	R. Escrivano, E. Masso	(BARC)
PROCARIO	93	PRL 70 1207	M. Procaro <i>et al.</i>	(CLEO Collab.)
ABREU	92N	ZPHY C55 555	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ACTON	92F	PL B281 405	D.P. Acton <i>et al.</i>	(OPAL Collab.)
ACTON	92H	PL B288 373	P.D. Acton <i>et al.</i>	(OPAL Collab.)
AKERIB	92	PRL 69 3610	D.S. Akerib <i>et al.</i>	(CLEO Collab.)
Also		PRL 71 3395 (errat.)	D.S. Akerib <i>et al.</i>	(CLEO Collab.)
ALBRECHT	92D	ZPHY C53 367	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALBRECHT	92K	ZPHY C55 179	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALBRECHT	92M	PL B292 221	H. Albrecht <i>et al.</i>	(ARGUS Collab.)

ALBRECHT	92Q	ZPHY C56 339	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
AMMAR	92	PR D45 3976	R. Ammar <i>et al.</i>	(CLEO Collab.)
ARTUSO	92	PRL 69 3278	M. Artuso <i>et al.</i>	(CLEO Collab.)
BAI	92	PRL 69 3021	J.Z. Bai <i>et al.</i>	(BES Collab.)
BATTLE	92	PL B291 488	M. Battle <i>et al.</i>	(CLEO Collab.)
BUSKULIC	92J	PL B297 459	D. Buskulic <i>et al.</i>	(ALEPH Collab.)
DECAMP	92C	ZPHY C54 211	D. Decamp <i>et al.</i>	(ALEPH Collab.)
ADEVA	91F	PL B265 451	B. Adeva <i>et al.</i>	(L3 Collab.)
ALBRECHT	91D	PL B260 259	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALEXANDER	91D	PL B266 201	G. Alexander <i>et al.</i>	(OPAL Collab.)
ANTREASYAN	91	PL B259 216	D. Antreasyan <i>et al.</i>	(Crystal Ball Collab.)
GRIFOLS	91	PL B255 611	J.A. Grifols, A. Mendez	(BARC)
ABACHI	90	PR D41 1414	S. Abachi <i>et al.</i>	(HRS Collab.)
ALBRECHT	90E	PL B246 278	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALBRECHT	90I	PL B250 164	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
BEHREND	90	ZPHY C46 537	H.J. Behrend <i>et al.</i>	(CELLO Collab.)
BOWCOCK	90	PR D41 805	T.J.V. Bowcock <i>et al.</i>	(CLEO Collab.)
DELAGUILA	90	PL B252 116	F. del Aguila, M. Sher	(BARC, WILL)
GOLDBERG	90	PL B251 223	M. Goldberg <i>et al.</i>	(CLEO Collab.)
WU	90	PR D41 2339	D.Y. Wu <i>et al.</i>	(Mark II Collab.)
ABACHI	89B	PR D40 902	S. Abachi <i>et al.</i>	(HRS Collab.)
BEHREND	89B	PL B222 163	H.J. Behrend <i>et al.</i>	(CELLO Collab.)
JANSSEN	89	PL B228 273	H. Janssen <i>et al.</i>	(Crystal Ball Collab.)
KLEINWORT	89	ZPHY C42 7	C. Kleinwort <i>et al.</i>	(JADE Collab.)
ADEVA	88	PR D38 2665	B. Adeva <i>et al.</i>	(Mark-J Collab.)
ALBRECHT	88B	PL B202 149	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALBRECHT	88L	ZPHY C41 1	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALBRECHT	88M	ZPHY C41 405	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
AMIDEI	88	PR D37 1750	D. Amidei <i>et al.</i>	(Mark II Collab.)
BEHREND	88	PL B200 226	H.J. Behrend <i>et al.</i>	(CELLO Collab.)
BRAUNSCH...	88C	ZPHY C39 331	W. Braunschweig <i>et al.</i>	(TASSO Collab.)
KEH	88	PL B212 123	S. Keh <i>et al.</i>	(Crystal Ball Collab.)
TSCHIRHART	88	PL B205 407	R. Tschirhart <i>et al.</i>	(HRS Collab.)
ABACHI	87B	PL B197 291	S. Abachi <i>et al.</i>	(HRS Collab.)
ABACHI	87C	PRL 59 2519	S. Abachi <i>et al.</i>	(HRS Collab.)
ADLER	87B	PRL 59 1527	J. Adler <i>et al.</i>	(Mark III Collab.)
AIHARA	87B	PR D35 1553	H. Aihara <i>et al.</i>	(TPC Collab.)
AIHARA	87C	PRL 59 751	H. Aihara <i>et al.</i>	(TPC Collab.)
ALBRECHT	87L	PL B185 223	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALBRECHT	87P	PL B199 580	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
BAND	87	PL B198 297	H.R. Band <i>et al.</i>	(MAC Collab.)
BAND	87B	PRL 59 415	H.R. Band <i>et al.</i>	(MAC Collab.)
BARINGER	87	PRL 59 1993	P. Baringer <i>et al.</i>	(CLEO Collab.)
BEBEK	87C	PR D36 690	C. Bebek <i>et al.</i>	(CLEO Collab.)
BURCHAT	87	PR D35 27	P.R. Burchat <i>et al.</i>	(Mark II Collab.)
BYLSMA	87	PR D35 2269	B.G. Bylsma <i>et al.</i>	(HRS Collab.)
COFFMAN	87	PR D36 2185	D.M. Coffman <i>et al.</i>	(Mark III Collab.)
DERRICK	87	PL B189 260	M. Derrick <i>et al.</i>	(HRS Collab.)
FORD	87	PR D35 408	W.T. Ford <i>et al.</i>	(MAC Collab.)
FORD	87B	PR D36 1971	W.T. Ford <i>et al.</i>	(MAC Collab.)
GAN	87	PRL 59 411	K.K. Gan <i>et al.</i>	(Mark II Collab.)
GAN	87B	PL B197 561	K.K. Gan <i>et al.</i>	(Mark II Collab.)
AIHARA	86E	PRL 57 1836	H. Aihara <i>et al.</i>	(TPC Collab.)
BARTEL	86D	PL B182 216	W. Bartel <i>et al.</i>	(JADE Collab.)
PDG	86	PL 170B 1	M. Aguilar-Benitez <i>et al.</i>	(CERN, CIT+)
RUCKSTUHL	86	PRL 56 2132	W. Ruckstuhl <i>et al.</i>	(DELCO Collab.)
SCHMIDKE	86	PRL 57 527	W.B. Schmidke <i>et al.</i>	(Mark II Collab.)
YELTON	86	PRL 56 812	J.M. Yelton <i>et al.</i>	(Mark II Collab.)
ALTHOFF	85	ZPHY C26 521	M. Althoff <i>et al.</i>	(TASSO Collab.)
ASH	85B	PRL 55 2118	W.W. Ash <i>et al.</i>	(MAC Collab.)
BALTRUSAIT...	85	PRL 55 1842	R.M. Baltrusaitis <i>et al.</i>	(Mark III Collab.)
BARTEL	85F	PL 161B 188	W. Bartel <i>et al.</i>	(JADE Collab.)
BEHREND	85	PR D32 2468	S. Behrends <i>et al.</i>	(CLEO Collab.)
BELTRAMI	85	PRL 54 1775	I. Beltrami <i>et al.</i>	(HRS Collab.)
BERGER	85	ZPHY C28 1	C. Berger <i>et al.</i>	(PLUTO Collab.)
BURCHAT	85	PRL 54 2489	P.R. Burchat <i>et al.</i>	(Mark II Collab.)
FERNANDEZ	85	PRL 54 1624	E. Fernandez <i>et al.</i>	(MAC Collab.)
MILLS	85	PRL 54 624	G.B. Mills <i>et al.</i>	(DELCO Collab.)
AIHARA	84C	PR D30 2436	H. Aihara <i>et al.</i>	(TPC Collab.)
BEHREND	84	ZPHY C23 103	H.J. Behrend <i>et al.</i>	(CELLO Collab.)
MILLS	84	PRL 52 1944	G.B. Mills <i>et al.</i>	(DELCO Collab.)

BEHREND	83C	PL 127B 270	H.J. Behrend <i>et al.</i>	(CELLO Collab.)
SILVERMAN	83	PR D27 1196	D.J. Silverman, G.L. Shaw	(UCI)
BEHREND	82	PL 114B 282	H.J. Behrend <i>et al.</i>	(CELLO Collab.)
BLOCKER	82B	PRL 48 1586	C.A. Blocker <i>et al.</i>	(Mark II Collab.)
BLOCKER	82D	PL 109B 119	C.A. Blocker <i>et al.</i>	(Mark II Collab.) J
FELDMAN	82	PRL 48 66	G.J. Feldman <i>et al.</i>	(Mark II Collab.)
HAYES	82	PR D25 2869	K.G. Hayes <i>et al.</i>	(Mark II Collab.)
BERGER	81B	PL 99B 489	C. Berger <i>et al.</i>	(PLUTO Collab.)
DORFAN	81	PRL 46 215	J.M. Dorfan <i>et al.</i>	(Mark II Collab.)
BRANDELIK	80	PL 92B 199	R. Brandelik <i>et al.</i>	(TASSO Collab.)
ZHOLENTZ	80	PL 96B 214	A.A. Zholents <i>et al.</i>	(NOVO)
Also		SJNP 34 814	A.A. Zholents <i>et al.</i>	(NOVO)
		Translated from YAF 34 1471.		
BACINO	79B	PRL 42 749	W.J. Bacino <i>et al.</i>	(DELCO Collab.)
KIRKBY	79	SLAC-PUB-2419	J. Kirkby	(SLAC) J
		Batavia Lepton Photon Conference.		
BACINO	78B	PRL 41 13	W.J. Bacino <i>et al.</i>	(DELCO Collab.) J
Also		Tokyo Conf. 249	J. Kirz	(STON)
Also		PL 96B 214	A.A. Zholents <i>et al.</i>	(NOVO)
BRANDELIK	78	PL 73B 109	R. Brandelik <i>et al.</i>	(DASP Collab.) J
FELDMAN	78	Tokyo Conf. 777	G.J. Feldman	(SLAC) J
JAROS	78	PRL 40 1120	J. Jaros <i>et al.</i>	(LGW Collab.)
PERL	75	PRL 35 1489	M.L. Perl <i>et al.</i>	(LBL, SLAC)

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PICH	90	MPL A5 1995	A. Pich	(VALE)
BARISH	88	PRPL 157 1	B.C. Barish, R. Stroynowski	(CIT)
GAN	88	IJMP A3 531	K.K. Gan, M.L. Perl	(SLAC)
HAYES	88	PR D38 3351	K.G. Hayes, M.L. Perl	(SLAC)
PERL	80	ARNPS 30 299	M.L. Perl	(SLAC)